

# Stress and Defect Effects on Electron Transport Properties at  $SnO<sub>2</sub>/$ Perovskite Interfaces: A First-Principles Insight

[Wenhua Pu,](https://pubs.acs.org/action/doSearch?field1=Contrib&text1="Wenhua+Pu"&field2=AllField&text2=&publication=&accessType=allContent&Earliest=&ref=pdf) [Wei Xiao,](https://pubs.acs.org/action/doSearch?field1=Contrib&text1="Wei+Xiao"&field2=AllField&text2=&publication=&accessType=allContent&Earliest=&ref=pdf)[\\*](#page-7-0) [Jianwei Wang,](https://pubs.acs.org/action/doSearch?field1=Contrib&text1="Jianwei+Wang"&field2=AllField&text2=&publication=&accessType=allContent&Earliest=&ref=pdf) [Xiao-Wu Li,](https://pubs.acs.org/action/doSearch?field1=Contrib&text1="Xiao-Wu+Li"&field2=AllField&text2=&publication=&accessType=allContent&Earliest=&ref=pdf) [and Ligen Wang](https://pubs.acs.org/action/doSearch?field1=Contrib&text1="Ligen+Wang"&field2=AllField&text2=&publication=&accessType=allContent&Earliest=&ref=pdf)

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ABSTRACT: The structural and electronic properties of interfaces play an important role in the stability and functionality of solar cell devices. Experiments indicate that the  $SnO<sub>2</sub>/perovskite$  interfaces always show superior electron transport efficiency and high structural stability even though there exists a larger lattice mismatch. Aiming at solving the puzzles, we have performed density-functional theory calculations to investigate the electronic characteristics of the  $SnO<sub>2</sub>/$ perovskite interfaces with various stresses and defects. The results prove that the  $PbI_2/SnO_2$  interfaces have better structural stability and superior characteristics for the electron transport. The tensile stress could move the conduction band minimum (CBM) of  $CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>$  upward, while the compressive stress could move the CBM of  $SnO<sub>2</sub>$  downward. By taking into account the stress effect, the



CBM offset is 0.07 eV at the PbI<sub>2</sub>/SnO<sub>2</sub> interface and 0.28 eV at the MAI/SnO<sub>2</sub> interface. Moreover, our calculations classify V<sub>I</sub> and  $I_i$  at the PbI<sub>2</sub>/SnO<sub>2</sub> interface and Sn–I, I<sub>i</sub> and Sn<sub>i</sub> at the MAI/SnO<sub>2</sub> interface as harmful defects. The I<sub>i</sub> defects are the most easily formed harmful defects and should be avoided at both interfaces. The calculated results are in agreement with the available experimental observations. The present work provides a theoretical basis for improving the stability and photovoltaic performance of the perovskite solar cells.

# 1. INTRODUCTION

Perovskites have been emerging as one of the important candidate materials for new photovoltaic cells because they possess superior photovoltaic performance of high light absorption coefficient, large carrier mobility, and long electron−hole diffusion length and also are easy to prepare in various synthetic methods with low manufacturing  $cost<sup>1−</sup>$ Perovskite organic metal halides were employed as a light harvester for the perovskite solar cells (PSCs) by Kojima et al. in 2009.<sup>[5](#page-7-0)</sup> With the efforts of many researchers, the efficiency of perovskite-based solar cells has increased from 3.8% to  $25.5\%$  $25.5\%$  $25.5\%$ <sup>5−[8](#page-8-0)</sup> PSCs are generally composed of the transparent conductive electrodes, electron transport layers (ETLs), perovskite light absorption layers, hole transport layers (HTLs), and metal electrodes.<sup>[9](#page-8-0)</sup> These layers are stacked together, and many interfaces are formed between the layers, all of which will affect the charge transport of the device. The bonding and defect properties of the interfaces play an important role in the interfacial stability and functionality. Therefore, it is very important to study the characteristics of the interfaces, which will help researchers overcome the shortcomings at the interfaces and continuously improve the efficiency of PSCs.

Electron transport layers are an important part of PSCs, which are responsible for transferring electrons to the electrodes. Electrons excited by light in the perovskite layer must transfer through the perovskite/ETL interfaces and then be collected by the ETL. Meanwhile, electrons and holes might recombine at the interfaces, which often affects the power conversion efficiency (PCE) of the device. TiO<sub>2</sub> is the most popular ETL used in PSCs because of its suitable conduction band minimum (CBM) and the history of being used as an ETL in dye sensitized solar cells.<sup>[10,11](#page-8-0)</sup> The PCE of TiO<sub>2</sub>-based PSCs has exceeded 25%.<sup>[12](#page-8-0)</sup> However, the development of TiO<sub>2</sub> is limited by a serious disadvantage, that is, its electron transport speed is very low, only  $0.1-0.4$  cm<sup>2</sup> V s<sup>-1</sup> and electrons are easy to accumulate at the  $TiO<sub>2</sub>/perovskite$ interfaces.<sup>[13](#page-8-0)</sup> Many other materials such as  $ZnO, ^{14}Zn_2SnO_4, ^{15}$  $ZnO, ^{14}Zn_2SnO_4, ^{15}$  $ZnO, ^{14}Zn_2SnO_4, ^{15}$ BaSnO<sub>3</sub>,<sup>[16](#page-8-0)</sup> SrTiO<sub>3</sub>,<sup>[17](#page-8-0)</sup> CdS,<sup>[18](#page-8-0)</sup> CdSe,<sup>[19](#page-8-0)</sup> WO<sub>3</sub>,<sup>[20](#page-8-0)</sup> In<sub>2</sub>O<sub>3</sub>,<sup>[21](#page-8-0)</sup>  $Nb_2O_5$ <sup>[22](#page-8-0)</sup> and  $CeO<sub>x</sub><sup>23</sup>$  $CeO<sub>x</sub><sup>23</sup>$  $CeO<sub>x</sub><sup>23</sup>$  have been investigated as ETL with different strengths and weaknesses. In comparison,  $SnO<sub>2</sub>$  as the promising ETL has attracted wide attention in recent years, and the PCE of  $SnO_2$ -based PSCs has also exceeded 25%.<sup>[24](#page-8-0)</sup>

Received: March 16, 2022 Accepted: April 6, 2022 Published: April 26, 2022





<span id="page-1-0"></span>

Figure 1. Optimized stable geometrical structures of  $SnO<sub>2</sub>/perovskite$  interfaces: (a) PbI<sub>2</sub>/SnO<sub>2</sub>; (b) MAI/SnO<sub>2</sub>.

