

Research Article

Riemannian Means on Special Euclidean Group and Unipotent Matrices Group

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Among the noncompact matrix Lie groups, the special Euclidean group and the unipotent matrix group play important roles in both theoretic and applied studies. The Riemannian means of a finite set of the given points on the two matrix groups are investigated, respectively. Based on the left invariant metric on the matrix Lie groups, the geodesic between any two points is gotten. And the sum of the geodesic distances is taken as the cost function, whose minimizer is the Riemannian mean. Moreover, a Riemannian gradient algorithm for computing the Riemannian mean on the special Euclidean group and an iterative formula for that on the unipotent matrix group are proposed, respectively. Finally, several numerical simulations in the 3-dimensional case are given to illustrate our results.

1. Introduction

A matrix Lie group, which is also a differentiable manifold simultaneously, attracts more and more researchers' attention from both theoretic interest and its applications [1–5]. The Riemannian mean on the matrix Lie groups is widely studied for its varied applications in biomedicine, signal processing, and robotics control [6–9]. Fiori and Tanaka [10] suggested a general-purpose algorithm to compute the average element of a finite set of matrices belonging to any matrix Lie group. In [11], the author investigated the Riemannian mean on the compact Lie groups and proposed a globally convergent Riemannian gradient descent algorithm. Different invariant notions of mean and average rotations on $SO(3)$ (it is compact) are given in [9]. Recently, Fiori [12] dealt with computing averages over the group of real symplectic matrices, which found applications in diverse areas such as optics and particle physics.

However, the Riemannian mean on the special Euclidean group $SE(n)$ and the unipotent matrix group $UP(n)$, which are the noncompact matrix Lie groups, has not been well studied. Fletcher et al. [6] proposed an iterative algorithm to obtain the approximate solution of the Riemannian mean on $SE(3)$ by use of the Baker-Cambell-Hausdorff formula. In [7], the

exponential mapping from the arithmetic mean of points on the Lie algebra $\mathfrak{se}(3)$ to the Lie group $SE(3)$ was constructed to give the Riemannian mean in order to get a mean filter.

In this paper, the Riemannian means on $SE(n)$ and those on $UP(n)$, which are both important noncompact matrix Lie groups [13, 14], are considered, respectively. Especially, $SE(3)$ is the spacial rigid body motion, and $UP(3)$ is the 3-dimensional Heisenberg group $H(3)$. Based on the left invariant metric on the matrix Lie groups, we get the geodesic distance between any two points and take their sum as a cost function. And the Riemannian mean will minimize it. Furthermore, the Riemannian mean on $SE(n)$ is gotten using the Riemannian gradient algorithm, rather than the approximate mean. An iterative formula for computing the Riemannian mean on $UP(n)$ is proposed according to the Jacobi field. Finally, we give some numerical simulations on $SE(3)$ and those on $H(3)$ to illustrate our results.

2. Overview of Matrix Lie Groups

In this section, we briefly introduce the Riemannian framework of the matrix Lie groups [15, 16], which forms the foundation of our study of the Riemannian mean on them.

2.1. *The Riemannian Structures of Matrix Lie Groups.* A group G is called a Lie group if it has differentiable structure: the group operators, that is, $G \times G \rightarrow G, (x, y) \mapsto x \cdot y$ and $G \rightarrow G, x \mapsto x^{-1}$, are differentiable, $x, y \in G$. A matrix Lie group is a Lie group with all elements matrices. The tangent space of G at identity is the Lie algebra \mathfrak{g} , where the Lie bracket is defined.

The exponential map, denoted by \exp , is a map from the Lie algebra \mathfrak{g} to the group G . Generally, the exponential map is neither surjective nor injective. Nevertheless, it is a diffeomorphism between a neighborhood of the identity I on G and a neighborhood of the identity $\mathbf{0}$ on \mathfrak{g} . The (local) inverse of the exponential map is the logarithmic map, denoted by \log .

The most general matrix Lie group is the general linear group $GL(n, \mathbb{R})$ consisting of the invertible $n \times n$ matrices with real entries. As the inverse image of $\mathbb{R} - \{0\}$ under the continuous map $A \mapsto \det(A)$, $GL(n, \mathbb{R})$ is an open subset of the set of $n \times n$ real matrices, denoted by $M_{n \times n}$, which is isomorphic to $\mathbb{R}^{n \times n}$, it has a differentiable manifold structure (submanifold). The group multiplication of $GL(n, \mathbb{R})$ is the usual matrix multiplication, the inverse map takes a matrix A on $GL(n, \mathbb{R})$ to its inverse A^{-1} , and the identity element is the identity matrix I . The Lie algebra $\mathfrak{gl}(n, \mathbb{R})$ of $GL(n, \mathbb{R})$ turns out to be $M_{n \times n}$ with the Lie bracket defined by the matrix commutator

$$[X, Y] = XY - YX, \quad \forall X, Y \in \mathfrak{gl}(n, \mathbb{R}). \quad (1)$$

All other real matrix Lie groups are subgroups of $GL(n, \mathbb{R})$, and their group operators are subgroup restrictions of the ones on $GL(n, \mathbb{R})$. The Lie bracket on their Lie algebras is still the matrix commutator.

Let S denote a matrix Lie group and \mathfrak{s} its Lie algebra. The exponential map for S turns out to be just the matrix exponential; that is, given an element $X \in \mathfrak{s}$, the exponential map is

$$\exp(X) = \sum_{m=0}^{\infty} \frac{X^m}{m!}. \quad (2)$$

The inverse map, that is, the logarithmic map, is defined as follows:

$$\log(A) = \sum_{m=1}^{\infty} (-1)^{m+1} \frac{(A - I)^m}{m}, \quad (3)$$

for A in a neighborhood of the identity I of S . The exponential of a matrix plays a crucial role in the theory of the Lie groups, which can be used to obtain the Lie algebra of a matrix Lie group, and it transfers information from the Lie algebra to the Lie group.

