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MHD Casson carbon nanotube flow with mass and heat transfer under thermosolutal Marangoni convection in a porous medium: analytical solution

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Current work portrays the flow of Marangoni convection Magneto hydrodynamics Casson fluid with carbon nanotubes under the effect of transpiration and radiation. The carbon nanotube particles namely water-single wall carbon nanotubes are inserted in the fluid to enhance better thermal efficiency. This type of flow problems is applicable for real life situations such as drying of silicon wafers, glues, crystal growth and heat exchangers and so on. The ordinary differential equations (ODEs) form of the result is yield to convert partial differential equations of the given equation by using similarity variables. Then this resulting ODEs are solved analytically, firstly using momentum equation to get solution domain and then by using this domain the energy equation solved to get the temperature profile in terms of Laguerre polynomial. Additionally, mass transpiration is also solved to get the concentration profile in terms of Laguerre polynomial. By using the different controlling parameters, the results can be discussed. And the effect of this parameters are discussed by using graphical arrangements. The newness of the present work is to explain the physically flow problem on the basis of chemically radiative thermosolutal Marangoni convective fluid.

Abbreviations

B. Cs Boundary conditions

- MHD Magneto hydrodynamics
- ODE Ordinary differential equation
- PDE Partial differential equation

List of symbols

- B_0 Strength of uniform magnetic field (Tesla)
- C Concentration of the fluid $(mol m^{-3})$
- C_{∞} Concentration of the fluid at infinity (mol m⁻³)
- C_w Concentration at wall (mol m⁻³)
- C_P Specific heat at constant pressure $(J \text{ Kg}^{-1} \text{ K}^{-1})$
- D Mass diffusivity (m² S⁻¹)
- Da^{-1} Inverse Darcy number
- *G* Chemical reaction parameter
- *I* Heat source/sink parameter
- *K* The permeability of porous medium (m^2)
- *L* Characteristic length
- Q Chandrasekhar's number
- Q₀ Heat source/sink coefficient

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- q_r Radiative heat flux
- \overline{T}_{∞} Free stream temperature (K)
- *T* Temperature of the fluid (K)
- (u, v) Components of velocities (ms^{-1})
- V_0 Mass transfer velocity (ms⁻¹)

Greek symbols

- α_r Mean absorption coefficient
- β Solution domain
- ϕ Concentration profile
- μ Dynamic viscosity (m² s⁻¹)
- θ Temperature profile
- κ Thermal conductivity (W/mK)
- σ Electrical conductivity (Sm⁻¹)
- σ_0 Equilibrium surface tension
- σ^* Stefan-Boltzmann constant (Wm⁻² K⁻⁴)
- ρ Density of the fluid (kg m⁻³)

Subscripts

- ∞ Free stream condition
- η Differentiation with respect to η

Basically in the interfaces of liquid–liquid or liquid–gas we found the layers of Marangoni convection, these layers are normally called as dissipative layers and these layers plays a great role in industrial applications. Gibbs¹ discovered this phenomenon in last century. Napolitano^{2,3} addressed the original work of this field. Temperature and concentration dependent surface tension is respectively called as Thermocapillary and destillocapillary effects^{4,5}. The application of Marangoni convection can be found in the fields of crystal growth, soap films and crystal growth. Chamkha et al.⁶ worked on Marangoni convection problem and he come to concluded that surface driven flows may build layers along the interfaces as well as buoyancy effect brought on by gravity and the external pressure gradient. When employing the arc length as coordinates, Napolitano et al.⁷ addressed the issue that non-Marangoni boundary layers in bulk fluids do not explicitly depend on the geometry of the interface. Only a few research and initiatives have been made to comprehend the fundamental laws of nature and the issues surrounding Marangoni convection.

The magnetohydrodynamics thermosolutal Marangoni convection over a flat surface in the presence of a heat source/sink parameter was addressed by Mudhaf and Chamkha⁸. The effects of heat transmission on MHD and radiation are examined by Aly and Ebaid⁹. Marangoni boundary layer nanofluid led him to conclude that a magnetic parameter causes a fluid's velocity to slow down and its temperature to rise. The double-diffusive convection in an open cavity was studied by Arbin et al.¹⁰ Nayak¹¹ investigated the magnetohydrodynamics viscoelastic fluid under the impact of chemical reaction effect with porous medium. See some other examples related to Marangoni convection are seen in^{12–15}.

