



OPEN

# High pressure Raman spectroscopy of H<sub>2</sub>O-CH<sub>3</sub>OH mixtures

SUBJECT AREAS:

GEODYNAMICS

PLANETARY SCIENCE

PHASE TRANSITIONS AND  
CRITICAL PHENOMENA

Wen-Pin Hsieh &amp; Yu-Hsiang Chien

Institute of Earth Sciences, Academia Sinica, Nankang, Taipei 11529, Taiwan.

Received  
18 November 2014Accepted  
23 January 2015Published  
23 February 2015Correspondence and  
requests for materials  
should be addressed toW.P.H. (wphsieh@  
earth.sinica.edu.tw)

Complex intra-molecular interactions and the hydrogen-bonding network in H<sub>2</sub>O-volatile mixtures play critical roles in many dynamics processes in physical chemistry, biology, and Earth and planetary sciences. We used high pressure Raman spectroscopy to study the pressure evolution of vibrational frequencies and bonding behavior in H<sub>2</sub>O-CH<sub>3</sub>OH mixtures. We found that the presence of low CH<sub>3</sub>OH content in H<sub>2</sub>O increases the transition pressure where water crystallizes to ice VI, but does not significantly change the pressure where ice VI transforms to ice VII. Furthermore, the stiffening rates of C-H stretching frequencies  $dw/dP$  in CH<sub>3</sub>OH significantly decrease upon the crystallization of water, and the softening rates of the O-H stretching frequencies of ice VII are suppressed over a narrow pressure range, after which the frequencies of these modes shift with pressure in ways similar to pure CH<sub>3</sub>OH and ice VII, respectively. Such complex pressure evolution of Raman frequencies along with pronounced variations in Raman intensities of CH<sub>3</sub>OH within the sample, and the hysteresis of the water-ice VI phase transition suggest pressure-induced segregation of low content CH<sub>3</sub>OH from ice VII. These findings indicate the significant influence of volatiles on the crystallization of sub-surface ocean and thermal evolution within large icy planets and satellites.

H<sub>2</sub>O is the archetypal hydrogen-bonded and the most abundant polyatomic molecule in the universe<sup>1</sup>. It displays a rich pressure-temperature phase diagram<sup>2</sup>, and the evolution of its molecular vibrational behavior under extreme pressures has been a major area of broad research interest<sup>3-7</sup>. In addition to pure H<sub>2</sub>O, knowledge of the high pressure physical and chemical properties of H<sub>2</sub>O mixed with other volatiles, such as methanol (CH<sub>3</sub>OH), methane (CH<sub>4</sub>), ammonia (NH<sub>3</sub>), and carbon dioxide (CO<sub>2</sub>), is critical to understanding many phenomena in Earth and planetary systems, e.g., the effect of volatiles on the geodynamics and thermal history of icy satellites<sup>8-12</sup>. Among simple volatiles, CH<sub>3</sub>OH is also a prototypical, polar and hydrogen-bonded molecule; when mixed with water it serves as an anti-freeze compound that lowers the freezing temperature of water at ambient pressure<sup>11</sup>. It has been shown that the presence of CH<sub>3</sub>OH in H<sub>2</sub>O could affect the crystallization behavior of the primordial ocean of icy satellites and determine whether a sub-surface ocean can be formed<sup>10</sup>.

The inter- and intra-molecular interactions and molecular structure of H<sub>2</sub>O-CH<sub>3</sub>OH mixtures with various relative concentrations have been extensively studied using computer simulations<sup>13-19</sup>, neutron and X-ray diffraction and spectroscopy techniques<sup>20-25</sup>, and Raman spectroscopy<sup>26-29</sup>. At ambient conditions, the H<sub>2</sub>O water is composed of a tetrahedral, hydrogen-bonded molecular structure<sup>30</sup>, whereas the CH<sub>3</sub>OH is composed of hydrogen-bonded chains or rings<sup>20</sup>. Upon mixing, the local arrangement and structure of these molecules deviate from those of each pure phase and could change with different relative concentrations due to the complex hydrogen-bonded, hydrophobic, and hydrophilic interactions. For instance, recent studies have shown that at ambient conditions, a mixture with higher than 50% molar fraction of CH<sub>3</sub>OH exhibits incomplete mixing at the molecular level, i.e., H<sub>2</sub>O molecules form clusters and segregate from the CH<sub>3</sub>OH<sup>15,20,21</sup>; the molecular segregation is enhanced by cooling and moderate compression due to the formation of larger water clusters<sup>24</sup>. However, previous studies have largely been performed near ambient conditions; the effect of extreme pressure on the local molecular bonding and structure have been rarely investigated and only up to the hundreds of MPa range, much lower than the relevant conditions in planetary and icy satellite interiors, where H<sub>2</sub>O is typically mixed with other volatiles (less than 20 wt%) under a few GPa pressures<sup>9,10</sup>.

Raman spectroscopy enables the probing of local vibrational modes of H<sub>2</sub>O-CH<sub>3</sub>OH mixtures and thus offers a powerful means for monitoring the evolution of both intra- and inter-molecular bonding under high pressure conditions that is complementary to X-ray and neutron diffraction techniques. In this work, we present a high pressure Raman spectroscopy study of H<sub>2</sub>O mixed with less than 20 wt% (<12% molar fraction) CH<sub>3</sub>OH at room temperature to approximately 10 GPa, a pressure range over which the liquid water undergoes pressure-induced phase transitions to crystalline ice VI and ice VII phases. Pressure tuning of the inter- and intra-molecular interactions and structures over a broad pressure range allows us to characterize the evolution of molecular



vibrational frequencies and the bonding behavior of the mixtures. We observed a significant influence of CH<sub>3</sub>OH on the crystallization of compressed H<sub>2</sub>O at high pressure, i.e., the pressure where liquid H<sub>2</sub>O crystallizes to ice VI at room temperature increases with increasing CH<sub>3</sub>OH content, whereas upon decompression ice VI melts to water at nearly the same pressure as in pure H<sub>2</sub>O, independent of the CH<sub>3</sub>OH content. In addition, it is well known that pressure typically enhances the inter-molecular interactions and could create additional bonding between different species, synthesizing novel molecular compounds<sup>31–34</sup>. In our experiments we observed intricate evolution of the C-H stretching frequencies of CH<sub>3</sub>OH and the O-H stretching frequencies of ice VII under pressure as well as pressure-induced variations in Raman intensities of the C-H stretching modes of CH<sub>3</sub>OH within the sample. These findings combined with the hysteresis of the water-ice VI phase transition suggest pressure-induced segregation of low content CH<sub>3</sub>OH from the ice VII.