The matrix Lie group also has the structure of a Riemannian manifold. For any $A, B \in S$ and $X \in T_A S$, the tangent space of S at A , we have the maps that

$$\begin{aligned} L_A B &= AB, & (L_A)_* X &= AX, \\ R_A B &= BA^{-1}, & (R_{A^{-1}})_* X &= XA, \end{aligned} \quad (4)$$

where L denotes the left translation, R denotes the right translation, and $(L_A)_*$ and $(R_{A^{-1}})_*$ are the tangent mappings associated with L_A and $R_{A^{-1}}$, respectively. The adjoint action $Ad_A : \mathfrak{s} \rightarrow \mathfrak{s}$ is

$$Ad_A X = AXA^{-1}. \quad (5)$$

It is also easy to see the formula that

$$Ad_A = L_A R_A. \quad (6)$$

Then, the left invariant metric on S is given by

$$\begin{aligned} \langle X, Y \rangle_A &= \langle (L_{A^{-1}})_* X, (L_{A^{-1}})_* Y \rangle_I \\ &= \langle A^{-1} X, A^{-1} Y \rangle_I := \text{tr} \left((A^{-1} X)^T A^{-1} Y \right) \end{aligned} \quad (7)$$

with $X, Y \in T_A S$ and tr denoting the trace of the matrix. Similarly, we can define the right invariant metric on S as well. It has been shown that there exist the left invariant metrics on all matrix Lie groups.

2.2. *Compact Matrix Lie Group.* A Lie group is compact if its differential structure is compact. The unitary group $U(n)$, the special unitary group $SU(n)$, the orthogonal group $O(n)$, the special orthogonal group $SO(n)$, and the symplectic group $Sp(n)$ are the examples of the compact matrix Lie groups [17]. Denote a compact Lie group by S_1 and its Lie algebra by \mathfrak{s}_1 . There exists an adjoint invariant metric $\langle \cdot, \cdot \rangle$ on S_1 such that

$$\langle Ad_A X, Ad_A Y \rangle = \langle X, Y \rangle \quad (8)$$

with $X, Y \in \mathfrak{s}_1$. Notice the fact that the left invariant metric of any adjoint invariant metric is also right invariant; namely, it is a bi-invariant metric; so all compact Lie groups have bi-invariant metrics. Furthermore, if the left invariant and the adjoint invariant metrics on S_1 deduce a Riemannian connection ∇ , then the following properties are valid:

$$\begin{aligned} \nabla_X Y &= \frac{1}{2} [X, Y], \\ \langle \mathcal{R}(X, Y) X, Y \rangle &= -\frac{1}{4} \langle [X, Y], [X, Y] \rangle, \end{aligned} \quad (9)$$

where $\mathcal{R}(X, Y)$ is a curvature operator about the smooth tangent vector field on the Riemannian manifold (S_1, ∇) . Therefore, the section curvature \mathcal{K} is given by

$$\mathcal{K}(X, Y) = \frac{\langle [X, Y], [X, Y] \rangle}{4(\langle X, X \rangle \langle Y, Y \rangle - \langle X, Y \rangle^2)} \geq 0, \quad (10)$$

which means that \mathcal{K} is nonnegative on the compact Lie group.

In addition, according to the Hopf-Rinow theorem, a compact connected Lie group is geodesically complete. It means that, for any given two points, there exists a geodesic curve connecting them and the geodesic curve can extend infinitely.

2.3. *The Riemannian Mean on Matrix Lie Group.* Let $\gamma : [0, 1] \rightarrow S$ be a sufficiently smooth curve on S . We define the length of $\gamma(t)$ by

$$\begin{aligned} \ell(\gamma) &:= \int_0^1 \sqrt{\langle \dot{\gamma}(t), \dot{\gamma}(t) \rangle_{\gamma(t)}} dt \\ &= \int_0^1 \sqrt{\text{tr} \left\{ \left(\gamma(t)^{-1} \dot{\gamma}(t) \right)^T \gamma(t)^{-1} \dot{\gamma}(t) \right\}} dt, \end{aligned} \tag{11}$$

where T denotes the transpose of the matrix. The geodesic distance between two matrices A and B on S considered as a differentiable manifold is the infimum of the lengths of the curves connecting them; that is,

$$\begin{aligned} d(A, B) &:= \inf \{ \ell(\gamma) \mid \gamma : [0, 1] \\ &\rightarrow S \text{ with } \gamma(0) = A, \gamma(1) = B \}. \end{aligned} \tag{12}$$

According to the Euclidean analogue (mean on Euclidean space), a definition of the mean of N matrices R_1, \dots, R_N is the minimizer of the sum of the squared distances from any matrix to the given matrices R_1, \dots, R_N on S . Now, we define the Riemannian mean based on the geodesic distance (12).

Definition 1. The mean of N given matrices R_1, \dots, R_N on S in the Riemannian sense corresponding to the metric (7) is defined as

$$\bar{R} = \arg \min_{R \in S} \frac{1}{2N} \sum_{k=1}^N d(R_k, R)^2. \tag{13}$$

3. The Riemannian Mean on SE(n)

In this section, we discuss the Riemannian mean on the special Euclidean group $SE(n)$, which is a subgroup of $GL(n+1, \mathbb{R})$. Moreover, the special rigid body motion group $SE(3)$ is taken as an illustrating example.

