# Ultra-low background measurements of decayed aerosol filters

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**Abstract** Aerosol samples collected on filter media were analyzed using HPGe detectors employing varying background-reduction techniques in order to experimentally evaluate the opportunity to apply ultra-low background measurement methods to samples collected, for instance, by the Comprehensive Test Ban Treaty International Monitoring System (IMS). In this way, realistic estimates of the impact of low-background methodology on the

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M. Laubenstein Laboratori Nazionali del Gran Sasso, S.S.17/bis km 18+910, 67010 Assergi, AQ, Italy sensitivity obtained in systems such as the IMS were assessed. The current detectability requirement of stations in the IMS is  $30 \mu Bq/m^3$  of air for <sup>140</sup>Ba, which would imply  $\sim 10^6$  fissions per daily sample. Importantly, this is for a fresh aerosol filter. One week of decay reduces the intrinsic background from radon daughters in the sample allowing much higher sensitivity measurement of relevant isotopes, including <sup>131</sup>I. An experiment was conducted in

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## Introduction

The Comprehensive Test Ban Treaty (CTBT), when in force, will prohibit nuclear explosions. The Verification Regime of this treaty exists to detect nuclear explosions. Seismic, hydroacoustic, and infrasound waveform technologies 'listen' for vibrations in the Earth's rock, oceans, and atmosphere, respectively. These technologies rapidly obtain information about the size and location of a suspect event. Radionuclide technologies, aerosol and xenon monitoring, exist to capture some of the radioactive atoms emitted by the explosion and by doing so confirm the nuclear nature of the event and possibly screen out unrelated civilian or natural phenomena, such as mining blasts and earthquakes. Aerosol samples of particular relevance to treaty compliance, i.e., samples that show evidence of two or more fission or activation products, are sent to specially certified IMS laboratories for confirmatory measurements at higher sensitivity levels. All these technologies comprise a network of 321 stations around the Earth: the International Monitoring System (IMS) which continuously monitors the environment and transmits data to a central data facility, the International Data Center (IDC), in Vienna, Austria.

The xenon IMS component is particularly valuable for detecting leakage from underground tests, as noble gases are the most likely to escape the containment of an underground test and be detectable via radioactive decay some distance downwind. Xenon was observed as a leakage from a recent underground test [1], showing the value of this method for verification, although xenon leakage from US tests at the Nevada Test Site has been documented previously in DOE/NV-317 [2]. The leakage of radioisotopes from historic US underground testing is more instructive, however, as many more isotopes were observed to leak, and leakage fractions were documented from  $10^{-3}$ to somewhat below  $10^{-7}$ . Iodine isotopes were the most commonly observed besides xenon. This is not unexpected, as iodine initially exists as a highly volatile gas. In the atmosphere a significant fraction of it reacts rapidly to form aerosol and organic species in addition to gas.

The aerosol monitoring IMS component was originally designed to detect atmospheric tests though the sensitivity of the aerosol equipment,  $30 \mu Bq/m^3$  for <sup>140</sup>Ba, is orders of

magnitude below that observed for historical atmospheric tests. While the system performance of the aerosol network is computed to verify that the network meets a notional design criterion of 90% probability of detection of a 1 kt atmospheric test within 2 weeks using <sup>140</sup>Ba, the consideration of the capability of the aerosol network to detect underground tests is an important but so far unexplored area. The detection of the 2006 event in the DPRK by just one IMS xenon system suggests that improving the chances of a marginal detection in aerosol could greatly increase the confidence of detecting an underground explosion. This is especially important since the xenon network is *less* dense than the aerosol network, 40 stations compared to 80.

## Source strength

To assess the capability of the aerosol network for underground tests, a source strength estimate is needed, together with average atmospheric dilution that would occur between source and station, and the sensitivity of the stations to iodine isotopes. Leakage from US NTS tests can be estimated by using stated yield and leaked activity from DOE/NV-317, assuming  $1.4 \times 10^{23}$  fissions/kt and a nominal decay time for the activity released. From this result, we take the leakage fraction to be  $10^{-7}$ , almost a worst case. That is, one in every 10 million fission atoms escapes containment. This is a very crude analysis, and does not consider many factors such as the impacts of differing rock type, yield level, and containment sophistication. For a hypothetical 1 kt explosion, this means that  $10^{16}$  fission atoms escape.

### Atmospheric dilution scenario

The dilution expected from a surface release at a distance downwind similar to the spacing of the IMS aerosol stations would be about  $10^{15}$ . For the hypothetical case above, the release of  $10^{16}$  atoms released, there would be 10 atoms/m<sup>3</sup> at the monitoring station. The IMS aerosol systems have a minimum sample size of 12,000 m<sup>3</sup>, such that, for a plume that covered the monitoring location for a full 24 h,  $1.2 \times 10^5$  fission debris atoms could be collected, leading to about 3600 <sup>131</sup>I atoms, and 7200 atoms each of <sup>140</sup>Ba and <sup>99</sup>Mo. The significance of only a part (perhaps as low as 1/3) of the iodine being in particulate form and the possibility that the aerosol might be too large to be supported by the atmosphere will be reintroduced later.