Recent advancements in nanotechnology helps to conduct innovative techniques to develop applications of nanofluids in many fields. The term nanofluid is initially addressed by Choi¹⁶. Arshad et al.¹⁷⁻²⁰ investigated flow problems in the presence of different nanofluids with various aspects such as heat source/sink parameter, radiation and so on. See some more references on nanofluid in^{21,22}. Similarly, carbon nanotubes take a lot of attention in many fields such as chemistry, physics, medicine, biology, and Engineering. These are the few examples for the importance of nanofluids in many fields²³⁻²⁵. Additionally, the fluid known as Casson fluid is quite interesting, and it is used to describe non-Newtonian phenomena. The researcher Casson addressed this flow model in 1995. This flow model is useful in many real life applications. See some of the recent works of this model in²⁶⁻²⁸.

Effect of Porous medium and thermal radiation take places major role in the fluid flow because these effects in the fluid flow is used in many industrial and real life applications namely metallurgic processes, geophysical and allied areas²⁹. There are many equations and derivations are available to describe the fluid flow process through porous medium and also the effect of thermal radiation. See some more articles published on porous medium and thermal radiation are given in ³⁰⁻³².

The current study is investigating Casson fluid flow in the presence of carbon nanotubes with thermal radiation and mass transpiration. It is prompted by the aforementioned studies. In this problem we use the new method to provide a similarity variable on the impact of chemically radiative thermosolutal Marangoni convective fluid flow, the partial differential equations of the governing equations are converted into ordinary differential equations. The novelty of the present work explains that the momentum energy and mass equation solved analytically to get the solution domain and the solution in terms of Laguerre polynomial. The impact of different parameters is examined with the help of graphical scenario. This work is also important in many industrial applications such as welding machines, metallurgical process, geosciences, space technology and so on. The current issue is persuasively argued in the work of Mahabaleshwar et al.³³.



Figure 1. Schematic diagram of the Casson fluid flow.

Mathematical formulation and solution

Flow of a Casson fluid with thermosolutal Marangoni convection thermal radiation and transpiration is analyzed in this study. The particles of carbon nanotubes are immersed inside the fluid to get better thermal efficiency. Temperature gradients and solute concentrations define surface tension. Figure 1 shows a schematic representation of fluid flow.

Let us assume the surface of the fluid move towards x axis. The governing equations can be defined as follows by taking into account the aforementioned premises (See^{34,35}).

$$\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = 0, \tag{1}$$

$$u\frac{\partial u}{\partial x} + v\frac{\partial v}{\partial y} = v_{nf}\left(1 + \frac{1}{\Lambda}\right)\frac{\partial^2 u}{\partial y^2} - \left(\frac{\mu_{nf}}{\rho_{nf}K} + \frac{\sigma_{nf}B_0^2}{\rho_{nf}}\right),\tag{2}$$

$$u\frac{\partial T}{\partial x} + v\frac{\partial T}{\partial y} = \frac{\kappa_{nf}}{(\rho C_P)_{nf}} \frac{\partial^2 T}{\partial y^2} + \frac{\mu_{nf}}{(\rho C_P)_{nf}} \left(\frac{\partial u}{\partial y}\right)^2 + \left(\frac{\mu_{nf}}{(\rho C_P)_{nf}K} + \frac{\sigma_{nf}B_0^2}{(\rho C_P)_{nf}}\right) u^2 - \frac{1}{(\rho C_P)_{nf}} \frac{\partial q_r}{\partial y} + \frac{Q_0}{(\rho C_P)_{nf}} (T - T_\infty),$$
(3)

$$u\frac{\partial C}{\partial x} + v\frac{\partial C}{\partial y} = D\frac{\partial^2 C}{\partial y^2} - G(C - C_{\infty}), \qquad (4)$$

Here the Casson fluid term is used for characterize the non-Newtonian fluid. Magnetic term and porous medium term is used for many scientific and technological phenomena. Effect of Porous medium and thermal radiation take places major role in the fluid flow because these effects in the fluid flow is used in many industrial and real life applications. Heat source/sink in the fluid flow influences the characteristics of heat transfer as there is a substantial amount of difference in the temperature between the surface and the fluid. Also the combination of mass transfer and heat source/sink helps in overcoming the problem of boundary layer separation.