3.1. *About SE(n).* The special Euclidean group $SE(n)$ in \mathbb{R}^n is the semidirect product of the special orthogonal group $SO(n)$ with \mathbb{R}^n itself [18]; that is,

$$SE(n) = SO(n) \ltimes \mathbb{R}^n. \tag{14}$$

The matrix representation of elements of $SE(n)$ is

$$SE(n) = \left\{ \begin{pmatrix} A & b \\ 0 & 1 \end{pmatrix} \mid A \in SO(n), b \in \mathbb{R}^n \right\}. \tag{15}$$

An element of $SE(n)$ physically represents a displacement, where A corresponds to the orientation, or attitude, of the rigid body and b encodes the translation. The Lie algebra $\mathfrak{se}(n)$ of $SE(n)$ can be denoted by

$$\mathfrak{se}(n) = \left\{ \begin{pmatrix} \Omega & v \\ 0 & 0 \end{pmatrix} \mid \Omega^T = -\Omega, v \in \mathbb{R}^n \right\}. \tag{16}$$

Specially, when $n = 3$, the skew-symmetric matrix Ω can be uniquely expressed as

$$\Omega = \begin{pmatrix} 0 & -\omega_z & \omega_y \\ \omega_z & 0 & -\omega_x \\ -\omega_y & \omega_x & 0 \end{pmatrix} \tag{17}$$

with $\omega = (\omega_x, \omega_y, \omega_z) \in \mathbb{R}^3$. $\|\omega\|_F$ gives the amount of rotation with respect to the unit vector along ω , where $\|\cdot\|_F$ denotes the Frobenius norm. Physically, ω represents the angular velocity of the rigid body, whereas v corresponds to the linear velocity [19]. In [18], the author presents a closed-form expression of the exponential map $\mathfrak{se}(3) \rightarrow SE(3)$ by

$$\exp(V) = I_4 + V + \frac{1 - \cos(\theta)}{\theta^2} V^2 + \frac{\theta - \sin(\theta)}{\theta^3} V^3 \tag{18}$$

with $V \in \mathfrak{se}(3)$ and $\theta^2 = \omega_x^2 + \omega_y^2 + \omega_z^2$. Note that it can be regarded as an extension of the well-known Rodrigues formula on $SO(3)$. The logarithmic map $SE(3) \rightarrow \mathfrak{se}(3)$ is yielded as

$$\log(Q) = q_1 (q_2 I_4 - q_3 Q + q_4 Q^2 - q_5 Q^3), \tag{19}$$

where

$$\begin{aligned} q_1 &= \frac{1}{8} \csc^3 \left(\frac{\theta}{2} \right) \sec \left(\frac{\theta}{2} \right), \\ q_2 &= \theta \cos(2\theta) - \sin(\theta), \\ q_3 &= \theta \cos(\theta) + 2\theta \cos(2\theta) - \sin(\theta) - \sin(2\theta), \\ q_4 &= 2\theta \cos(\theta) + \theta \cos(2\theta) - \sin(\theta) - \sin(2\theta), \\ q_5 &= \theta \cos(\theta) - \sin(\theta), \end{aligned} \tag{20}$$

$\text{tr}(Q) = 2 + 2 \cos(\theta)$, for $-\pi < \theta < \pi$.

3.2. *Algorithm for the Riemannian Mean on SE(n).* Denote $P, Q \in SE(n)$ by

$$P = \begin{pmatrix} A_1 & b_1 \\ 0 & 1 \end{pmatrix}, \quad Q = \begin{pmatrix} A_2 & b_2 \\ 0 & 1 \end{pmatrix}. \tag{21}$$

Taking the corresponding exponential mappings on manifolds $SO(n)$ and \mathbb{R}^n into consideration, the geodesic $\gamma_{P,Q}$ between P and Q on the Lie group $SE(n)$ is given by

$$\begin{aligned} \gamma_{P,Q}(t) &= \begin{pmatrix} \alpha_{A_1, A_2}(t) & \beta_{b_1, b_2}(t) \\ 0 & 1 \end{pmatrix} \\ &= \begin{pmatrix} A_1 (A_1^T A_2)^t & b_1 + (b_2 - b_1)t \\ 0 & 1 \end{pmatrix}, \end{aligned} \tag{22}$$

where $\alpha : [0, 1] \rightarrow SO(n)$ and $\beta : [0, 1] \rightarrow \mathbb{R}^n$ are the geodesics expressed, respectively, by

$$\begin{aligned} \alpha_{A,B}(t) &= \exp_A(t \log(A^T B)) = A(A^T B)^t, \quad A, B \in SO(n), \\ \beta_{b_1, b_2}(t) &= \exp_{b_1}(t(b_2 - b_1)) \\ &= b_1 + (b_2 - b_1)t, \quad b_1, b_2 \in \mathbb{R}^n. \end{aligned} \tag{23}$$

Then, the midpoint of P and Q is defined by

$$P \circ Q = \begin{pmatrix} A_1 (A_1^T A_2)^{1/2} & \frac{1}{2}(b_1 + b_2) \\ 0 & 1 \end{pmatrix}. \tag{24}$$

Before the geodesic distance on $SE(n)$ is given, we first introduce a lemma which is a known conclusion in linear algebra [20].

Lemma 2. *If $E \in \mathbb{R}^{n \times n}$ and $H \in \mathbb{R}^{m \times m}$ are invertible matrices, then the block matrix*

$$\begin{pmatrix} E & F \\ 0 & H \end{pmatrix} \quad (25)$$

is invertible, where $F \in \mathbb{R}^{n \times m}$. Furthermore,

$$\begin{pmatrix} E & F \\ 0 & H \end{pmatrix}^{-1} = \begin{pmatrix} E^{-1} & -E^{-1}FH^{-1} \\ 0 & H^{-1} \end{pmatrix}. \quad (26)$$

Now, we give the geodesic distance on $SE(n)$ as follows.