 $SnO<sub>2</sub>$  has superior physical properties for being used as an ETL. Its electron transfer efficiency is much higher, reaching  $240 \text{ cm}^2 \text{ V s}^{-1}$ ,<sup>[13](#page-8-0)</sup> although it has a cell and band structure similar to those of TiO<sub>2</sub>. The CBM of  $SnO<sub>2</sub>$  is lower than  $TiO<sub>2</sub>$ , which is favorable for electronic transport.<sup>[25](#page-8-0)</sup> Also, the larger band gap of  $SnO<sub>2</sub>$  makes it less susceptible to ultraviolet rays.<sup>[26](#page-8-0)</sup> Moreover, SnO<sub>2</sub> can be prepared at low temperatures, which is more convenient and can be used in flexible solar cells.<sup>[27](#page-8-0)</sup> In 2015, Li et al. were the first batch to use mesoporous  $SnO<sub>2</sub>$  films as ETL and got a PCE of 10.18%.<sup>[28](#page-8-0)</sup> Ke et al. prepared  $SnO<sub>2</sub>$  thin films by thermal decomposition of  $SnCl<sub>2</sub>$ .  $2H<sub>2</sub>O$  precursor on fluorine doped tin oxide (FTO), and the yield PCE reached 17.21%.<sup>[29](#page-8-0)</sup> Anaraki et al. obtained 20.7% PCE by the spin coating and chemical bath deposition method.<sup>30</sup> Jiang et al. fabricated planar PSCs with a certificated PCE of 23.32% by using the organic halide salt phenylethylammonium iodide (PEAI) to passivate the surface defects of HC(NH<sub>2</sub>)<sub>2</sub>−CH<sub>3</sub>NH<sub>3</sub> mixed perovskite films.<sup>[31](#page-8-0)</sup> Recently, Yoo et al. tuned the chemical bath deposition of  $SnO<sub>2</sub>$  and got better ETL with ideal film coverage, thickness, and composition, which could be beneficial for the interface properties, and thus a certified PCE of 25.2% was finally obtained.<sup>2</sup>

Many experimenters have reached a consensus on the excellent performance of  $SnO<sub>2</sub>$  as an ETL and they have employed a lot of approaches to improve the properties of the  $\text{SnO}_2$ /perovskite interfaces.<sup>[32](#page-8-0)–[35](#page-8-0)</sup> Yang et al. found that the recombination rate of quantum dot  $SnO<sub>2</sub>/perovskite interfaces$ is less than that of nanocrystalline  $SnO<sub>2</sub>/perovskite$  interfaces. $26$  Zhang et al. used density functional theory (DFT) and experiments to comprehensively investigate the interface structures and transport properties of the  $CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>/$  $SnO<sub>2</sub>$  interfaces and found that the interface contacted by the  $PbI<sub>2</sub>$  layer is more stable than that contacted by the  $CH<sub>3</sub>NH<sub>3</sub>I$ layer. $34$  In order to adjust the interfacial stress caused by the lattice mismatch, Du et al. introduced the amino acid selfassembled layer as the buffer layer, which improved the quality of perovskite films, enhanced the charge transfer/extraction at the interfaces, and finally made it achieve a PCE of 20.68%.<sup>[35](#page-8-0)</sup> Kim et al. studied the orbital hybridization and the effects of O vacancy and Ti/Sn interstitial at TiO<sub>2</sub>/perovskite and  $SnO<sub>2</sub>/$ perovskite interfaces and indicated that  $SnO<sub>2</sub>$  is better than  $TiO<sub>2</sub>$  as an ETL material.<sup>[36](#page-8-0)</sup> Until now the understanding of the influences of the microstructures on the performance of the  $SnO<sub>2</sub>/$  perovskite interfaces, such as the stress produced by the

mismatch and the defect effects, is far from enough. So in this paper, we will carry out a systematic study on the structural and electronic properties of the  $SnO<sub>2</sub>/perovskite$  interfaces. The remainder of the paper is organized as follows. In Section 2, the theoretical methods and the computational details are described. [Section 3](#page-2-0) presents the interfacial properties of the clean heterointerfaces, and the stress and defect effects on the electron transport at the interfaces are explored. Finally, a short summary is given in [Section 4.](#page-7-0)

# 2. COMPUTATIONAL METHODS

The first-principles calculations in the framework of DFT were performed using Vienna ab initio Simulation Package (VASP).[37,38](#page-8-0) The electron−ion interaction was described using the projector augmented wave method.<sup>[39](#page-8-0),[40](#page-8-0)</sup> The energy cutoff for the plane wave basis set was 500 eV for all calculations. In the simulation, an empirical pairwise correction proposed by Grimme in terms of the DFT+D2 scheme had also been included for more precisely depicting the dispersion interactions in the systems.  $41$  For both bulk and interface systems, their structures were optimized by using the generalized gradient approximation (GGA) in the Perdew– Burke−Ernzerhof (PBE) form.[42](#page-8-0) For the accuracy of the band alignment and electronic transport properties at the interfaces, the electronic properties were calculated using the Heyd− Scuseria−Ernzerhof screened Coulomb hybrid functional  $(HSE06).$ <sup>[43](#page-8-0)</sup> The supercell parameters and the atomic positions were allowed to relax until the forces on all atoms were converged to 0.05 eV/Å.