The surface tension along with heat and mass boundaries is given by (See³⁶⁻³⁸)

$$\sigma = \sigma_0 [1 - \gamma_T (T - T_\infty) - \gamma_C (C - C_\infty)], \tag{5}$$

Coefficients of surface tension respectively for heat and mass is given by

$$\gamma_T = -\frac{1}{\sigma_0} \left(\frac{\partial \sigma}{\partial T} \right)_T$$
, and $\gamma_C = -\frac{1}{\sigma_0} \left(\frac{\partial \sigma}{\partial C} \right)_T$. (6)

The terms from Eqs. (1-6) are specified in this section under Nomenclature. Associated B. Cs

$$\mu\left(\frac{\partial u}{\partial y}\right)_{y=0} = -\left(\frac{\partial \sigma}{\partial x}\right)_{y=0} = \sigma_0 \left(\gamma_T \left(\frac{\partial T}{\partial x}\right)_{y=0} + \gamma_C \left(\frac{\partial C}{\partial x}\right)_{y=0}\right),\tag{7}$$

$$V(x,0) = V_0, u(x,\infty) = 0$$

$$T(x,0) = T_{\infty} + T_0 X^2, T(x,\infty) = T_{\infty}$$

$$C(x,0) = C_{\infty} + C_0 X^2, C(x,\infty) = C_{\infty}$$
(8)

here $X = \frac{x}{L}$, and $L = -\frac{\mu v}{\sigma_0 T_0 \gamma_T}$ is the characteristic length, T_0 and C_0 are constants. In addition, the following transformations are defined.

$$\psi(x,y) = \nu X f(\eta) \quad \eta = \frac{y}{L}
T(x,y) = T_{\infty} + T_0 X^2 \theta(\eta)
C = C_{\infty} + C_0 X^2 \phi(\eta)$$
(9)

By using dimensional form of velocity components are given by

$$u = \frac{\nu}{L} f_{\eta}(\eta), \quad v = -\frac{\nu}{L} f(\eta).$$
(10)

The value q_r can be defined on the basis of Rosseland's approximation as follows (See³⁹⁻⁴².

$$q_r = -\frac{4\sigma^*}{3\alpha_r} \frac{\partial T^4}{\partial y},\tag{11}$$

where the ambient temperature T^4 expands in terms of the Taylor's series as

$$T^{4} = T_{\infty}^{4} + 4T_{\infty}^{3}(T - T_{\infty}) + 6T_{\infty}^{2}(T - T_{\infty})^{2} + \cdots$$
(12)

when the higher order elements in this equation are ignored, this results in

$$T^4 \cong 3T^3_\infty + 4T^3_\infty T. \tag{13}$$

On applying Eq. (8) into Eq. (6), then the first order derivative of heat flux can be given by

$$\frac{\partial q_r}{\partial y} = -\frac{16\sigma^* T_{\infty}^3}{3\alpha_r} \frac{\partial^2 T}{\partial y^2}.$$
(14)

Therefore, the Eq. (3) can b rewritten as

$$u\frac{\partial T}{\partial x} + v\frac{\partial T}{\partial y} = \frac{\kappa_{nf}}{(\rho C_P)_{nf}}\frac{\partial^2 T}{\partial y^2} + \frac{\mu_{nf}}{(\rho C_P)_{nf}} \left(\frac{\partial u}{\partial y}\right)^2 + \left(\frac{\mu_{nf}}{(\rho C_P)_{nf}K} + \frac{\sigma_{nf}B_0^2}{(\rho C_P)_{nf}}\right)u^2 + \frac{1}{(\rho C_P)_{nf}}\frac{16\sigma^*T_\infty^3}{3k^*}\frac{\partial^2 T}{\partial y^2} + \frac{Q_0}{(\rho C_P)_{nf}}(T - T_\infty).$$
(15)