Lemma 3. *The geodesic distance between two points P and Q on $SE(n)$ induced by the scale-dependent left invariant metric (7) is given by*

$$d(P, Q) = \left(\|\log(A_1^T A_2)\|_F^2 + \|b_2 - b_1\|_F^2 \right)^{1/2}. \quad (27)$$

Proof. As mentioned above, the geodesic distance between two matrices P and Q on $SE(n)$ is achieved by the length of geodesics connecting them; thus, we will compute it through substituting (22) into (11).

From Lemma 2, we get

$$\begin{aligned} \gamma_{P,Q}^{-1}(t) &= \begin{pmatrix} (A_1^T A_2)^{-t} A_1^T & -(A_1^T A_2)^{-t} A_1^T (b_1 + (b_2 - b_1)t) \\ 0 & 1 \end{pmatrix}. \end{aligned} \quad (28)$$

Then, according to the principle about the derivatives of the matrix-valued functions, the following formula is valid:

$$\dot{\gamma}_{P,Q}(t) = \begin{pmatrix} A_1 (A_1^T A_2)^t \log(A_1^T A_2) & b_2 - b_1 \\ 0 & 0 \end{pmatrix}. \quad (29)$$

Moreover, we have that

$$\begin{aligned} \text{tr} \left(\left(\gamma_{P,Q}^{-1}(t) \dot{\gamma}_{P,Q}(t) \right)^T \gamma_{P,Q}^{-1}(t) \dot{\gamma}_{P,Q}(t) \right) &= -\log^2(A_1^T A_2) + (b_2 - b_1)^T (b_2 - b_1). \end{aligned} \quad (30)$$

Therefore, the geodesic distance on $SE(n)$ between P and Q is given by

$$\begin{aligned} d(P, Q) &= \int_0^1 \left(-\log^2(A_1^T A_2) + (b_2 - b_1)^2 \right)^{1/2} dt \\ &= \left(\|\log(A_1^T A_2)\|_F^2 + \|b_2 - b_1\|_F^2 \right)^{1/2}. \end{aligned} \quad (31)$$

This completes the proof of Lemma 2. □

In addition, it is valuable to mention that the distance $\|\log(A_1^T A_2)\|_F$, induced by the standard bi-invariant metric on $SO(n)$, stands for the rotation motion from the point P to Q and the distance $\|b_2 - b_1\|_F$ stands for the translation motion on \mathbb{R}^n . Therefore, considering an object undergoing a rigid body Euclidean motion, then, this motion can be decomposed into a rotation with respect to the center of mass of the object and a translation of the center of mass of the object.

Theorem 4. *For N given points on $SE(n)$*

$$P_k = \begin{pmatrix} A_k & b_k \\ 0 & 1 \end{pmatrix}, \quad (32)$$

where $A_k \in SO(n)$ and $b_k \in \mathbb{R}^n$, $k = 1, 2, \dots, N$, if the Riemannian mean of A_1, A_2, \dots, A_N and the Riemannian mean of b_1, b_2, \dots, b_N (i.e., arithmetic mean) are denoted by \bar{A} and \bar{b} , respectively, then, one has the Riemannian mean \bar{P} of $P_1, P_2, \dots, P_N \in SE(n)$ by

$$\bar{P} = \begin{pmatrix} \bar{A} & \bar{b} \\ 0 & 1 \end{pmatrix}. \quad (33)$$

Proof. In the Riemannian sense, by (13), the mean \bar{P} is defined as

$$\begin{aligned} \bar{P} &= \arg \min_{P \in SE(n)} \frac{1}{2N} \sum_{k=1}^N d(P_k, P)^2 \\ &= \arg \min_{P \in SE(n)} \frac{1}{2N} \sum_{k=1}^N \left(\|\log(A_k^T A)\|_F^2 + \|b_k - b\|_F^2 \right) \\ &= \arg \min_{A \in SO(n)} \frac{1}{2N} \sum_{k=1}^N \|\log(A_k^T A)\|_F^2 \\ &\quad + \arg \min_{b \in \mathbb{R}^n} \frac{1}{2N} \sum_{k=1}^N \|b_k - b\|_F^2. \end{aligned} \quad (34)$$

From [9], the geodesic distance between A_k and A on $SO(n)$ is given by

$$d(A_k, A)^2 = \|\log(A_k^T A)\|_F^2, \quad (35)$$

so we have that

$$\arg \min_{A \in SO(n)} \frac{1}{2N} \sum_{k=1}^N \|\log(A_k^T A)\|_F^2 = \arg \min_{A \in SO(n)} d(A_k, A)^2 = \bar{A}. \quad (36)$$

On the other hand, for $b_k \in \mathbb{R}^n$, $k = 1, 2, \dots, N$, it is easy to see that

$$\arg \min_{b \in \mathbb{R}^n} \frac{1}{2N} \sum_{k=1}^N \|b_k - b\|_F^2 = \frac{1}{N} \sum_{k=1}^N b_k = \bar{b}. \quad (37)$$

Therefore, the fact is shown that the Riemannian mean \bar{b} of $\{b_k\}$ is equivalent to the arithmetic mean.