The calculated parameters of bulk tetragonal  $CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>$  $(MAPbI<sub>3</sub>)$  perovskites are a = 8.69 and c = 12.80 Å, which match well with the previous DFT calculations and experimental observations.  $44-46$  $44-46$  In order to build a typical  $SnO<sub>2</sub>/perovskite interface, we also optimized the lattice$ constants of  $SnO<sub>2</sub>$  to be a = 4.83 and c = 3.24 Å, which are in good agreement with other researches. $47,48$  $47,48$  $47,48$  Then, we constrained the lattice constants of the perovskite slab with these of the  $SnO<sub>2</sub>$  slab, which is always employed as the substrate in experiments. Haruyama et al. listed all possible surface of tetragonal  $CH_3NH_3PbI_3$ , and we chose the (001) surface to stack interface models because the (001) surface has the best match with the  $SnO_2$  lattice.<sup>[49](#page-9-0)</sup> Due to the structural characteristics of the (001) slab of perovskites, we produced two kinds of interfaces based on the different terminations: the MAI termination with  $MA^+$  and I<sup>−</sup> ions and the PbI<sub>2</sub>

<span id="page-2-0"></span>termination with  $Pb^{2+}$  and  $I^-$  ions, as shown in [Figure 1.](#page-1-0) The heterointerfaces were built by connecting the five-layer  $SnO<sub>2</sub>$ slabs with seven-layer perovskite slabs together and left 20 Å vacuum along the nonperiodic direction. The  $SnO<sub>2</sub>$  slab contains  $10 \text{ SnO}_2$  units, namely, 60 atoms. The perovskite slab contains 6.5 MAPbI<sub>3</sub> units with 78 atoms for  $PbI_2$  termination, and 7.5 MAPbI<sub>3</sub> units with 90 atoms for MAI termination. The bottom atoms of the  $SnO<sub>2</sub>$  substrate were fixed to keep the bulk environment. The corresponding *k*-point mesh was  $4 \times 4$  $\times$  1.

#### 3. RESULTS AND DISCUSSION

3.1. Stability and Bonding Characteristics of  $SnO<sub>2</sub>/$ Perovskite Interfaces. In order to evaluate the stability of the  $SnO<sub>2</sub>/perovskite interfaces, we calculate the interfacial$ adhesion energy  $(E_{ad})$ , which is defined as

$$
E_{\rm ad} = -(E_{\rm interface} - E_{\rm SnO_2} - E_{\rm perovskite})/S \tag{1}
$$

where  $E_{\rm interface}$ ,  $E_{\rm SnO_2}$ , and  $E_{\rm perovskite}$  denote the total energies of the heterointerface supercell, the  $SnO<sub>2</sub>$  substrate and the perovskite slab, respectively and S is the area of the interfaces. The slabs of  $SnO<sub>2</sub>$  and perovskite are constrained to match the lattice constants at the interfaces. A positive interfacial adhesion energy implies that it is energetically favorable to form the interface.

By using eq 1, we have obtained that the  $PbI_2/SnO_2$ interfaces have higher interfacial adhesion energy of 1.22 J/  $m<sup>2</sup>$  as shown in Table 1. For the MAI/SnO<sub>2</sub> interfaces, it is

Table 1. Calculated Interfacial Adhesion Energy, Lattice Mismatch and Bader Charge of Different  $SnO<sub>2</sub>/Perovskite$ Interfaces

interface	interfacial adhesion energy $(I/m^2)$	lattice mismatch (%)	Bader charge (e)
$PbI_2/$ $\overline{\text{SnO}}_2$	1.22	11	0.74
MAI/ SnO <sub>2</sub>	0.66	11	0.54

only 0.66 J/m<sup>2</sup>. These results show that the  $PbI_2/SnO_2$ interfaces are more stable than the  $MAI/SnO<sub>2</sub>$  interfaces. Since perovskites are supposed to grow on the  $SnO<sub>2</sub>$  surface in experiments, their lattice constants parallel to the interface are compressed to match  $SnO<sub>2</sub>$ . Then, we calculated the lattice mismatch of the  $SnO_2/$  perovskite heterointerfaces to be about 11% (Table 1), larger than that  $(7.35%)$  of TiO<sub>2</sub>/perovskite heterointerfaces. However, compared to the stable  $PbI_2/TiO_2$ interfaces (0.93 J/m<sup>2</sup>), the  $\text{SnO}_2\text{/perovskite}$  interface is energetically more stable even though it has a large mismatch[.50](#page-9-0) But when referred to the heterointerfaces of traditional CZTS/CdS solar cells  $(3.05 \text{ J/m}^2)$ , it can be deduced that the binding at the  $SnO_2/perovskite$  interfaces is still relatively weak.<sup>[51](#page-9-0)</sup>

In order to study why the stability of the  $PbI_2/SnO_2$ interface is better, we continue to study the bonding characteristics at  $SnO<sub>2</sub>/perovskite interfaces from the atomic$ and electronic levels. First, the analyses of Bader charge are listed in Table 1 to address the interactions between two slabs. It shows that for the  $PbI_2/SnO_2$  interfaces, the perovskite slab lost more electron (0.74 e) than the  $MAI/SnO<sub>2</sub>$  interfaces (0.54 e). With more electron transfer between the two slabs, the binding of the  $PbI_2/SnO_2$  interfaces might be stronger.

Second, the proportion of bonded atoms of the perovskite surface and bond length of the interface structures are obtained in Table 2. The bonding ratio of one kind of bond is defined as

# Table 2. Proportion of Bonded Atoms of the Perovskite Surface and Bond Length of Interface Structures



the surface chemical bond/surface atom number (O and Sn). In the PbI<sub>2</sub> termination, the bonding ratio of Pb−O bonds is 50%, and the bond length of Pb−O is 2.29 Å. For Sn−I bonds, the bonding ratio is 100% and the bond length is 2.82−2.94 Å. In the MAI termination, H−O and Sn−I bonds dominate the stability of the interfaces, and the bonding ratios (bond length) are 37.5% (1.41−1.63 Å) and 50% (2.87 Å), respectively. On the basis of the above analyses of the atomic structures at the interfaces, we can conclude that the  $PbI_2/SnO_2$  interfaces not only have more atoms bonded but also contain some Pb−O bonds with high strength when compared to the  $MAI/SnO<sub>2</sub>$ interfaces, which can be considered another strong evidence for its high interfacial adhesion energy.