## Results

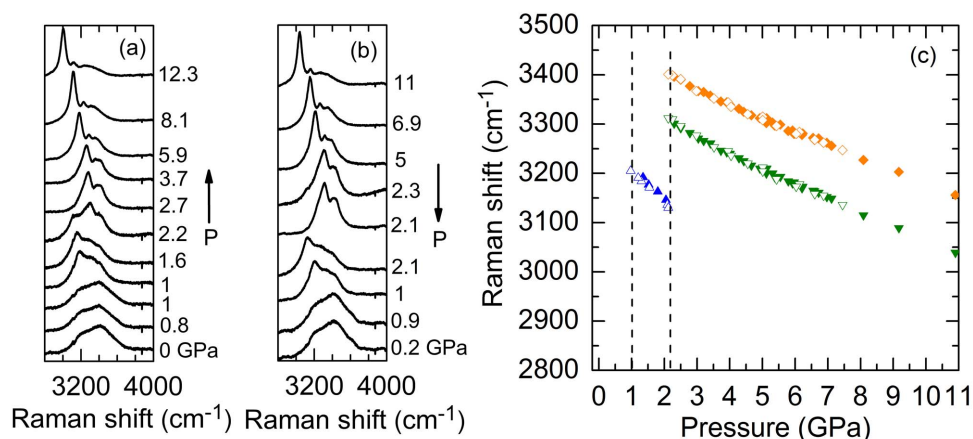
Figures 1a and 1b show representative Raman spectra of H<sub>2</sub>O at room temperature as a function of pressure during compression and decompression cycles. We focused on the frequency range of 2800 to 4000 cm<sup>-1</sup> and observed two broad Raman bands around 3235 and 3400 cm<sup>-1</sup> due to the O-H stretching modes of liquid water. Upon increasing pressure, water then crystallized to ice VI at 1 GPa accompanied by the appearance of a characteristic, intense Raman mode around 3180 cm<sup>-1</sup>. Ice VI further transformed to ice VII at approximately 2.2 GPa, and coexistence of the ice VI and VII was observed during a pressure range from 2.2 to 2.3 GPa (Fig. 1a). Upon decompression, the ice VII reversibly returned to ice VI at 2.1 GPa, which then transformed back to the liquid phase at 0.9 GPa. In Fig. 1c, we plot the pressure dependent Raman shifts of the O-H stretching modes of ice VI around 3200 cm<sup>-1</sup> and ice VII around 3300 and 3400 cm<sup>-1</sup>. The vibrational frequencies of the O-H stretching modes all red-shifted with increasing pressure due to the strong hydrogen-bonded network between the H<sub>2</sub>O molecules. All the frequency shifts were reversible upon decompression. The pressure dependent Raman spectra, phase transition pressures between liquid and crystalline phases, and Raman shifts under pressure are well consistent with previous studies<sup>6,7,35</sup>.

The representative pressure dependence of Raman spectra of CH<sub>3</sub>OH, the other end member of H<sub>2</sub>O-CH<sub>3</sub>OH mixture, is shown in Figs. 2a–d. At ambient pressure, we observed four dominant modes, including a C-O stretching mode at 1031 cm<sup>-1</sup>, a C-H bend-

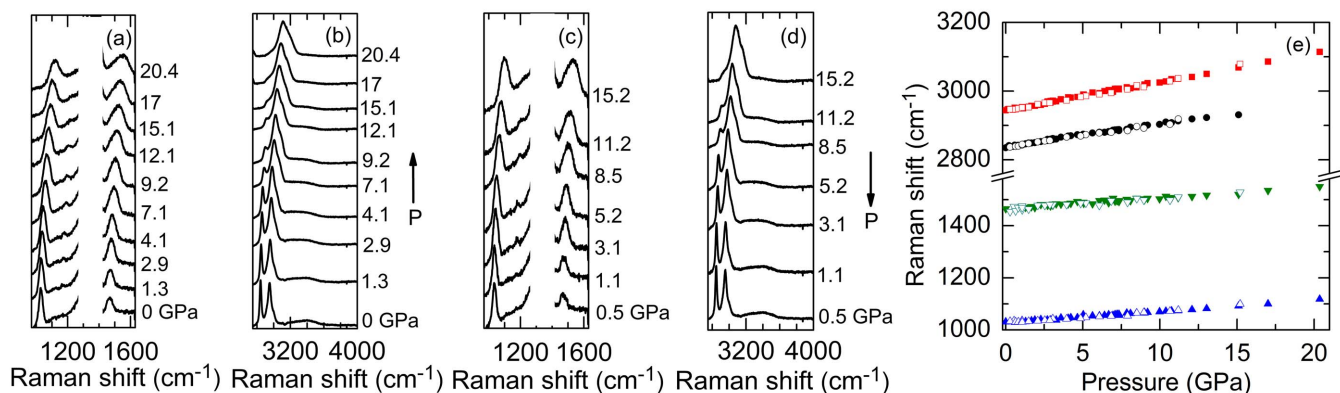
ing mode at 1465 cm<sup>-1</sup>, and two C-H stretching modes at 2835 cm<sup>-1</sup> (symmetric stretching) and 2944 cm<sup>-1</sup> (anti-symmetric stretching) that are due to the Fermi resonance<sup>36</sup>. In addition, we also detected small, broad bands around 3300 and 3400 cm<sup>-1</sup> due to the O-H stretching modes. All the characteristic Raman modes are in good agreement with literature results<sup>26,29,36–38</sup>. Upon compression, all the dominant modes blue-shifted and broadened with increasing pressure<sup>38</sup>. Moreover, the ratio of the Raman intensity of the C-H anti-symmetric stretching to symmetric stretching mode increased dramatically, resulting in a broad spectrum when the pressure is higher than 17 GPa. During the decompression cycle, a reversible evolution with pressure was observed where the C-O stretching and C-H bending modes returned to the initial spectrum before compression, and the pressure-induced broad spectrum of the C-H stretching modes transformed back to the starting two peaks with the same initial Raman intensity ratio.

We summarize the Raman shifts of the four dominant modes of CH<sub>3</sub>OH as a function of pressure in Fig. 2e. All the frequencies shifted linearly with pressure during both the compression (filled symbols) and decompression (open symbols) cycles. We did not observe mode splitting and discontinuous frequency shifts that are signatures of the pressure-induced crystallization of liquid CH<sub>3</sub>OH<sup>37</sup>, indicating that in our experiments the CH<sub>3</sub>OH was superpressed and did not transform into a crystalline phase, as reported in previous studies that under pressure CH<sub>3</sub>OH could be difficult to crystallize and usually forms a superpressed liquid<sup>37,39</sup>.