By using Eqs. (9) and (10) in Eqs. (2) and (3) to get

$$\varepsilon_1 \left(1 + \frac{1}{\Lambda} \right) f_{\eta\eta\eta} + \varepsilon_2 \left(f f_{\eta\eta} - f_{\eta}^2 \right) - \left(\varepsilon_1 D a^{-1} + \varepsilon_3 Q \right) f_{\eta} = 0, \tag{16}$$

$$(\varepsilon_5 + R)\theta_{\eta\eta} + \Pr \varepsilon_4 \left(f \theta_\eta + \left(I - 2f_\eta \right) \theta \right) + Ec \left(\varepsilon_1 f_{\eta\eta}^2 + \left(\varepsilon_1 D a^{-1} + \varepsilon_3 Q \right) f_\eta \right) = 0, \tag{17}$$

$$\phi_{\eta\eta} + Sc \left(f \phi_{\eta} - \left(\delta + 2f_{\eta} \right) \phi \right) = 0, \tag{18}$$

the B. Cs reduces to

$$f(0) = V_C, \ f_\eta(\infty) = 0, \quad f_{\eta\eta}(0) = -2(1+M_a)$$

$$\theta(0) = 1, \ \theta(\infty) = 0, \ \phi(0) = 1, \ \phi(\infty) = 0,$$
(19)

here $V_C = -\frac{\gamma}{L}v_0$ is the mass transpiration, here $V_C = 0$, $V_C > 0$ and $V_C < 0$ respectively indicates suction, injection and no-permeability cases. Pr $= \frac{\kappa}{\mu C_p}$, $S_C = \frac{\nu}{D}$ and $\delta = \frac{GL^2}{\nu}$ respectively indicates the Prandtl number, Schmidt number, chemical reaction coefficient. $R = \frac{16\sigma^*T_\infty}{3\alpha_r\kappa}$ is the radiation number, $I = \frac{Q_0L^2}{\epsilon_4\rho C_p\nu}$ is the heat source or sink parameter, $Da^{-1} = \frac{L^2}{K}$ is inverse Darcy number, $Q = \frac{\sigma B_0L^2}{\rho\nu}$ is Chandrasekhar's number, $E_C = \frac{\gamma^2}{L^2 T_0 C_p}$ is Eckert number, and finally $Ma = \frac{Ma_C}{k}$ is the Marangoni number (Thermosolutal surface tension ratio), $Ma_C = \frac{\sigma_0\gamma_C C_0LC_p}{\kappa}$ and $Ma_T = \frac{\sigma_0\gamma_T T_0LC_p}{\kappa}$ are the solutal and thermal Marangoni numbers. Carbon nanofluid quantities used in Eqs. (16) and (17) can be defined as (See^{43,44})

$$\varepsilon_1 = \frac{\mu_{nf}}{\mu_f}, \ \varepsilon_2 = \frac{\rho_{nf}}{\rho_f}, \ \varepsilon_3 = \frac{\sigma_{nf}}{\sigma_f}, \ \varepsilon_4 = \frac{(\rho C_P)_{nf}}{(\rho C_P)_f}, \ \varepsilon_4 = \frac{\kappa_{nf}}{\kappa_f}$$

Exact solutions

Exact solution for momentum equation. Consider Eq. (16)'s solution, which has the following structure (See^{45})

$$f(\eta) = f_{\infty} + (V_C - f_{\infty}) Exp(-\beta\eta), \qquad (20)$$

here,
$$f_{\infty} = \beta - \frac{Da^{-1}}{\beta}$$
, (21)

Although, from the governing B. Cs $f(0) = V_C$, $f_{\eta}(\infty) = 0$, and $f_{\eta\eta}(0) = -2(1 + M_a)$ is simultaneously satisfied f_{∞} for $\beta > 0$ is as follows

$$f_{\infty} = V_C + \frac{2(1+M_a)}{\beta^2}.$$
 (22)