Furthermore, the charge density differences at the interfaces are given to clarify the interactions of different atoms. [Figure 2](#page-3-0) depicts the charge density differences (left panels) and planaraveraged charge density differences (right panels) of the two interfaces. We can clearly observe the strong Pb−O and Sn−I coupling at the  $PbI_2/SnO_2$  interfaces, while only some Sn–I coupling at the  $MAI/SnO<sub>2</sub>$  interfaces. Electrons are favorable to accumulate around O and I atoms (yellow region) and deplete around Sn and Pb atoms (green region). The weak charge transfer at the  $MAI/SnO<sub>2</sub>$  interfaces can also be quantitatively described by the planar-averaged charge density differences along the Z direction. The charge distribution at the interfaces could form an internal electric field to affect the charge transport properties of the solar cells. Thus, more stable  $PbI_2/SnO_2$  interfaces are supposed to cause a stronger internal electric field and make the electron−hole pairs separate more easily at the interfaces.<sup>[52](#page-9-0)</sup> We can speculate that the  $PbI_2/SnO_2$ interfaces could be easier to form during the device manufacturing process as evidenced by experiments,<sup>[53](#page-9-0)</sup> and it is beneficial to the performance improvement of the electron transport.

3.2. Characteristics of the Electron Transport at SnO<sub>2</sub>/Perovskite Interfaces. The electrical properties at the interfaces are closely related to the electron transport at the interfaces. In order to understand the electrical properties of  $SnO<sub>2</sub>/perovskite interfaces, we first resorted to the partial$ density of states (PDOSs) of the interfaces, as shown in [Figure](#page-3-0) [3](#page-3-0). It can be seen from [Figure 3](#page-3-0) that the CBM of  $SnO<sub>2</sub>$  at two different interfaces are all significantly lower than that of perovskites. Because with a proper band gap, perovskites always act as the light absorption layer, electrons should be excited from the top of the valence band maximum (VBM) (Ip and Pb-s orbitals) of perovskites to the conduction band minimum (CBM) of perovskites (Pb-p orbitals) and then transferred from the conduction bands of the perovskites to the conduction bands of  $SnO<sub>2</sub>$  (Sn-s orbitals). It also clearly shows

<span id="page-3-0"></span>

Figure 2. Main views of the 3D charge density differences (left panels) and planar-averaged charge density differences (right panels) along the Z direction of the (a)  $PbI_2/SnO_2$  and (b)  $MAI/SnO_2$  interfaces, respectively. The location of two different defects I-1 and I-2 at the latter interfaces are also marked. The yellow region denotes the electron accumulation, and the green region represents the electron depletion.

e/Bohr<sup>3</sup>



Figure 3. Partial density of states (PDOSs) for the upper and lower two-layer atoms at the (a)  $PbI_2/SnO_2$  and (b)  $MAI/SnO_2$  interfaces.



Figure 4. CBM and VBM of the (a)  $PbI_2/SnO_2$  and (b)  $MAI/SnO_2$  interfaces with their deformation rates in brackets below.

that MA molecules have little effect on CBM and VBM at the interfaces and almost no participation in electron transport. By comparing the band gaps of the two interfaces by the HSE06, it can be found that that the band gap of the  $PbI_2/SnO_2$  (1.76 eV) interface is slightly smaller than that of the  $MAI/SnO<sub>2</sub>$ interface (2.26 eV). This is because the outflow of electrons from the Pb atoms causes the Pb states to shift to the left, thereby reducing the band gap.<sup>[54](#page-9-0)</sup> The band gap of the  $PbI_2$ /  $SnO<sub>2</sub>$  interfaces is closer to the ideal one of single-junction solar cells, which indicates that the PCE of the  $PbI_2/SnO_2$ interfaces might be higher. Moreover, the rate of electron transport at the interfaces can be determined by the CBM

e/Bohr<sup>3</sup>

difference between the perovskite and  $SnO<sub>2</sub>$  layers. A larger energy difference could accelerate the electron transport. Thus, comparing the positions of Pb-p and Sn-s orbitals in [Figure 3](#page-3-0), we can safely say that the electron transport at the  $MAI/SnO<sub>2</sub>$ interfaces could have a higher efficiency. To deeply understand the characteristics of the electronic transport at the interfaces, the stress and defect effects will be further discussed in the following.