Addition of CH<sub>3</sub>OH to H<sub>2</sub>O changes the local molecular bonding in both species and therefore, the phase transition behavior under pressure. The representative pressure dependence of Raman spectra of H<sub>2</sub>O mixed with 2.5 wt% CH<sub>3</sub>OH at room temperature during compression and decompression cycles are shown in Figs. 3a and 3b. At ambient pressure we observed two strong and broad Raman bands of liquid water around 3240 and 3410 cm<sup>-1</sup> as well as two small peaks at 2844 and 2952 cm<sup>-1</sup> due to the C-H stretching of CH<sub>3</sub>OH, in good agreement with literature<sup>26,40</sup>. At 1.3 GPa the liquid water crystallized to the ice VI, which then transformed to the ice VII phase at 2.3 GPa. The higher crystallization pressure compared to pure H<sub>2</sub>O indicates that the CH<sub>3</sub>OH serves as an anti-freeze compound that stabilizes liquid water (expanding the pressure range of stability of water) and delays its crystallization under pressure, similar to its anti-freeze effect whereby it lowers the freezing temperature of liquid water at ambient pressure. During decompression, the ice VII reversibly returned to ice VI at 2.2 GPa, whereas the ice VI transformed back to the liquid water at 1.1 GPa. Since the uncertainties of the pressure



**Figure 1** | Representative pressure dependence of the Raman spectra of H<sub>2</sub>O during (a) compression and (b) decompression cycles. Pressure is labeled next to each spectrum which is displaced vertically for clarity. During the compression cycle, the liquid water crystallizes to ice VI at 1 GPa and ice VI transforms to ice VII at 2.2 GPa where coexistence of the two phases is observed in the spectrum. (c) Raman shifts of the O-H stretching modes of H<sub>2</sub>O as a function of pressure during compression (filled symbols) and decompression (open symbols) cycles. The dashed lines mark the phase transition pressures for compressed water-ice VI and ice VI–VII.



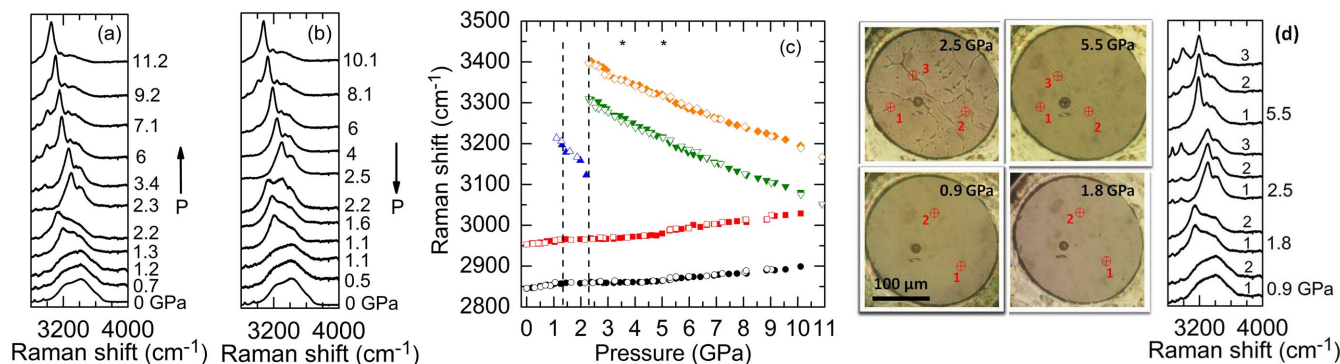
**Figure 2** | Representative pressure dependence of the Raman spectra of  $\text{CH}_3\text{OH}$  during (a) (b) compression and (c) (d) decompression cycles. Pressure is labeled next to each spectra which are displaced vertically for clarity. (e) Raman shifts of the  $\text{CH}_3\text{OH}$  for C-O stretching around  $1030\text{ cm}^{-1}$  (upper triangles), C-H bending around  $1465\text{ cm}^{-1}$  (lower triangles), and C-H stretching around  $2835\text{ cm}^{-1}$  (circles) and  $2944\text{ cm}^{-1}$  (squares) during compression (filled symbols) and decompression (open symbols) cycles. The frequencies of all four modes shifted approximately linearly and reversibly over the pressure range studied. In (a) and (c), the strong Raman signal from the diamond anvils around  $1330\text{ cm}^{-1}$  is removed for clarity.

measurements are typically  $\sim 0.1\text{ GPa}$ , the slightly lower transition pressure from ice VI to water compared to that during compression ( $1.3\text{ GPa}$ ) indicates hysteresis water-ice VI transition under pressure due to the presence of  $\text{CH}_3\text{OH}$ . This hysteresis effect is enhanced with higher  $\text{CH}_3\text{OH}$  contents as presented below.

Figure 3c summarizes the pressure dependent Raman shifts of the O-H stretching modes of ice VI and VII as well as the C-H stretching modes of  $\text{CH}_3\text{OH}$  for  $\text{H}_2\text{O}$  mixed with 2.5 wt%  $\text{CH}_3\text{OH}$ . Three experimental runs yielded consistent results. The frequencies of the two C-H stretching modes initially increased linearly with pressure, but the slopes  $d\omega/dP$  substantially decreased around  $1.3\text{ GPa}$ , when the liquid water crystallized to ice VI. Similar to the pressure dependence of the Raman shifts in pure  $\text{H}_2\text{O}$ , the frequencies of the O-H stretching modes of ice VI and VII all red-shifted with increasing pressure. The softening rates of the O-H stretching modes of ice VII, however, were suppressed during a pressure range from approxi-

mately 3.5 to 5 GPa (labeled by two stars in Fig. 3c). After 5 GPa, the softening rates increased and became similar to the softening rates of pure ice VII; meanwhile the C-H frequencies of  $\text{CH}_3\text{OH}$  started to increase linearly with larger slopes than those during 1.3 to 5 GPa. During the decompression cycle, all the frequency shifts were reversible, except that around 1.1 GPa the O-H frequency of ice VI was still observed due to the delayed transition from ice VI to liquid water, and that when  $P < 1\text{ GPa}$  the small slopes for the C-H frequencies with pressure,  $d\omega/dP$ , increased and became similar to the slopes during initial compression.