On applying Eqs. (20, 21) in Eq. (16) to yield the following cubic equation

$$\varepsilon_1 \left(1 + \frac{1}{\Lambda} \right) \beta^3 - \varepsilon_2 V_C \beta^2 - \left(\varepsilon_1 D a^{-1} + \varepsilon_3 Q \right) a - 2\varepsilon_2 (1 + M_a) = 0.$$
⁽²³⁾

Then the velocity can be required as

$$f_{\eta}(\eta) = -\beta \left(V_C - f_{\infty} \right) Exp(-\beta \eta).$$
(24)

Exact solution for temperature and concentration equation. For the purpose of solving temperature and concentration equation we introduce the following new variable for temperature and concertation respectively as follows

$$\xi = \left(\frac{\Pr\left(V_C - f_{\infty}\right)}{\beta(\varepsilon_5 + R)}\right) Exp(-\beta\eta), \text{ for temperature}$$
(25)

$$\varsigma = \left(\frac{Sc(V_C - f_{\infty})}{\beta}\right) Exp(-\beta\eta), \text{ for concentration}$$
(26)

Using Eqs. (25) and (26) respectively in Eqs. (17) and (18) to get

$$\xi \frac{\partial^2 \theta}{\partial \xi^2} + (1 - \varepsilon_4 S_1 - \varepsilon_4 \xi) \frac{\partial \theta}{\partial \xi} + \varepsilon_4 \left(2 - \frac{\gamma_1}{\xi}\right) \theta = -Ec_1 \xi, \tag{27}$$

$$\varsigma \frac{\partial^2 \phi}{\partial \varsigma^2} + \left(1 - p - \varsigma\right) \frac{\partial \phi}{\partial \varsigma} + \left(2 + \frac{q}{\varsigma}\right) \phi = 0, \tag{28}$$

here

$$S_{1} = \frac{\Pr f_{\infty}}{\beta(\varepsilon_{5} + R)}, \quad \gamma_{1} = \frac{I \Pr}{\beta(\varepsilon_{5} + R)}$$
$$Ec_{1} = -\frac{Ec\beta^{2}}{\Pr^{2}}(\varepsilon_{5} + R)(\varepsilon_{1}\beta^{2} + (\varepsilon_{1}Da^{-1} + \varepsilon_{3}Q))$$
$$p = \frac{Scf_{\infty}}{\beta} \quad q = -\frac{Sc\delta}{\beta^{2}},$$

The B. Cs are also reducing to

$$\theta(\xi = -1) = 1, \ \theta(\xi = 0) = 0,$$
(29)

$$\phi(\varsigma = -1) = 1, \quad \phi(\varsigma = 0) = 0, \tag{30}$$

on solving Eqs. (27) and (28) by using Frobenius method to yield the following equations

$$\theta(\eta) = (1 - A_3) Exp(-\beta A_2 \eta) \frac{L(2 - A_2, A_1, \varepsilon_4 \xi)}{L(2 - A_2, A_1, -\varepsilon_4 \xi_0)} + A_3 \xi^2,$$
(31)

$$\phi(\eta) = Exp(-\beta B_2 \eta) \frac{L(2 - B_2, B_1, \zeta)}{L(2 - B_2, B_1, \zeta_0)}.$$
(32)

where

$$\begin{aligned} A_{1} &= \sqrt{\varepsilon_{4}^{2}S_{1}^{2} + 4\varepsilon_{4}\gamma_{1}}, \quad A_{2} &= \frac{\varepsilon_{4}S_{1}}{2} + \frac{\sqrt{\varepsilon_{4}^{2}S_{1}^{2} + 4\varepsilon_{4}\gamma_{1}}}{2}, \quad A_{3} &= \frac{\varepsilon_{4}Ec_{1}}{\varepsilon_{4}(2S_{1} + \gamma_{1}) - 4} \\ B_{1} &= \sqrt{P^{2} - 4q}, \quad B_{2} &= \frac{P}{2} + \frac{\sqrt{P^{2} - 4q}}{2} \\ \xi_{0} &= \left(\frac{\Pr(V_{C} - f_{\infty})}{\beta(\varepsilon_{5} + R)}\right), \quad \zeta_{0} &= \left(\frac{Sc(V_{C} - f_{\infty})}{\beta}\right) \end{aligned}$$