3.2.1. Stress Effect on the Electronic Transport at SnO $\sqrt$ Perovskite Interfaces. Because of the lattice mismatch between the perovskite and  $SnO<sub>2</sub>$  layers, there necessarily exists stress at the interfaces. The stress could have a nonnegligible impact on the electron transport at the interfaces and therefore should be taken into account. [Figure 4](#page-3-0) shows the CBM and VBM of the perovskite and  $SnO<sub>2</sub>$  layers under different deformation rates, where positive and negative values indicate that the material bears tensile stress and compressive stress, respectively. We considered the following situations: (1) all the deformation caused by the lattice mismatch occurred on the perovskite layer and it was subjected to tensile strain at the deformation rate of 9%; (2) all the deformation was applied on the SnO<sub>2</sub> layer and it was subjected to compressive strain at the deformation rate of −8.28%; (3) the deformation was equally applied to the perovskite and  $SnO<sub>2</sub>$  layers, and the deformation rate of the perovskite and  $SnO<sub>2</sub>$  layers is 4.5% and −4.14%, respectively. The deformation rate here refers to the degree of deformation of the matched lattice constant at the interfaces relative to the initial lattice constant. The cell volume remains unchanged under different conditions of deformation. The core level alignment method has been employed to obtain the band alignment of two different  $SnO<sub>2</sub>/perovskite interfaces$ by using the PBE-HSE functional.<sup>[55,56](#page-9-0)</sup> From [Figure 4](#page-3-0), we can see for both  $PbI_2/SnO_2$  and  $MAI/SnO_2$  interfaces, the CBM of perovskite is uplifted gradually under the tensile stress and the corresponding band gap increases; the CBM of  $SnO<sub>2</sub>$  is decreased gradually under the compressive stress with a reduced band gap. When the deformation at the interfaces mainly occurs on the perovskite layer, the CBM difference between the perovskite and  $SnO<sub>2</sub>$  is relatively small, i.e., 0.07 eV at  $PbI_2/SnO_2$  interfaces and 0.28 eV at  $MAI/SnO_2$ interfaces, which could be beneficial to improve the PCE of the PSCs;<sup>[57,58](#page-9-0)</sup> when the SnO<sub>2</sub> layer carries all the deformation at the interfaces, the CBM difference between the two layers could be enlarged, i.e.,  $0.55$  eV at  $PbI_2/SnO_2$  interfaces and 0.76 eV at  $MAI/SnO<sub>2</sub>$  interfaces, which is conducive to improve the electron transport efficiency at the interfaces.[54,59](#page-9-0) From [Figure 4,](#page-3-0) it can be also found that the CBM offset between the two layers at the  $MAI/SnO<sub>2</sub>$  interfaces is greater than that at the  $PbI_2/SnO_2$  interfaces, indicating that the MAI/  $SnO<sub>2</sub>$  interfaces might be more suitable for electron transport. As in the preparation process of PSCs,  $MAPbI<sub>3</sub>$  always grows on the  $SnO<sub>2</sub>$  layer, the perovskite layer is supposed to withstand the major deformation caused by the lattice mismatch at the interfaces. By considering the stress effect, we find that for the PSCs the CBM offset at the  $PbI_2/SnO_2$ interfaces is close to 0.07 eV, and at the  $MAI/SnO<sub>2</sub>$  interfaces is approximately 0.28 eV. There are many researchers measuring the CBM offset at perovskite/ $SnO<sub>2</sub>$  interfaces in the experiments. For example, Park et al. and Chen et al. measured the optical properties of  $SnO_2/CH_3NH_3PbI_3$ heterojunction by the ultraviolet−visible spectrophotometer, and they both showed the difference was  $0.27$  eV.<sup>[60,61](#page-9-0)</sup> Yang et al. measured the Fermi level by Kelvin probe force microscopy

(KPFM) and obtained the value of 0.10  $eV<sub>1</sub><sup>62</sup>$  and Xiong et al. measured the VBM and CBM of  $SnO<sub>2</sub>$  by the UV photoelectron spectroscopy (UPS) and speculated it to be 0.30 eV. $63$  The predicted values by the high accurate HSE06 method are very consistent with the experimental results, and our theoretical calculations have proved that the stress can be used to adjust the PCE and the electron transport efficiency at the interfaces.<sup>[64](#page-9-0),[65](#page-9-0)</sup>

3.2.2. Defect Effect on the Electronic Transport at SnO $\sqrt$ Perovskite Interfaces. Defects are more likely to occur at the interfaces, and the produced defects are likely to affect the electron transport properties at the interfaces. Therefore, we have studied the formation of several defects that may occur at the two types of interfaces and analyzed their influence on the electron transport properties at the interfaces. The formula for calculating defect formation energy is shown in eq 2.

$$
E_{\rm f} = E_{\rm d} - E_{\rm SC} - \sum_{\rm i} n_{\rm i} \mu_{\rm i} \tag{2}
$$

 $E_d$  and  $E_{SC}$  are the total energies of the defect system and defect-free supercell structure, respectively.  $\mu_i$  and  $n_i$  are the chemical potentials and the number of defect species i that have been added to  $(n_i > 0)$  or removed from  $(n_i < 0)$  the supercell structure. When the number of all the species in the supercell does not change, such as antisite substitutions, then  $n_i$  $= 0.$ 

The formation energies of defects depend on the chemical potentials  $(\mu_i)$  of the constituent elements in preparation environments. In this work, we use the chemical potentials at the point B ( $\Delta \mu_{\rm Pb}$  = -1.06 eV,  $\Delta \mu_i$  = -0.6 eV,  $\Delta \mu_{\rm MA}$  = -2.41 eV) mentioned by Yin's work as a reference, which is an intermediate state for the growth of the perovskites.<sup>66</sup>  $\Delta \mu_i$  and  $\mu_i$  are linked by

$$
\Delta \mu_i = \mu_i - \mu_i^{\text{ref}} \tag{3}
$$

where  $\mu_i^{\text{ref}}$  is the chemical potential of the most stable elemental phase.

For the defects related with  $SnO<sub>2</sub>$ , the relevant chemical potentials  $\mu_{Sn}$  and  $\mu_{O}$  can be determined by the following formula:

$$
\Delta \mu_{\text{Sn}} + 2\Delta \mu_0 = \Delta H(\text{SnO}_2) = -4.77 \text{eV}
$$
\n(4)

where  $\Delta H(SnO_2)$  is the formation enthalpy of  $SnO_2$ . Here, we continue to employ an intermediate state  $(\Delta \mu_{o}^{\text{med}})$  and  $\Delta \mu_{Sn}^{\text{med}})$ as the growth condition, which is the average value of the two extreme environments, namely, O-richest (Sn-poorest) and Opoorest (Sn-richest) conditions. In the former case,  $\Delta \mu_{o}^{\rm rich} = 0$ , and the corresponding Sn-poorest chemical potential  $\Delta \mu_{\rm Sn}^{\rm pool}$ can be obtained by eq 4. In the latter case, the O-poorest chemical potential  $\Delta \mu_{o}^{\rm poor}$  can be obtained through eq 4 by setting  $\Delta \mu_{\rm Sn}^{\rm rich} = 0$ . On the basis of the above, we compute  $\Delta \mu_{\rm o}^{\rm med}$  =  $-1.19$  eV and  $\Delta \mu_{\rm Sn}^{\rm med}$  =  $-2.39$  eV, for the intermediate state. Therefore, the related defect formation energies by using the chemical potentials  $(\Delta \mu_{\rm o}^{\rm med}$  and  $\Delta \mu_{\rm Sn}^{\rm med})$  can be derived.