During the compression cycle, we also observed clear changes in the distribution of the mixture within the sample chamber of the diamond anvil cell (DAC) (Fig. 3d). When the  $\text{H}_2\text{O}$  is in the liquid phase ( $P < 1.3\text{ GPa}$ ), the 2.5 wt%  $\text{CH}_3\text{OH}$  mixture was transparent and the Raman spectra taken at different locations within the DAC were nearly the same, see for instance, the photomicrograph and



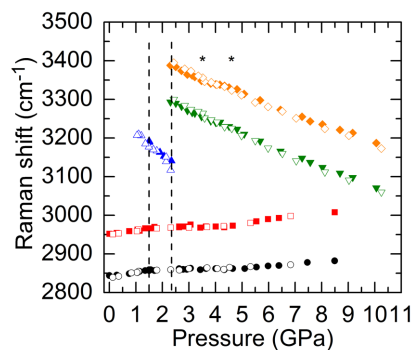
**Figure 3** | Representative pressure dependence of the Raman spectra of  $\text{H}_2\text{O}$  mixed with 2.5 wt%  $\text{CH}_3\text{OH}$  during (a) compression and (b) decompression cycles. Pressure is labeled next to each spectrum which is displaced vertically for clarity. With the addition of 2.5 wt%  $\text{CH}_3\text{OH}$ , the compressed liquid water crystallizes to ice VI at  $1.3\text{ GPa}$  and ice VI transforms to ice VII at  $2.3\text{ GPa}$ . (c) Raman shifts of the O-H stretching modes of  $\text{H}_2\text{O}$  and the C-H stretching modes of  $\text{CH}_3\text{OH}$  as a function of pressure during compression (filled symbols) and decompression (open symbols) cycles. The slopes  $d\omega/dP$  of the C-H stretching modes decrease upon crystallization of the compressed liquid water and then increase around 5 GPa. The softening rates of the O-H stretching modes of ice VII decrease around 3.5 GPa and slightly increase after 5 GPa. The two stars indicate the pressure range over which the reduction in the softening rates of the O-H modes occurs. The dashed lines mark the phase transition pressures for compressed water-ice VI and ice VI-VII, respectively. (d) Representative photomicrographs of the 2.5 wt%  $\text{CH}_3\text{OH}$  mixture at high pressures. The cross-circles in each photomicrograph mark the locations where the Raman spectra were taken within the DAC and their corresponding spectra numbered above each one are shown in the right panel. Significant variations in the Raman intensities of the C-H stretching modes of  $\text{CH}_3\text{OH}$  and the O-H stretching modes of ice VII are observed at 2.5 and 5.5 GPa. Note that the Raman spectra were taken at locations similarly distant from the center of the DAC to minimize the pressure gradient among different locations. The minimal pressure difference is confirmed by the observation that the Raman shifts measured at different locations are essentially identical.



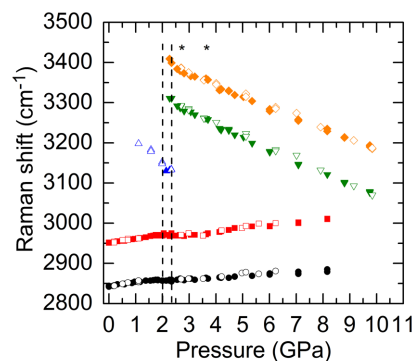
spectra at 0.9 GPa in Fig. 3d. However, upon crystallization of water to the ice VI and VII, the Raman spectra taken at different locations within the DAC showed pronounced variations in the intensities of the C-H stretching modes of CH<sub>3</sub>OH and the O-H stretching modes of ice VI and VII, e.g., the spectra at 2.5 and 5.5 GPa in Fig. 3d, indicating a pressure-induced inhomogeneous distribution that large amount of CH<sub>3</sub>OH molecules likely aggregates at the boundaries of the ice domains. Such pressure-induced molecular segregation was also observed in all other higher CH<sub>3</sub>OH content mixtures as discussed below.

The transition pressure and pressure dependent Raman shifts were further investigated with H<sub>2</sub>O mixed with 5, 10, and 20 wt% CH<sub>3</sub>OH. Similar to the 2.5 wt% CH<sub>3</sub>OH mixture, at ambient pressure all these mixtures also showed two broad bands of liquid water around 3240 and 3410 cm<sup>-1</sup> and two small peaks at 2844 and 2952 cm<sup>-1</sup> thanks to the C-H stretching of CH<sub>3</sub>OH. During compression, the pressure where water crystallized to ice VI was further raised with higher CH<sub>3</sub>OH content, i.e., it increased to 1.5 GPa for 5 wt% CH<sub>3</sub>OH, 2 GPa for 10 wt% CH<sub>3</sub>OH, and 2.1 GPa for 20 wt% CH<sub>3</sub>OH, respectively. The ice VI–VII transition, however, all occurred around 2.3 to 2.4 GPa, the same as that in 2.5 wt% CH<sub>3</sub>OH mixture. When releasing pressure, the ice VII reversibly transformed back to ice VI at 2.3 GPa in all these mixtures, but ice VI underwent a hysteresis phase transition and melted to liquid water at pressures around 0.8 to 1.1 GPa, much lower than the crystallization pressure during compression.

The pressure dependent Raman shifts of the O-H stretching modes of ice VI and VII and the C-H stretching modes of CH<sub>3</sub>OH for H<sub>2</sub>O mixed with 5, 10, and 20 wt% CH<sub>3</sub>OH are plotted in Figs. 4, 5, and 6, respectively. Similar to the H<sub>2</sub>O–2.5 wt% CH<sub>3</sub>OH mixture, the frequencies of the two C-H stretching modes in all these mixtures initially increased linearly with pressure until the pressures where liquid water crystallized to ice VI and then the slopes  $d\omega/dP$  became very small. The O-H stretching modes of ice VI and VII, on the other hand, all softened with pressure. The softening rates for the ice VII were also suppressed during a narrow pressure range (labeled by two stars in Figs. 4, 5, and 6), after which they softened with larger slopes similar to the pure ice VII, and meanwhile the small slopes of the C-H stretching modes of CH<sub>3</sub>OH increased. Note that the suppression in the softening rates of O-H modes for ice VII in 5 wt% CH<sub>3</sub>OH mixture occurs at higher pressure regime compared to 10 and 20 wt% CH<sub>3</sub>OH mixtures. During decompression, all frequencies



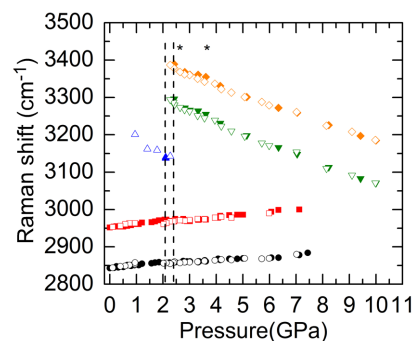
**Figure 4** | Pressure dependent Raman shifts of the O-H stretching modes of H<sub>2</sub>O and the C-H stretching modes of CH<sub>3</sub>OH during compression (filled symbols) and decompression (open symbols) cycles for H<sub>2</sub>O mixed with 5 wt% CH<sub>3</sub>OH. The stiffening rates of the C-H stretching modes decrease around 1.5 GPa when water crystallizes to ice VI, and then increase again around 4.6 GPa with larger slopes than those between 1.5 and 4.6 GPa. The softening rates of the O-H stretching modes for ice VII are reduced during a pressure range from approximately 3.4 to 4.6 GPa, which is labeled by two stars. The dashed lines mark the phase transition pressures for compressed water–ice VI and ice VI–VII, respectively.