Consequently, we prove that equality (33) is valid. □

In addition, let L denote the cost function of the minimization problem (34) on $SE(n)$; that is,

$$L(P) = L_{\text{rota}}(A) + L_{\text{trans}}(b) = \frac{1}{2N} \sum_{k=1}^N \|\log(A_k^T A)\|_F^2 + \frac{1}{2N} \sum_{k=1}^N \|b_k - b\|_F^2, \quad (38)$$

where L_{rota} and L_{trans} stand for the rotation and the translation components of the cost function L , respectively. We have the gradient of $L_{\text{rota}}(A)$ for $A \in SO(n)$ as follows [21]:

$$\text{grad}(L_{\text{rota}}) = -A \sum_{k=1}^N \log(A^T A_k). \quad (39)$$

Consequently, the Riemannian gradient descent algorithm is applied to calculate \bar{A} , taking the geodesic on $SO(n)$ as the trajectory and the negative gradient (39) as the descent direction.

Finally, we achieve the following algorithm for computing the Riemannian mean \bar{P} on $SE(n)$.

Algorithm 5. Given N matrices P_k , $k = 1, 2, \dots, N$, on $SE(n)$, their Riemannian mean \bar{P} is computed by the following iterative method.

- (1) Store $(1/N) \sum_{k=1}^N b_k$ to \bar{b} .
- (2) Set $\bar{A} = A_1$ as an initial input, and choose a desired tolerance $\varepsilon > 0$.
- (3) If $\|\sum_{k=1}^N \log(\bar{A}^T A_k)\|_F < \varepsilon$, then stop.
- (4) Otherwise, update $\bar{A} = \bar{A} \exp\{-\varepsilon \sum_{k=1}^N \log(\bar{A}^T A_k)\}$, and go to step (3).

3.3. Simulations on $SE(3)$. Let us consider a rigid object W in the Euclidean space undergoing a rigid body Euclidean motion $SE(3)$. Suppose that the coordinate of the center of gravity in W is $d_W \in \mathbb{R}^3$; then, the optimal trajectory from the configuration P to Q is the curve $D(t)$ such that

$$\begin{pmatrix} D(t) \\ 1 \end{pmatrix} = \gamma_{P,Q}(t) \begin{pmatrix} d_W \\ 1 \end{pmatrix}, \quad (40)$$

where $t \in [0, 1]$ and $\gamma_{P,Q}(t)$ denotes the geodesic connecting P and Q on $SE(3)$ (see Figure 1). For the configuration of two points P and Q , as shown in Figure 2, given by the angular velocity ω_P, ω_Q of the rigid body and the linear velocity v_P, v_Q , we choose $\omega_P = (\pi/2)(0, 1, 1)$, $v_P = (0, 0, 0)$, $\omega_Q = \pi(1/4, 0, -1/2)$, and $v_Q = (4.380, -1.348, 3.690)$; then, we obtain their Riemannian mean according to Algorithm 5, which is just the middle point $P \circ Q$ from (24).

4. The Riemannian Mean on $UP(n)$

In this section, the Riemannian mean of N given points on the unipotent matrix group $UP(n)$ is considered. $UP(n)$ is a noncompact matrix Lie group as well. Moreover, in the special case $n = 3$, it is the Heisenberg group $H(3)$.

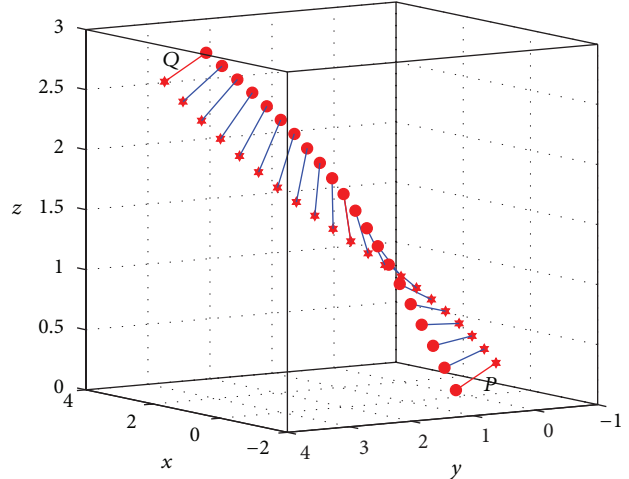


FIGURE 1: The rigid motion $D(t)$ from P to Q .

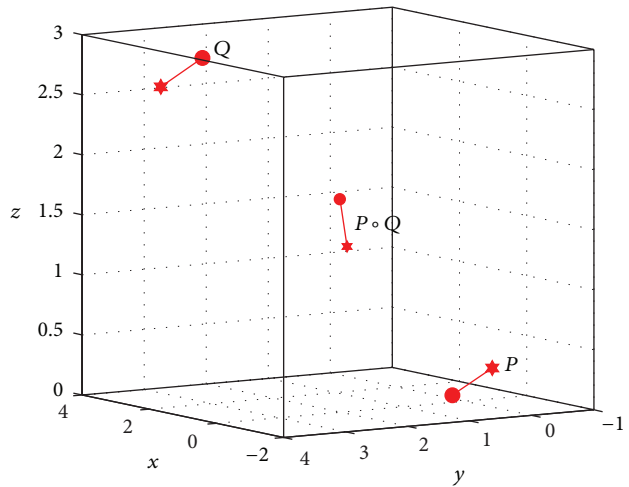


FIGURE 2: The Riemannian mean $P \circ Q$.

4.1. About $UP(n)$. The set of all of the uppertriangular $n \times n$ matrices with diagonal elements that are all one is called unipotent matrices group, denoted by $UP(n)$.