Results and discussion

This article portrays the Casson fluid flow with Marangoni convection with Carbon nanoparticles are immersed in the fluid flow to enhance the thermal efficiency of the fluid. Analytical results are examined with the help of different controlling parameters namely Casson fluid parameter, inverse Darcy number, Chandrasekhar's number, Marangoni number and so on. The significant effect of Prandtl number, Schmidt number, chemically reaction coefficient and heat source/sink parameter on temperature, concentration and heat source/sink parameter is discussed as follows. The graphical scenario can be disused as follows.

Figure 2a,b demonstrated that the one of the results of Eq. (23), The red solid and dashed lines of the figure represents non-physical solution for various values of *Ma* and keeping other parameters with suitable values. The effect of the physical solution varied directly with V_C , Da^{-1} and *Ma*. Similarly, Fig. 3a,b portrays the plots of physical solution verses *Ma* for different Casson fluid parameter Λ for $V_C > 0$ and $V_C < 0$ cases respectively. Figures 4 and 5 represents the relation associated with velocity $f_\eta(0)$ with roots β_1 , $\beta_2 \beta_3$ for various values of *Ma*. The physical and nonphysical surfaces depending upon the positive and negative roots respectively. From Fig. 4a,b we observe that the Λ directly affected the surface velocity and *Ma*. Also from Fig. 5a–c it is cleared that the V_C is directly affected the surface velocity and *Ma*. Figure 6 indicates $f(\eta)$ verses η for different values of Λ . In this Fig. 6a represents suction case, Fig. 6b indicates injection case and Fig. 6c indicates no permeability cases. Figure 7a–c indicates $f_\eta(\eta)$ verses η for various values of Λ for $V_C > 0$, $V_C < 0$ and $V_C = 0$ respectively. From this it is cleared that $f_\eta(\eta)$ decreases with increasing the values of Λ for $V_C > 0$, $V_C < 0$ and $V_C = 0$.

Impact of $f(\eta)$ verses η and $f_{\eta}(\eta)$ verses η for various values of M_a is respectively indicated at Fig. 8a,b for $V_C > 0$, and keeping all other parameters with suitable values. Here $f(\eta)$ and $f_{\eta}(\eta)$ is more for more values of M_a for $V_C > 0$. Figure 9a,b indicates $f(\eta)$ verses η and $f_{\eta}(\eta)$ verses η for different values of M_a at $V_C < 0$ respectively. Here $f(\eta)$ is more for more values of M_a for injection case. Also $f_{\eta}(\eta)$ less for more values of M_a for $V_C < 0$. Figure 10a,b portrays the $f(\eta)$ verses η and $f_{\eta}(\eta)$ verses η for various values of V_C respectively. Here $f(\eta)$ is more for more values of Λ_a for $V_C < 0$. Figure 10a,b portrays the $f(\eta)$ verses η and $f_{\eta}(\eta)$ verses η for various values of V_C respectively. Here $f(\eta)$ is more for more values of Λ_a to $V_C < 0$. Figure 10a,b portrays the $f(\eta)$ verses η and $f_{\eta}(\eta)$ verses η for various values of V_C respectively. Here $f(\eta)$ is more for more values of Λ_a to $V_c < 0$. The effect of $\theta(\eta)$ on η for various values of Λ , I, and Da^{-1} is respectively represented at Fig. 11a–c. Here, $\theta(\eta)$

The effect of $\theta(\eta)$ on η for various values of Λ , I, and Da^{-1} is respectively represented at Fig. 11a–c. Here, $\theta(\eta)$ more for more values of Λ and Da^{-1} , but $\theta(\eta)$ less for more values of I. And also we observe that after certain values of Λ lines are merging each other. The effect of $\phi(\eta)$ on η for different choices of Sc, δ , and V_C is respectively represented at Fig. 12a–c. from these graphs it is cleared that $\phi(\eta)$ decreases with increasing the values of Sc, δ , and V_C . The inclusion of porous media, heat source/sink parameter, thermal radiation and mass transpiration greatly useful in many fields, porous media prevents heat loss/gain and also accelerates the heat source/sink. Heat source/sink results in thinning of the thermal boundary, Marangoni convection results in more induced flows.