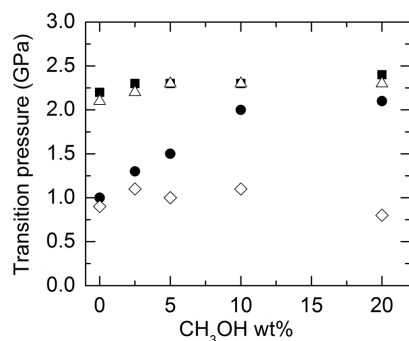
**Figure 5** | Raman shifts of the O-H stretching modes of H<sub>2</sub>O and the C-H stretching modes of CH<sub>3</sub>OH during compression (filled symbols) and decompression (open symbols) cycles for H<sub>2</sub>O mixed with 10 wt% CH<sub>3</sub>OH. The stiffening rates of the C-H stretching modes of CH<sub>3</sub>OH decrease upon the crystallization of water and then increase around 3.6 GPa with larger slopes. On the other hand, the softening rates of the O-H stretching modes for ice VII experience a reduction between approximately 2.7 and 3.6 GPa (labeled by two stars). The dashed lines mark the phase transition pressures for compressed water–ice VI and ice VI–VII, respectively.

reversibly shifted until the pressures where compressed water crystallized to ice VI, below which the ice VI phase was still present due to the delayed ice VI–water transition; when  $P < 1$  GPa the small slopes of C-H frequencies in all these mixtures increased to similar slopes upon initial compression.

The transition pressures for water–ice VI–ice VII at room temperature during compression and decompression cycles as a function of CH<sub>3</sub>OH wt% is summarized in Fig. 7. The crystallization pressure of water (filled circles) increases with increasing CH<sub>3</sub>OH wt% and saturates when the concentration of CH<sub>3</sub>OH is higher than 10 wt%. The transition pressure from ice VI to VII (filled squares), however, depends weakly on the concentration of CH<sub>3</sub>OH. When decreasing pressure, the ice VII in all the mixtures reversibly transformed back to ice VI at nearly the same pressures as during compression (open triangles), independent of the CH<sub>3</sub>OH wt%. The ice VI melted to water around 1 GPa (open diamonds) for all of the CH<sub>3</sub>OH wt% we studied, indicating a pronounced hysteresis for the water–ice VI phase transition in the mixtures.



**Figure 6** | Raman shifts of the O-H stretching modes of H<sub>2</sub>O and the C-H stretching modes of CH<sub>3</sub>OH during compression (filled symbols) and decompression (open symbols) cycles for H<sub>2</sub>O mixed with 20 wt% CH<sub>3</sub>OH. The slopes  $d\omega/dP$  of the C-H modes decrease when the compressed water crystallizes to ice VI and then increase after 3.6 GPa. The softening rates of the O-H modes decrease between approximately 2.6 and 3.6 GPa (labeled by two stars), after which they increase and become similar to those of pure ice VII. The dashed lines mark the phase transition pressures for compressed water–ice VI and ice VI–VII, respectively.



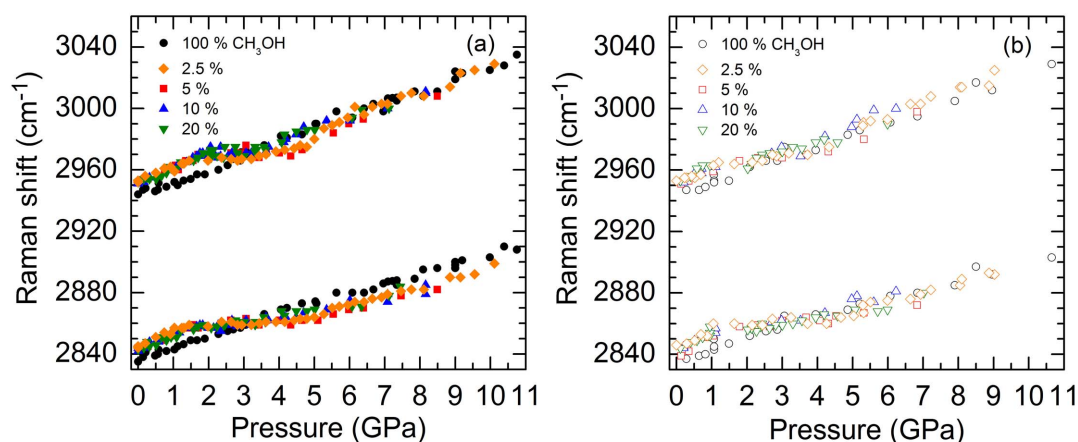
**Figure 7** | Transition pressures for water-ice VI and ice VI-VII as a function of  $\text{CH}_3\text{OH}$  wt%. In the compression cycle, the crystallization pressure of water (filled circles) significantly increases with the  $\text{CH}_3\text{OH}$  wt%, but the transition pressure of ice VI to VII (filled squares) is weakly dependent on the  $\text{CH}_3\text{OH}$  wt%. During the decompression, the ice VII transforms back to the ice VI (open triangles) at approximately the same pressure as being compressed; the ice VI, however, experiences a hysteresis and melts to water around 0.9 GPa (open diamonds), independent of the  $\text{CH}_3\text{OH}$  wt%.