In fact, given an invertible matrix $C \in UP(n)$, there is a neighborhood U of C such that every matrix in U is also in $UP(n)$, so $UP(n)$ is an open subset of $\mathbb{R}^{n \times n}$. Furthermore, the matrix product $P \cdot Q$ is clearly a smooth function of the entries of P and Q , and P^{-1} is a smooth function of the entries of P . Thus, $UP(n)$ is a Lie group. On the other hand, it can be verified that $UP(n)$ is of dimension $n(n-1)/2$ and is nilpotent. Since we can use the nonzero elements C_{ij} , $i < j$, directly as global coordinate functions for $UP(n)$, the manifold underlying $UP(n)$ is diffeomorphic to $\mathbb{R}^{n(n-1)/2}$. Therefore, $UP(n)$ is not compact, but simply connected.

The Lie algebra $\mathfrak{up}(n)$ of $UP(n)$ consists of uppertriangular matrices T with diagonal elements $T_{ii} = 0$, $i = 1, \dots, n$. It is an indispensable tool which gives a realization of the Heisenberg commutation relations of quantum mechanics in the 3-dimensional case [17].

Moreover, it is the fact that both $C - I$ and T are all nilpotent matrices, for any $C \in \text{UP}(n)$ and $T \in \mathfrak{up}(n)$. Thus, from (2) and (3), the infinite series representations of the exponential mapping in $\mathfrak{up}(n)$ and the logarithm mapping in $\text{UP}(n)$ can be given, respectively, by

$$\log(C) = \sum_{m=1}^n (-1)^{m+1} \frac{(C - I)^m}{m}, \tag{41}$$

where $C \in \text{UP}(n)$, $\|C - I\|_F < 1$, and

$$\exp(T) = \sum_{m=0}^n \frac{T^m}{m!} \tag{42}$$

with $T \in \mathfrak{up}(n)$.

Notice that $\text{UP}(n)$ is connected, which means that, for any given pair A, B , we can find a geodesic curve $\gamma(t)$ such that $\gamma(0) = A$ and $\gamma(1) = B$, namely, by taking the initial velocity as $\dot{\gamma}(0) = \log(A^{-1}B)$. Let the geodesic curve $\gamma(t)$ be

$$\gamma(t) = A \exp(t \log(A^{-1}B)) \in \text{UP}(n) \tag{43}$$

with $\gamma(0) = A$, $\gamma(1) = B$, and $\dot{\gamma}'(0) = \log(A^{-1}B) \in \mathfrak{up}(n)$. Then, the midpoint of A and B is given by

$$A \circ B = A \exp\left(\frac{1}{2} \log(A^{-1}B)\right), \tag{44}$$

and from (11) the geodesic distance $d(A, B)$ can be computed explicitly by

$$d(A, B) = \|\log(A^{-1}B)\|_F. \tag{45}$$

4.2. Algorithm for the Riemannian Mean on $\text{UP}(n)$. For N given points B^1, B^2, \dots, B^N in $\text{UP}(n)$, L denotes the cost function of the minimization problem (13); that is,

$$\min_{A \in \text{UP}(n)} L(A) = \min_{A \in \text{UP}(n)} \frac{1}{2N} \sum_{k=1}^N d(B^k, A)^2. \tag{46}$$

Following [22, 23], it has been shown that the Jacobi field is equal to zero at the Riemannian mean. The Jacobi field for the Riemannian mean is equal to the sum of tangent vectors to all geodesics (from mean to each point). Noticing the fact that the geodesic between two points A and B has already been given by (43), we can then compute the Jacobi field at point A to N points B^k (at $t = 0$) such that

$$\begin{aligned} \gamma_k(t) &= A(A^{-1}B^k)^t = A \exp(t \log(A^{-1}B^k)), \\ \left. \frac{d\gamma_k(t)}{dt} \right|_{t=0} &= A \log(A^{-1}B^k), \quad k = 1, \dots, N. \end{aligned} \tag{47}$$

Then, we suppose that the summation of all these vectors should be equal to zero; that is,

$$L_A = \sum_{k=1}^N \left. \frac{d\gamma_k(t)}{dt} \right|_{t=0} = A \sum_{k=1}^N \log(A^{-1}B^k) = 0, \tag{48}$$

so the Riemannian mean A of the N matrices $\{B^k\}$ should satisfy

$$\sum_{k=1}^N \log(A^{-1}B^k) = 0. \tag{49}$$

From the logarithm of the matrices on $\text{UP}(n)$ given by (41), we can rewrite (49) as

$$\sum_{k=1}^N \sum_{m=1}^n (-1)^{m+1} \frac{(A^{-1}B^k - I)^m}{m} = 0. \tag{50}$$

Therefore, the Riemannian mean A of the N given matrices $\{B^k\}$ can be given explicitly by solving (50).

For the case of $n = 2$, from (50), it is shown that the Riemannian mean \bar{A}_2 of N given matrices $\{B_2^k\}$ in $\text{UP}(2)$ is their arithmetic mean; that is,

$$\bar{A}_2 = \frac{1}{N} \sum_{k=1}^N B_2^k. \tag{51}$$

Next, for $n = 3$, we obtain the Riemannian mean on $\text{UP}(3)$ ($H(3)$) as follows.

Theorem 6. Given N matrices $\{B_3^k\}$ on the Heisenberg group $H(3)$ by

$$B_3^k = \begin{pmatrix} 1 & b_{12}^k & b_{13}^k \\ 0 & 1 & b_{23}^k \\ 0 & 0 & 1 \end{pmatrix}, \tag{52}$$

where $k = 1, 2, \dots, N$, then, one has the Riemannian mean \bar{A}_3 on the Heisenberg group $H(3)$ such that

$$\bar{A}_3 = \begin{pmatrix} 1 & \bar{b}_{12} & \bar{b}_{13} - \frac{1}{2} \text{cov}(b_{12}, b_{23}) \\ 0 & 1 & \bar{b}_{23} \\ 0 & 0 & 1 \end{pmatrix}, \tag{53}$$

where $\bar{b}_{ij} := (1/N) \sum_{k=1}^N b_{ij}^k$, $i, j = 1, 2, 3$ ($i < j$), and $\text{cov}(b_{12}, b_{23}) := (1/N) \sum_{k=1}^N (b_{12}^k - \bar{b}_{12})(\bar{b}_{23} - b_{23}^k)$.