In previous studies, Yin et al.'s DFT calculations demonstrated that  $V_{Pb}$ ,  $V_{I}$ ,  $I_{i}$ ,  $V_{MA}$  and  $MA_{Pb}$  defects with low formation energies have transition levels less than 0.05 eV above (below) VBM (CBM) in  $\mathrm{CH_{3}NH_{3}PbI_{3}}^{66}$  $\mathrm{CH_{3}NH_{3}PbI_{3}}^{66}$  $\mathrm{CH_{3}NH_{3}PbI_{3}}^{66}$  Freeman et al. found that the point defects in  $SnO<sub>2</sub>$  are mainly Schottky defects, and the interstitial defects  $Sn_i$  and  $O_i$  are easily formed. $67$  Noh et al. found there are a lot of oxygen vacancies in  $SnO<sub>2</sub>$  films.<sup>[68](#page-9-0)</sup> Shi et al. indicated that there might be some

Table 3. Defect Formation Energies at the  $PbI_2/SnO_2$  Interfaces

	defect									
	$Sn-$	$ -$	<b>TT</b> $V_{MA}$			$MA_{Pb}$		$Pb - O$	Sn,	$V_{Pb}$
formation energy (eV)	2.20	0.37	1.58	0.10	0.26	$-0.60$	1.37	2.69	$-0.21$	3.60

element exchanges at the interfaces. $69$  On the basis of these results, we have considered all these possible point defects that might be easily formed at the  $CH_3NH_3PbI_3/SnO_2$  interfaces, including four vacancies ( $V_O$ ,  $V_{Pb}$ ,  $V_L$ ,  $V_{MA}$ ), three interstitials  $(Sn_i, O_i, I_i)$ , one cation substitution  $(MA_{Pb})$ , and two antisite substitutions (Sn−I, Pb−O). Table 3 lists the formation energies of ten possible defects at the  $PbI<sub>2</sub>/SnO<sub>2</sub>$  interfaces. Among these defects, these five types of defects, such as  $V_I$ ,  $V_O$ ,  $\rm I_{i}$ , MA $_{\rm Pb}$ , and Sn $_{\rm i}$  are more likely to form at the interfaces with a formation energy less than  $1$  eV.  $V<sub>O</sub>$  has a low formation energy of 0.10 eV, which is consistent with the previous experiments' observations that O vacancy can easily form at the interface.<sup>[68](#page-9-0),[69](#page-9-0)</sup> For  $V_I$  and  $I_i$ , their formation energies at the interface are 0.37 and 0.26 eV, respectively, relatively smaller when compared to the bulk. It can be attributed to the bonding environment at the interface that the lattices of the perovskites are expanded by the tensile stress, which weakens the bond strength of Pb−I bonds and enlarges the interstice space. Therefore,  $V_I$  and  $I_i$  become easily formed at the interface. For the MA<sub>Pb</sub>, the formation energy at the interface is largely reduced to −0.60 eV. By checking the bonding characteristics, we can observe that the MA loses an H to form  $CH<sub>3</sub>NH<sub>2</sub>$ , and the H preferably passivates the dangling bonds of O atoms at the interface, which largely reduces the formation energy of MAPb. [See [Figure S1](https://pubs.acs.org/doi/suppl/10.1021/acsomega.2c01584/suppl_file/ao2c01584_si_001.pdf).] Similarly, one interstitial Sn atom could form two strong Sn−O bonds at the interface, which makes the formation energy of  $Sn_i$  much lower than the bulk. However, it is worth noting that the free volume and the dangling O atoms might be exhausted by a small content of defects, and then the interstitial defects could become hard to form.

In order to address the impact of the defects on the electron transport efficiency at the  $PbI_2/SnO_2$  interface, the PDOSs of these five defect systems by the HSE06 method have been depicted in Figure 5. From Figure 5a, we can see the formation of  $V_I$  at the interfaces causes Pb- $p$  states slightly shift to the left, which reduces the band gap and makes the CBM of  $CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>$  significantly lower than that of  $SnO<sub>2</sub>$ . Therefore, the electron transport from  $CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>$  to  $SnO<sub>2</sub>$  becomes denpendent on thermal effects. For  $V<sub>O</sub>$  at the interface, it almost has little effect on the band gap, VBM and CBM (Figure 5b), which demonstrates that a small amount of oxygen vacancies at the interface might be acceptable. As for  $I_i$ at the interface, it causes I- $p$  states to shift slightly to the left (Figure 5c), which is close to or just below the CBM of  $SnO<sub>2</sub>$ . To some extent, it could also affect the electron transport mechanism at the interface. For  $Sn_i$  at the interfaces, we can observe a shallow defect level (Sn-s) lying at  $E = -0.7$  eV below the VBM (Figure 5d). The shallow defect level caused by  $Sn<sub>i</sub>$  could not become the recombination center of electrons and holes and might not be harmful for the electron transport at the interfaces. The formation of  $MA_{Pb}$  at the interface causes the Pb- $p$  states slightly shift to the right (Figure 5e); this does not affect the transport mechanism of electrons but may cause electrons accumulation at the interface. This change could increase the open circuit voltage of the PSCs, which finally affects the PCE of solar cells.<sup>7</sup>



Figure 5. PDOSs for the upper and lower two-layer atoms at the  $PbI_2/SnO_2$  interfaces with different defects: (a) I vacancy, (b) O vacancy,  $(c)$  I interstitial,  $(d)$  Sn interstitial, and  $(e)$  MA<sub>Pb</sub> cation substitution.