To understand the effect of pressure on the molecular bonding and phase transition behaviors in  $\text{H}_2\text{O}$ - $\text{CH}_3\text{OH}$  mixtures, we further compare the C-H stretching frequencies of  $\text{CH}_3\text{OH}$  in pure methanol with those in  $\text{H}_2\text{O}$  mixed with low (2.5, 5, 10 and 20 wt%) methanol content, as shown in Fig. 8. At ambient pressure, the C-H frequencies in all low methanol content mixtures are essentially the same and higher than the frequencies in pure methanol by  $\sim 9\text{ cm}^{-1}$  (Fig. 8a), in good agreement with previous studies<sup>26,28,40</sup>. Note that the C-O stretching frequencies (not shown) in all low methanol content mixtures ( $1017\text{ cm}^{-1}$ ) are the same and  $14\text{ cm}^{-1}$  lower than the frequency in pure methanol ( $1031\text{ cm}^{-1}$ )<sup>26,28,40,41</sup>. Upon compression, the C-H frequencies in low  $\text{CH}_3\text{OH}$  content mixtures increase with similarly constant slopes  $d\omega/dP$  larger than the constant slopes of pure methanol until the crystallization of water at pressures that vary with the  $\text{CH}_3\text{OH}$  wt%. The flattening of the C-H frequencies for 10 and 20 wt%  $\text{CH}_3\text{OH}$  persisted until 3.6 GPa, lower than the pressure for 5 wt% (4.6 GPa) and 2.5 wt%  $\text{CH}_3\text{OH}$

(5 GPa), after which all the frequencies increased with a constant slope similar to that of the pure methanol. Fig. 8b shows that the C-H frequencies reversibly shifted during decompression until approximately 1 GPa where the small slopes of all of the low methanol content mixtures started to increase. On the other hand, in the compression cycle, the decreasing rates of the O-H stretching frequencies of ice VII in all low methanol content mixtures were suppressed during a narrow pressure range, after which the frequencies red-shifted with nearly the same slope as that of pure  $\text{H}_2\text{O}$ . Such evolution for C-H and O-H frequencies under pressure along with the observed pressure-induced variations in Raman intensities of these modes within the sample, and the hysteresis transition where the decompressed ice VI melted to water around 1 GPa as in the pure  $\text{H}_2\text{O}$  (Fig. 7) suggest that the low content  $\text{CH}_3\text{OH}$  molecules undergo a pressure-induced segregation from the  $\text{H}_2\text{O}$  cage.

## Discussion

The C-H stretching frequencies of  $\text{CH}_3\text{OH}$  are sensitive to local hydrogen-bonding configurations. In pure methanol, the  $\text{CH}_3\text{OH}$  molecules form a series of chains and rings<sup>20,28,42</sup>, in which a  $\text{CH}_3\text{OH}$  molecule accepts a hydrogen bond from a neighboring molecule and also donates a hydrogen to form another hydrogen bond. The C-H stretching frequencies increase when the  $\text{CH}_3\text{OH}$  molecule serves as a proton acceptor (A) and the frequencies decrease when the  $\text{CH}_3\text{OH}$  is a proton donor (D)<sup>27</sup>. On the other hand, in the  $\text{H}_2\text{O}$  mixtures with low  $\text{CH}_3\text{OH}$  content, a  $\text{CH}_3\text{OH}$  molecule which is surrounded by an  $\text{H}_2\text{O}$  cage accepts two hydrogen bonds from neighboring  $\text{H}_2\text{O}$  molecules and donates a hydrogen atom to form a  $\text{CH}_3\text{OH}$ - $\text{H}_2\text{O}$  hydrogen bond. Therefore, the transformation from the acceptor-donor (AD) configuration in pure  $\text{CH}_3\text{OH}$  to the AAD configuration in low  $\text{CH}_3\text{OH}$  content mixtures results in higher C-H frequencies in low  $\text{CH}_3\text{OH}$  content mixtures at ambient pressure and a larger slope ( $d\omega/dP$ ) under compression. When water crystallizes to ice VI, the  $\text{H}_2\text{O}$  molecules form a tetragonal lattice arrangement that changes the local hydrogen-bonding configuration around the  $\text{CH}_3\text{OH}$  molecule and suppresses the increase in the C-H stretching frequencies under pressure. Upon further compression, the structural transition from ice VI (tetragonal) to VII (cubic) does not significantly change the bonding around methanol molecules and thus the evolution of C-H frequen-



**Figure 8** | Raman shifts of the C-H stretching modes of  $\text{CH}_3\text{OH}$  as a function of pressure and  $\text{CH}_3\text{OH}$  wt% during (a) compression and (b) decompression cycles. The 100 wt%, pure  $\text{CH}_3\text{OH}$  shows an approximately constant slope  $d\omega/dP$  in both compression and decompression cycles. The frequencies of both C-H modes for 2.5, 5, 10, and 20 wt%  $\text{CH}_3\text{OH}$  are higher than those for pure  $\text{CH}_3\text{OH}$  at ambient pressure, and initially increase linearly with a similar slope larger than those of pure  $\text{CH}_3\text{OH}$  until the crystallization of water, in which the slopes all become very small. The frequencies then increase again at higher pressures with a slope similar to those of the pure  $\text{CH}_3\text{OH}$ . The pressure induced frequency shifts in pure  $\text{CH}_3\text{OH}$  are reversible over the pressure range studied, whereas the frequency shifts in all low  $\text{CH}_3\text{OH}$  content mixtures are only reversible until around 1–2 GPa, where a hysteresis water-ice VI transition occurs and the small slopes of the C-H stretching modes of all low  $\text{CH}_3\text{OH}$  content mixtures increase around 1 GPa.



cies with pressure remains nearly the same. The O-H stretching modes of ice VII soften with pressure and experience a reduction in the slopes during a narrow pressure range, which may be due to the pressure-driven perturbation of CH<sub>3</sub>OH molecules that move and segregate from the ice VII cages. This hypothesis is supported by the observation that in lower CH<sub>3</sub>OH content mixtures the reduction in the softening rates occurs at higher pressure regime. When large numbers of CH<sub>3</sub>OH molecules aggregates, the local hydrogen bonding configuration becomes similar to that in the pure CH<sub>3</sub>OH with an AD configuration. As a result, the C-H stretching frequencies increase and the O-H frequencies decrease with slopes similar to those of the pure CH<sub>3</sub>OH and ice VII, respectively.