Proof. First, let us denote the Riemannian mean \bar{A}_3 by

$$\bar{A}_3 = \begin{pmatrix} 1 & a_{12} & a_{13} \\ 0 & 1 & a_{23} \\ 0 & 0 & 1 \end{pmatrix}. \tag{54}$$

Then, note that, for the given matrices $\{B_3^k\}$ on $H(3)$, their Riemannian mean \bar{A}_3 has to satisfy (50), so we get the following solutions:

$$\begin{aligned} a_{12} &= \frac{1}{N} \sum_{k=1}^N b_{12}^k, \\ a_{23} &= \frac{1}{N} \sum_{k=1}^N b_{23}^k, \\ a_{13} &= \frac{1}{N} \sum_{k=1}^N (a_{12} - b_{12}^k)(a_{23} - b_{23}^k). \end{aligned} \tag{55}$$

As a matter of convenience, supposing that $\bar{b}_{ij} := (1/N) \sum_{k=1}^N b_{ij}^k$, $i, j = 1, 2, 3$ ($i < j$), and $\text{cov}(b_{12}, b_{23}) := (1/N) \sum_{k=1}^N (\bar{b}_{12} - b_{12}^k)(\bar{b}_{23} - b_{23}^k)$, we show that (54) is valid.

This completes the proof of Theorem 6. □

More generally, while $n > 1$, we can get the Riemannian mean on $UP(n)$ given by the following theorem.

Theorem 7. Take $n > 1$. For N given matrices $\{B_n^k\}$ in $UP(n)$, one assumes that they are in the form of

$$B_n^k = \begin{pmatrix} B_{n-1}^k & b_{n-1}^k \\ 0 & 1 \end{pmatrix} \tag{56}$$

with $B_{n-1}^k \in UP(n-1)$ and $b_{n-1}^k \in \mathbb{R}^{n-1}$; then, the Riemannian mean \bar{A}_n of the N matrices B_n^k is given by

$$\bar{A}_n = \begin{pmatrix} \bar{A}_{n-1} & a_{n-1} \\ 0 & 1 \end{pmatrix}, \tag{57}$$

where \bar{A}_{n-1} is the Riemannian mean of $\{B_{n-1}^k\}$ and a_{n-1} is given by the formula that

$$\begin{aligned} a_{n-1} &= \bar{A}_{n-1} \left(\sum_{k=1}^N \sum_{m=0}^{n-1} \frac{(-1)^m}{m+1} (\bar{A}_{n-1}^{-1} B_{n-1}^k - I)^m \right)^{-1} \\ &\times \left(\sum_{k=1}^N \sum_{m=0}^{n-1} \frac{(-1)^m}{m+1} (\bar{A}_{n-1}^{-1} B_{n-1}^k - I)^m \bar{A}_{n-1}^{-1} b_{n-1}^k \right). \end{aligned} \tag{58}$$

Proof. For simplicity of exposition, we suppose that the Riemannian mean \bar{A}_n is the block matrix in the form of

$$\bar{A}_n = \begin{pmatrix} A_{n-1} & a_{n-1} \\ 0 & 1 \end{pmatrix} \tag{59}$$

with $A_{n-1} \in UP(n-1)$ and $a_{n-1} \in \mathbb{R}^{n-1}$. Since the Riemannian mean \bar{A}_n of the N matrices $\{B_n^k\}$ should satisfy (50), we substitute the block matrix forms (59) and (57) into (50). Then, we obtain the following matrix equation for the Riemannian mean \bar{A}_n :

$$\begin{aligned} &\sum_{k=1}^N \sum_{m=1}^n \frac{(-1)^{m-1}}{m} \\ &\times \left((A_{n-1}^{-1} B_{n-1}^k - I)^m (A_{n-1}^{-1} B_{n-1}^k - I)^{m-1} A_{n-1}^{-1} (b_{n-1}^k - a_{n-1}) \right) = 0, \end{aligned} \tag{60}$$

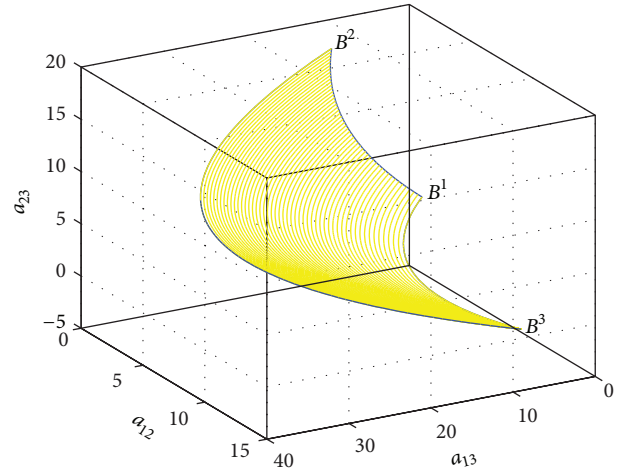


FIGURE 3: The geodesic triangle on $H(3)$.

which means that (58) is valid and A_{n-1} satisfies the equation

$$\sum_{k=1}^N \sum_{m=1}^n \frac{(-1)^{m-1}}{m} (A_{n-1}^{-1} B_{n-1}^k - I)^m = 0. \tag{61}$$

Moreover, from (41), we have that

$$\log(A_{n-1}^{-1} B_{n-1}^k) = 0. \tag{62}$$

Furthermore, it is shown that A_{n-1} is the Riemannian mean of $\{B_{n-1}^k\}$. At last, we write A_{n-1} as \bar{A}_{n-1} , so the proof of Theorem 7 is completed. □

As shown above, we give the iterative formula for computing the Riemannian mean for any dimension $n > 1$. Either (51) or (54) can be chosen as the initial formula.