As for the  $MAI/SnO<sub>2</sub>$  interfaces, we have also chosen several possible point defects including five vacancies ( $V_{\text{O}}$ ,  $V_{\text{Pb}}$ ,  $V_{\text{I-1}}$ ,  $V_{I-2}$ ,  $V_{MA}$ ), three interstitials  $(Sn_i, O_i, I_i)$ , one cation substitution (MA<sub>Pb</sub>), and one antisite substitution (Sn− I).[66](#page-9-0)−[69](#page-9-0) [Table 4](#page-6-0) lists the corresponding formation energies of these possible defects. Due to the different bonding environments at the  $MAI/SnO<sub>2</sub>$  interfaces, two different iodine atoms labeled as I-1 and I-2 in [Figure 2](#page-3-0) are involved in this study. Among these ten defects, eight types of defects have a formation energy less than 1 eV, namely, Sn−I,  $V_{I-1}$ ,  $V_{I-2}$ ,  $V_{O}$ ,  $I_{i}$  $\text{MA}_{\text{Pb}}$ ,  $\text{O}_{\text{p}}$  and  $\text{Sn}_{\text{i}}$ . Obviously, compared to the defects at the  $PbI_2/SnO_2$  interfaces, Sn–I, V<sub>O</sub>, I<sub>i</sub>, O<sub>i</sub>, and Sn<sub>i</sub> at the MAI/  $SnO<sub>2</sub>$  interfaces are favorable to form with a relatively smaller formation energy. We can attribute these to the poor stability and bonding characteristics of the  $MAI/SnO<sub>2</sub>$  interfaces. The Sn−I antisite substitution becomes easily formed at this weakbinding interface with the formation energy of 0.89 eV. For the defect of I vacancy, the formation energy of  $V_{1-2}$  (0.33 eV) is



<span id="page-6-0"></span>Table 4. Defect Formation Energies at the  $MAI/SnO<sub>2</sub>$  Interfaces

Figure 6. PDOSs for upper and lower two-layer atoms at the MAI/SnO2 interfaces with different defects: (a) Sn−I antisite substitution, (b) O vacancy, (c) I-1 vacancy, (d) I-2 vacancy, (e) I interstitial, (f) Sn interstitial, (g) O interstitial, and (h)  $MA_{Pb}$  cation substitution.

less than that of  $V<sub>I-1</sub>$  (0.63 eV), because the strength of Sn–I bonds is weaker than that of Pb−I bonds.  $V_O$  has a low formation energy of 0.01 eV at the  $MAI/SnO<sub>2</sub>$  interfaces, which is similar to that for the  $PbI_2/SnO_2$  interfaces and consistent with experimental observations. $67,68$  $67,68$  $67,68$  Compared with the  $PbI_2/SnO_2$  interfaces, the formation energies of  $I_i$  and  $O_i$ are becoming lower. This is because the  $MAI/SnO<sub>2</sub>$  interfaces have larger interstice space at the interfaces. For  $MA_{Pb}$  at the  $MAI/SnO<sub>2</sub>$  interfaces, its formation energy becomes larger due to that the MA molecule replacing a Pb atom is far away from the  $SnO<sub>2</sub>$  layer and could not be bonded to the dangling O atom. As for  $Sn_i$  at the  $MAI/SnO_2$  interfaces, we could also observe the similar phenomenon as the  $PbI_2/SnO_2$  interfaces. It has very low formation energy of −0.32 eV because of the free volume and the dangling O atoms. [See [Figure S2\]](https://pubs.acs.org/doi/suppl/10.1021/acsomega.2c01584/suppl_file/ao2c01584_si_001.pdf).

For these defects that might form at the  $MAI/SnO<sub>2</sub>$ interfaces, we also give the PDOSs in Figure 6 for further discussing their effects on the electron transport. From Figure 6a, it is noticed that there appear I-p and  $O-p$  levels near the middle of the band gap for the Sn−I antisite substitution defect. The deep levels in the band gap produced by the defect could attract electrons/holes and act as Shockley−Reid−Hall nonradiative recombination centers, which could adversely affect the efficiency of the solar cells. On the other hand, the

band edges of VBM and CBM are totally changed, which are not conducive to electron transfer from  $CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>$  to SnO<sub>2</sub>. In Figure 6b,  $V_{\Omega}$  creates a defect level (Sn-s) at  $E = -0.4$  eV, where the level is located just above the VBM of  $CH_3NH_3PbI_3$ . Because the electronic transport path does not involve the VBM of  $SnO<sub>2</sub>$ , the production of an oxygen vacancy could have no impact on the electron transport properties at the interfaces. From Figure 6c,d, we can see I-1 and I-2 vacancy defects have slightly moved the CBM of  $SnO<sub>2</sub>$  upward and have little effect on the electron transport process at the interfaces. For Sn<sub>i</sub> and  $I_i$  defects, they both create defect levels near the middle of the band gap. As shown in Figure 6e,f, the defect levels are composed of Sn-s, Sn-p, O-p, and I-p at  $E = 0.5$  eV for Sn<sub>i</sub> and I-p and Pb-p at  $E = 1.1$  eV for I<sub>i</sub>, respectively. Thus, Sn<sub>i</sub> and I<sub>i</sub> can be classified as harmful defects and the deep levels might become recombination centers for electrons and holes. Besides,  $Sn_i$  makes the Pb- $p$  shift to the CBM edge, which seriously influences the electron transfer from  $CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>$  to  $SnO<sub>2</sub>$ . Unlike other defects,  $O<sub>i</sub>$  at the interfaces does not create any defect levels but narrows the band gap of  $SnO<sub>2</sub>$  to 1.58 eV as observed from Figure 6g. The reduced band gap could decrease the open circuit voltage of PSCs and therefore affects the PCE of the solar cells.<sup>[70](#page-9-0)</sup> As for  $MA_{Pb}$  at the interfaces, it creates a defect level (Pb-p) at  $E = 3.2$  eV (Figure 6h). As the

<span id="page-7-0"></span>defect level is just located above the CBM of  $CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>$ , the effects on the electron transport can be neglected.