Conclusion

The investigation of results from the 2-D Casson fluid with mass transpiration, thermal radiation and chemically reaction parameter. The ODEs of equations are yielded when we mapped PDEs equation with similarity variables. These ODE equations are solved exactly then the momentum equation is solved to get solution domain, this domain is used in energy and concentration equation to get the temperature profile and concentration profile. The outlook of the present work explains the importance of porous media, thermal radiation, Marangoni convection, thermal radiation and heat source/sink parameter in the physically modelling of the flow. The outcomes we discovered using the graphical scenario are as follows.

- 1. Effect of the physical solution is directly affected by the V_C , Da^{-1} and Q.
- 2. A and V_C directly affected the surface velocity and *Ma*.
- 3. $f(\eta)$ and $f_{\eta}(\eta)$ decreases for the instance of suction, injection and no permeability case for rising the values of Λ for both suction, injection and no permeability cases.
- 4. $f(\eta)$ and $f_n(\eta)$ more if we rising the values of M_a for suction case.
- 5. $f(\eta)$ more for more values of M_a for suction case. But $f_{\eta}(\eta)$ less for more values of M_a for injection case. And $f(\eta)$ more for more values of V_C but $f_{\eta}(\eta)$ less for more values of V_C .
- 6. $\theta(\eta)$ more for more values of Λ and Da^{-1} , but $\theta(\eta)$ less for more values of *I*.
- 7. $\phi(\eta)$ less for more values of *Sc*, δ , and *V_C*.
- 8. Presence of porous media, prevents heat loss/gain and also accelerates the heat source/sink. Chemical reaction term thinning the thermal boundary, Marangoni convection results in more induced flows.
- 9. The future perspectives of the present work motivate to explain the physically flow problem on the basis of chemically radiative thermosolutal Marangoni convective fluid also helps to conduct flow problems with porous media.
- 10. Following conditions explain the comparison pf present work with previous works.





- a. If $Q = \phi = Ec = 0$ (In Eqs. 16 and 17) \Rightarrow Mahabaleshwar et al.³³. b. If $Q = \phi = Ec = R = 0$ (In Eqs. 16 and 17) \Rightarrow Mudhaf and Chamkha⁸.
- c. If $Q = \phi = Ec = Da^{-1} = 0$ (In Eqs. 16 and 17) \Rightarrow Magyari and Chamkha¹⁴,



Figure 3. The plots of β_1 , β_2 , β_3 verses *Ma* for $\Lambda = 5$ and $\Lambda = 1$ cases at (**a**) $V_C = 3$, and (**b**) $V_C = -3$.







Figure 5. The plots of $f_{\eta}(0)$ verses Ma for $\Lambda = 5$ and $\Lambda = \infty$ at (a) $V_C = 3$, (b) $V_C = -3$ and (c) $V_C = 0$.



Figure 6. $f(\eta)$ verses η for various values of Λ at (a) $V_C = 3$, (b) $V_C = -3$ and (c) $V_C = 0$.



Figure 7. $f_{\eta}(\eta)$ verses η for various values of Λ at (a) $V_C = 3$, (b) $V_C = -3$ and (c) $V_C = 0$.









Figure 9. The plots of (a) $f(\eta)$ verses η and (b) $f_{\eta}(\eta)$ verses η for different choices of M_a at injection case.



Figure 10. The plots of (a) $f(\eta)$ verses η and (b) $f_{\eta}(\eta)$ verses η for different choices of V_C .







Figure 12. The plots of $\phi(\eta)$ verses η for different choices of (a) Sc (b) δ and (c) V_C .

Data availability

The datasets used and/or analysed during the current study available from the corresponding author on reasonable request.

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Author contributions

All authors participated in all sections of numerical modeling, results and analysis.

Competing interests

The authors declare no competing interests.

Additional information

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