4.3. Simulations on $H(3)$. In this section, we take two examples to illustrate the results about the Riemannian mean on the Heisenberg group $H(3)$, which is the 3-dimensional space.

Example 8. Consider the Riemannian mean of three points B^1, B^2, B^3 on the Heisenberg group $H(3)$. Using (43), we can get the geodesics of three points on $H(3)$, which form a geodesic triangle. In Figure 3, all of the curves are geodesics. Moreover, as shown in Figure 4, the midpoint of each geodesic is easy to be obtained by (44). Thus, each centerline connects a vertex to the midpoint of its opposing side. On $H(3)$, these centerlines always meet in a single point which is coincident with the Riemannian mean computed by (54), denoted by a red dot as shown in Figure 4.

Example 9. Given four points B^1, B^2, B^3, B^4 on the Heisenberg group $H(3)$, we can get a geodesic tetrahedron from (43) (see Figure 5), where all curves are geodesics. Moreover, similar to Example 8, the Riemannian means of three vertexes on each curved face are obtained, denoted by red circles (see Figure 6). Then, we plot each centerline which connects a vertex to the Riemannian mean of its opposing side. It

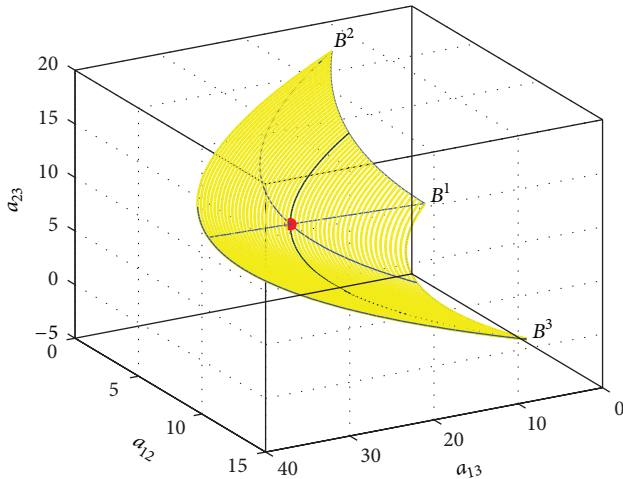


FIGURE 4: The Riemannian mean of three points.

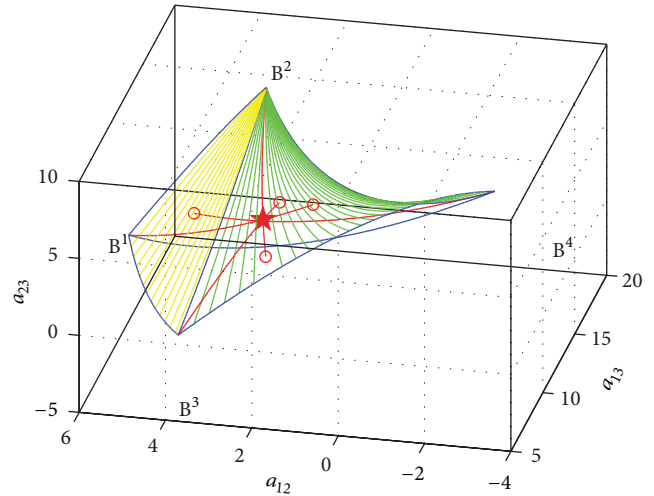


FIGURE 6: The Riemannian mean of four points.

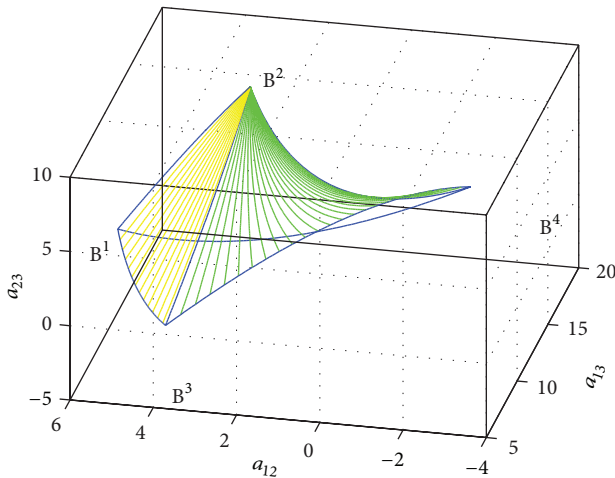


FIGURE 5: The geodesic tetrahedron on $H(3)$.

is shown that these centerlines still meet in a single point, denoted by a red pentacle. In fact, the point is the Riemannian mean of B^1, B^2, B^3, B^4 applying (54).

5. Conclusion

In this paper, we consider the Riemannian means on the special Euclidean group $SE(n)$ and the unipotent matrix group $UP(n)$, respectively. Based on the left invariant metric on the matrix Lie groups, we get the geodesic distance between any two points and take their sum as a cost function. Furthermore, we get the Riemannian mean on $SE(n)$ using the Riemannian gradient algorithm. Moreover, we give an iterative formula for computing the Riemannian mean on $UP(n)$ according to its Jacobi field. Finally, we make advantages of several numerical simulations on $SE(3)$ and $H(3)$ to illustrate our results.

Conflict of Interests

The authors declare that there is no conflict of interests regarding the publication of this paper.

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