## Conclusion

To summarize, we have used Raman spectroscopy to investigate the effect of pressure on the inter- and intra-molecular bonding in H<sub>2</sub>O-CH<sub>3</sub>OH mixtures. The presence of small amounts of CH<sub>3</sub>OH stabilizes the liquid water phase under pressure, but does not significantly affect the ice VI–VII phase transition, reducing the pressure range ice VI exists. Pressure evolution of the C-H stretching frequencies of CH<sub>3</sub>OH and the O-H stretching frequencies of H<sub>2</sub>O combined with the pressure-induced variations in the Raman intensities of these modes reveals changes in the hydrogen-bonding configuration in H<sub>2</sub>O-CH<sub>3</sub>OH mixtures and suggests a pressure-induced segregation of low content CH<sub>3</sub>OH from the ice VII. These findings could be general in other low temperature crystalline phases of ice as well as in other H<sub>2</sub>O-volatile mixtures, and are of importance for Earth and planetary sciences. For instance, within large icy satellites and potential icy super-Earths H<sub>2</sub>O may be in low temperature crystalline phases, e.g., ice II, VI, and VIII, and the segregation of volatiles from H<sub>2</sub>O ices will substantially reduce the thermal conductivity and sound velocity of H<sub>2</sub>O-volatile mixtures that would hinder the heat transfer through planetary interiors. The suppressed heat transfer would affect the thermal evolution and geodynamics processes, and prevent the sub-surface ocean from crystallizing within large icy bodies. In addition, the hysteresis of the water-ice VI transition may result in inhomogeneous mixing of crystalline ices and liquid water-volatile mixture near the boundary between icy layers and sub-surface ocean when the geothermal gradient within large icy bodies fluctuates due to the competition between the heat source from the core and from tidal dissipation within icy layers, and the heat sink through convection in the outer ice I layer. It is worth noting that such hysteresis cycle implies that re-melting occurs at higher temperatures than the initial crystallization temperature and thus would require a substantial amount of heat to occur<sup>10</sup>. We point out, however, that since a dramatic change in the slope  $dw/dP$  of the Raman frequency could be a signature of pressure-induced phase transition, we cannot exclude the possibility of additional, novel phase transitions in the H<sub>2</sub>O-CH<sub>3</sub>OH mixtures. Future X-ray and neutron diffraction experiments and computer simulations are necessary to provide complementary insights into the bonding behavior and phase relationships in the H<sub>2</sub>O-CH<sub>3</sub>OH mixtures under extreme pressures.

## Methods

Distilled H<sub>2</sub>O mixed with specific wt% of CH<sub>3</sub>OH was loaded, together with a ruby ball, into a symmetric diamond-anvil cell (DAC) with a culet size of 500 μm. The mixture itself served as a pressure transmitting medium. The pressure was determined by ruby fluorescence<sup>43</sup> and the uncertainties of the pressure measurements were typically ~0.1 GPa. The compression and decompression rates of the samples were typically ~0.1 GPa s<sup>-1</sup>. After the pressure within the DAC equilibrated and reached a stable value, the high pressure Raman spectra of H<sub>2</sub>O-CH<sub>3</sub>OH mixtures were measured at room temperature using a Raman microscope (Horiba Jobin Yvon) that employs a CW 514.5 nm Ar-ion laser to excite the mixture within the DAC and probes the Raman scattering and vibrational frequency shifts with a spectral resolution of ~2 cm<sup>-1</sup>.

1. Nisini, B. Water's role in making stars. *Science* **290**, 1513–1514 (2000).

2. Cai, Y. Q. *et al.* Ordering of hydrogen bonds in high-pressure low-temperature H<sub>2</sub>O. *Phys. Rev. Lett.* **94**, 025502 (2005).
3. Aoki, K., Yamawaki, H. & Sakashita, M. Pressure-tuned Fermi resonance in ice VII. *Science* **268**, 1322–1324 (1995).
4. Goncharov, A., Struzhkin, V., Somayazulu, M., Hemley, R. & Mao, H. Compression of ice to 210 Gigapascals: Infrared evidence for a symmetric hydrogen-bonded phase. *Science* **273**, 218–220 (1996).
5. Chou, L., Blank, J., Goncharov, A., Mao, H. & Hemley, R. In situ observations of a high-pressure phase of H<sub>2</sub>O ice. *Science* **281**, 809–812 (1998).
6. Pruzan, P., Chervin, J. C. & Gauthier, M. Raman spectroscopy investigation of ice VII and deuterated ice VII to 40 GPa. Disorder in ice VII. *Eur. Lett.* **13**, 81–87 (1990).
7. Walrafen, G. E. *et al.* Raman and x-ray investigations of ice VII to 36.0 GPa. *J. Chem. Phys.* **77**, 2166–2174 (1982).
8. Fu, R., O'Connell, R. J. & Sasselov, D. D. The interior dynamics of water planets. *Astrophys. J.* **708**, 1326–1334 (2010).
9. Freeman, J. Non-Newtonian stagnant lid convection and the thermal evolution of Ganymede and Callisto. *Planet. Space Sci.* **54**, 2–14 (2006).
10. Deschamps, F., Mousis, O., Sanchez-Valle, C. & Lunine, J. I. The role of methanol in the crystallization of Titan's primordial ocean. *Astrophys. J.* **724**, 887–894 (2010).
11. Fortes, A. D. & Choukroun, M. Phase behaviour of ices and hydrates. *Space Sci. Rev.* **153**, 185–218 (2010).
12. Shin, K. *et al.* Methanol incorporation in clathrate hydrates and the implications for oil and gas pipeline flow assurance and icy planetary bodies. *Proc. Natl. Acad. Sci. U. S. A.* **110**, 8437–8442 (2013).
13. Fidler, J. & Rodger, P. M. Solvation structure around aqueous alcohols. *J. Phys. Chem. B* **103**, 7695–7703 (1999).
14. Laaksonen, A., Kusalik, P. G. & Svishchev, I. M. Three-dimensional structure in water - methanol mixtures. *J. Phys. Chem. A* **101**, 5910–5918 (1997).
15. Allison, S., Fox, J., Hargreaves, R. & Bates, S. Clustering and microimmiscibility in alcohol-water mixtures: Evidence from molecular-dynamics simulations. *Phys. Rev. B* **71**, 024201 (2005).
16. Meng, E. C. & Kollman, P. A. Molecular dynamics studies of the properties of water around simple organic solutes. *J. Phys. Chem.* **100**, 11460–11470 (1996).
17. Moin, S. T., Hofer, T. S., Randolf, B. R. & Rode, B. M. Structure and dynamics of methanol in water: a quantum mechanical charge field molecular dynamics study. *J. Comput. Chem.* **32**, 886–892 (2011).
18. Pascal, T. A. & Goddard, W. A. Hydrophobic segregation, phase transitions and the anomalous thermodynamics of water/methanol mixtures. *J. Phys. Chem. B* **116**, 13905–13912 (2012).
19. Da Silva, J. A. B., Moreira, F. G. B., dos Santos, V. M. L. & Longo, R. L. On the hydrogen bond networks in the water-methanol mixtures: topology, percolation and small-world. *Phys. Chem. Chem. Phys.* **13**, 6452–6461 (2011).
20. Guo, J.-H. *et al.* Molecular structure of alcohol-water mixtures. *Phys. Rev. Lett.* **91**, 157401 (2003).
21. Dixit, S., Crain, J., Poon, W. C. K., Finney, J. L. & Soper, A. K. Molecular segregation observed in a concentrated alcohol-water solution. *Nature* **416**, 829–832 (2002).
22. Soper, A. & Finney, J. L. Hydration of methanol in aqueous solution. *Phys. Rev. Lett.* **71**, 4346 (1993).
23. Guo, J.-H. *et al.* The molecular structure of alcohol-water mixtures determined by soft-X-ray absorption and emission spectroscopy. *J. Electron Spectros. Relat. Phenomena* **137–140**, 425–428 (2004).
24. Dougan, L. *et al.* Segregation in aqueous methanol enhanced by cooling and compression. *J. Chem. Phys.* **122**, 174514 (2005).
25. Dougan, L. *et al.* Methanol-water solutions: a bi-percolating liquid mixture. *J. Chem. Phys.* **121**, 6456–6462 (2004).
26. Ebukuro, T., Takami, A., Oshima, Y. & Koda, S. Raman spectroscopic studies on hydrogen bonding in methanol and methanol/water mixtures under high temperature and pressure. *J. Supercrit. Fluids* **15**, 73–78 (1999).
27. Gruenloh, C. J., Florio, G. M., Carney, J. R., Hagemeyer, F. C. & Zwier, T. S. C-H stretch modes as a probe of H-bonding in methanol-containing clusters. *J. Phys. Chem. A* **103**, 496–502 (1999).
28. Dixit, S., Poon, W. & Crain, J. Hydration of methanol in aqueous solutions: a Raman spectroscopic study. *J. Phys. Condens. Matter* **12**, L323–L328 (2000).
29. Del Corro, E., Cáceres, M., Taravillo, M., Núñez, J. & Baonza, V. G. Raman spectroscopy of aqueous methanol solutions under pressure. *High Press. Res.* **26**, 407–410 (2006).
30. Kühne, T. D. & Khaliullin, R. Z. Electronic signature of the instantaneous asymmetry in the first coordination shell of liquid water. *Nat. Commun.* **4**, 1450; DOI:10.1038/ncomms2459 (2013).
31. Mao, W. L. & Mao, H.-K. Hydrogen storage in molecular compounds. *Proc. Natl. Acad. Sci. U. S. A.* **101**, 708–710 (2004).
32. Lin, Y., Mao, W. L. & Mao, H.-K. Storage of molecular hydrogen in an ammonia borane compound at high pressure. *Proc. Natl. Acad. Sci. U. S. A.* **106**, 8113–8116 (2009).
33. Somayazulu, M., Finger, L. W., Hemley, R. J. & Mao, H. K. High-pressure compounds in methane-hydrogen mixtures. *Science* **271**, 1400–1402 (1996).
34. Vos, W. L., Finger, L. W., Hemley, R. & Mao, H. Novel H<sub>2</sub>-H<sub>2</sub>O clathrates at high pressures. *Phys. Rev. Lett.* **71**, 3150–3153 (1993).