In general,  $V<sub>I</sub>$  and I<sub>i</sub> can be classified as the harmful defects, and  $\rm V_{O}$ ,  $\rm Sn_{\it i}$  and  $\rm MA_{\rm Pb}$  are the benign defects at the  $\rm PbI_2/SnO_2$ interfaces. At the MAI/SnO<sub>2</sub> interfaces, Sn−I, I<sub>i</sub>, and Sn<sub>i</sub> are harmful, and  $\rm V_{O}$ ,  $\rm V_{I}$ ,  $\rm O_{\it j}$ , and  $\rm MA_{Pb}$  are benign. Obviously, harmful defects are more easily formed at the  $MAI/SnO<sub>2</sub>$ interfaces.  $I_i$  is the most easily formed harmful defect at both  $PbI_2/SnO_2$  and  $MAI/SnO_2$  interfaces, which is consistent with the previous calculations and experimental results.<sup>71-[73](#page-9-0)</sup> The harmful defects affect the transport of electrons from  $CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>$  to SnO<sub>2</sub> or create deep levels that could become the recombination center of electrons/holes, thereby injuring the performance of PSCs. Therefore, minimizing the formation of the  $MAI/SnO<sub>2</sub>$  interfaces and avoiding the harmful defects could be beneficial to achieve high-performance and stable PSCs.

#### 4. CONCLUSION

In this paper, first-principles calculations have been performed to study structural and electronic properties of  $CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>/$  $SnO<sub>2</sub>$  interfaces by involving two different terminations (PbI<sub>2</sub>) and MAI). It is expected that the  $PbI_2/SnO_2$  interfaces have a high interfacial adhesion energy of 1.22  $J/m^2$  due to the interfacial Pb−O and Sn−I bonding as well as a stronger internal electric field for the electron−hole pairs separation. The effects of the stress and defects at the interfaces on the electron transport are also thoroughly addressed. The tensile stress could move the CBM of  $CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>$  upward, and the compressive stress could move the CBM of  $SnO<sub>2</sub>$  downward. By considering the stress effect, the CBM offset at the  $PbI_2/$  $SnO<sub>2</sub>$  interfaces is close to 0.07 eV, and at the MAI/SnO<sub>2</sub> interfaces is approximately 0.28 eV. Moreover, according to our research on possible interface defects,  $V<sub>L</sub>$ , and  $I<sub>i</sub>$  can be classified as the harmful defects at the  $PbI_2/SnO_2$  interfaces, and Sn–I, I<sub>i</sub> and Sn<sub>i</sub> are harmful at the MAI/SnO<sub>2</sub> interfaces. The harmful defects affect the transport of electrons from  $CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>$  to SnO<sub>2</sub> or create deep levels that could become the recombination center of electrons/holes. For both  $PbI_2/$  $SnO<sub>2</sub>$  and  $MAI/SnO<sub>2</sub>$  interfaces,  $I<sub>i</sub>$  is the most easily formed harmful defect and should be avoided in experiments. These results might contribute to the understanding of the  $CH<sub>3</sub>NH<sub>3</sub>PbI<sub>3</sub>/SnO<sub>2</sub>$  interfaces and provide some references for finding ways to improve the stability and photovoltaic performance of the PSCs.

# ■ ASSOCIATED CONTENT

#### **9** Supporting Information

The Supporting Information is available free of charge at [https://pubs.acs.org/doi/10.1021/acsomega.2c01584.](https://pubs.acs.org/doi/10.1021/acsomega.2c01584?goto=supporting-info)

> Figure S1, optimized stable geometrical structures of  $MA<sub>Pb</sub>$  at  $PbI<sub>2</sub>/SnO<sub>2</sub>$  interfaces; Figure S2, optimized stable geometrical structures of  $Sn_i$  at  $MAI/SnO_2$ interfaces ([PDF\)](https://pubs.acs.org/doi/suppl/10.1021/acsomega.2c01584/suppl_file/ao2c01584_si_001.pdf)

#### ■ AUTHOR INFORMATION

#### Corresponding Author

Wei Xiao − State Key Laboratory of Nonferrous Metals and Processes, GRINM Group Co., Ltd., Beijing 100088, PR China; GRIMAT Engineering Institute Co., Ltd., Beijing 101407, PR China; General Research Institute for

Nonferrous Metals, Beijing 100088, PR China; O[orcid.org/](https://orcid.org/0000-0001-9447-3279) [0000-0001-9447-3279](https://orcid.org/0000-0001-9447-3279); Email: [xiaowei@grinm.com](mailto:xiaowei@grinm.com)

### Authors

- Wenhua Pu <sup>−</sup> State Key Laboratory of Nonferrous Metals and Processes, GRINM Group Co., Ltd., Beijing 100088, PR China; GRIMAT Engineering Institute Co., Ltd., Beijing 101407, PR China; General Research Institute for Nonferrous Metals, Beijing 100088, PR China; Department of Materials Physics and Chemistry, School of Materials Science and Engineering, Key Laboratory for Anisotropy and Texture of Materials, Ministry of Education, Northeastern University, Shenyang 110819, PR China
- Jianwei Wang <sup>−</sup> State Key Laboratory of Nonferrous Metals and Processes, GRINM Group Co., Ltd., Beijing 100088, PR China; GRIMAT Engineering Institute Co., Ltd., Beijing 101407, PR China; General Research Institute for Nonferrous Metals, Beijing 100088, PR China
- Xiao-Wu Li <sup>−</sup> Department of Materials Physics and Chemistry, School of Materials Science and Engineering, Key Laboratory for Anisotropy and Texture of Materials, Ministry of Education, Northeastern University, Shenyang 110819, PR China; [orcid.org/0000-0002-0238-9107](https://orcid.org/0000-0002-0238-9107)
- Ligen Wang <sup>−</sup> State Key Laboratory of Nonferrous Metals and Processes, GRINM Group Co., Ltd., Beijing 100088, PR China; GRIMAT Engineering Institute Co., Ltd., Beijing 101407, PR China; General Research Institute for Nonferrous Metals, Beijing 100088, PR China

Complete contact information is available at: [https://pubs.acs.org/10.1021/acsomega.2c01584](https://pubs.acs.org/doi/10.1021/acsomega.2c01584?ref=pdf)

#### Notes

The authors declare no competing financial interest.

#### ■ ACKNOWLEDGMENTS

This work was supported by the National Natural Science Foundation of China (Grant No. 11704041) and GRINM Youth Foundation funded project.

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