35. Carey, D. M. & Korenowski, G. M. Measurement of the Raman spectrum of liquid water. *J. Chem. Phys.* **108**, 2669–2675 (1998).
36. Devendorf, G. S., Hu, M. A. & Ben-amotz, D. Pressure dependent vibrational Fermi resonance in liquid CH<sub>3</sub>OH and CH<sub>2</sub>Cl<sub>2</sub>. *J. Phys. Chem. A* **102**, 10614–10619 (1998).
37. Mammone, J. F., Sharma, S. K. & Nicol, M. Raman spectra of methanol and ethanol at pressures up to 100 kbar. *J. Phys. Chem* **84**, 3130–3134 (1980).
38. Arencibia, A., Taravillo, M., Cáceres, M., Núñez, J. & Baonza, V. G. Pressure tuning of the Fermi resonance in liquid methanol: implications for the analysis of high-pressure vibrational spectroscopy experiments. *J. Chem. Phys.* **123**, 214502 (2005).
39. Bridgman, P. W. Viscosities to 30,000 kg/cm<sup>2</sup>. *Proc. Am. Acad. Arts Sci.* **77**, 117–128 (1949).
40. Kabisch, G. & Pollmer, K. Hydrogen bonding in methanol-organic solvent and methanol-water mixtures as studied by the νCO and νOH Raman bands. *J. Mol. Struct.* **31**, 35–50 (1982).
41. Zerda, T. W., Thomas, H. D., Bradley, M. & Jonas, J. High pressure isotropic bandwidths and frequency shifts of the C–H and C–O modes of liquid methanol. *J. Chem. Phys.* **86**, 3219–3224 (1987).
42. Yamaguchi, T., Hidaka, K. & Soper, A. K. The structure of liquid methanol revisited: a neutron diffraction experiment at –80°C and +25°C. *Mol. Phys.* **96**, 1159–1168 (1999).
43. Mao, H. K., Bell, P. M., Shaner, J. W. & Steinberg, D. J. Specific volume measurements of Cu, Mo, Pd, and Ag and calibration of the ruby R1 fluorescence pressure gauge from 0.06 to 1 Mbar. *J. Appl. Phys.* **49**, 3276–3283 (1978).

## Acknowledgments

This work was supported by the Academia Sinica and the Ministry of Science and Technology of Taiwan, Republic of China, under Contract NSC 103-2112-M-001-001-MY3. The authors would like to thank Prof. Wendy Mao of Stanford University and Dr. Frederic Deschamps of Academia Sinica for their helpful comments on the manuscript.

## Author contributions

W.P.H. conceived and designed research. W.P.H. and Y.H.C. conducted the experiments and analyzed the data. W.P.H. wrote the manuscript. Both authors reviewed the manuscript.

## Additional information

**Competing financial interests:** The authors declare no competing financial interests.

**How to cite this article:** Hsieh, W.-P. & Chien, Y.-H. High pressure Raman spectroscopy of H<sub>2</sub>O–CH<sub>3</sub>OH mixtures. *Sci. Rep.* **5**, 8532; DOI:10.1038/srep08532 (2015).



This work is licensed under a Creative Commons Attribution 4.0 International License. The images or other third party material in this article are included in the article's Creative Commons license, unless indicated otherwise in the credit line; if the material is not included under the Creative Commons license, users will need to obtain permission from the license holder in order to reproduce the material. To view a copy of this license, visit <http://creativecommons.org/licenses/by/4.0/>