#### System sensitivity

The sensitivity (in fissions atoms) for an IMS aerosol system and for several alternative detection systems [3] has

Isotope	<sup>140</sup> Ba Minimum detectab	$^{140}La$	<sup>99</sup> Mo	<sup>131</sup> I		
Detector	Minimum detectable fission atoms at KASA $I_0$					
RASA (normal IMS)	641,535	315,505	67,421	268,713		
P-type, shallow	29,079	14,745	9,548	14,725		
Well-type, shallow	11,453	4,922	691	3,697		
TWIN: p-type, deep	10,764	4,756	1,553	3,627		
RNL: multi p-type, deep	3,472	1,831	643	1,539		

 Table 1
 Estimated sensitivity for IMS measurement and for alternative detection approaches in which a 7 day decay and a 7 day measurement are made

Minimum detectable fission atoms are the minimum detectable atoms of a species divided by the cumulative fission yield, which is  $\sim 3\%$  for <sup>131</sup>I, and  $\sim 6\%$  for the others

been estimated using a simple calculation based on the Currie formulation [4] of the critical limit (Lc) with a 95% confidence level. An IMS aerosol system would detect several isotopes clearly when  $10^6$  fission atoms are collected, or about 100 atoms/m<sup>3</sup>. By comparison, Table 1 compares estimated sensitivity values for a normal IMS station to a variety of detector types (p-type, well type, multi-crystal) in various low background configurations (a 30 m water equivalent (mwe) underground site and a site located deeper than 1000 mwe). The RNL multi-crystal system is described by another paper in this volume (Keillor et al.), and for the purposes of this table is operated as a single large detector, i.e., not in coincidence mode.

# **Experimental approach**

To test the estimates made previously, an experiment was designed to send IMS filters to a series of underground laboratories. A collection and measurement station in Vienna, Austria, compliant to IMS requirements was used; i.e.,  $>12,000 \text{ m}^3$  of air collected in 24 h. Each of these

filters was handled and measured in the normal IMS procedure, in which a large volume of air is drawn for 24 h, a subsequent 24 h decay period reduces Rn decay products, then an HPGe assay is performed. Afterwards, each filter was compressed to a disk 5 mm thick by 50 mm diameter. The result was a cylindrical geometry that could be placed directly on the endcap of the germanium detector, the closest possible geometry. This leads to summing effects for isotopes with multiple simultaneous gamma emissions, but is nevertheless the most sensitive geometry available for this type of sample and detector where additional reduction of the sample volume by ashing is disallowed.

The samples were rapidly shipped to several laboratories employing background reduction techniques of several kinds: careful shielding, special clean materials for shielding and detector construction, muon veto systems, and location below a burden of shielding material, ranging over three orders of magnitude. The participating laboratories are listed in Table 2.

The goal was for the compressed samples to be measured for about 1 week after 1 week of decay. Some samples were received and measured about a week late, but

Abbreviation	Meters water equivalent	Location	Institution
IRSN/LMRE (O)	7.5	Orsay, France	IRSN/LMRE
MPI-K	15	Heidelberg, Germany	Max-Planck-Institut für Kernphysik
IAEA-MEL	35	Monaco, Monaco	Marine Environment Laboratories
VKTA	110	Dresden, Germany	VKTA Rossendorf
UI	300	Como, Italy	Universita' dell'Insubria
IRMM	500	Geel, Belgium	EC-JRC-IRMM
IFIN-HH	560	Magurele, Romania	National Institute of RD for Physics and Nuclear Engr
РТВ	1100	Braunschweig, Germany	PTB
LNGS	3800	Gran Sasso, Italy	Laboratori Nazionali del Gran Sasso
IRSN/LMRE (M)	4800	Modane, France	IRSN/LMRE

**Table 2** Participatinglaboratories, depth in meterswater equivalent, and location

measurements done at one underground lab (IRMM) covering 3.5 weeks showed that a week of excess decay on the filter was insignificant for the purposes of this work, since the resultant spectrum did not change appreciably and the back-correction for decay could be artificially adjusted to eliminate the unwanted decay. This would not be possible in real operations, of course, as the desired Treaty isotopes would have undergone sensitivity-reducing decay, even though the backgrounds do not reduce appreciably after the first week.

# **Results and discussion**

Given that the filters varied somewhat in the mix of Rn progeny and Be activity, no comparison of this type would fairly reflect the capability of the laboratories involved. The results are intended to show generally what might be accomplished on average, but with only one sample per laboratory, variations in the samples due to local weather in Vienna skew the results to some degree. For instance, one laboratory (IAEA-MEL) received a filter with 1/5 the <sup>7</sup>Be activity of the average of the others.

One conclusion is that while low background systems could improve the detection sensitivity to  $^{131}$ I (364 keV, 82% probability) the Compton continuum below the 477 keV <sup>7</sup>Be line creates a limiting factor for ultra-low background detectors for all isotopes detected in this energy range, seen in spectra published previously [3]. Above 477 keV, the main limitations of the ultra-low background approach are the efficiency of the detector, the time available for measurement. In fact, the <sup>131</sup>I sensitivity via 637 keV line emitted with 7.2% probability is likely better than the 364 keV line in all underground detectors.

The preliminary results from the laboratory measurement are shown in Table 3. Because of the masking effect of <sup>7</sup>Be Compton scatters, the <sup>99</sup>Mo and <sup>140</sup>Ba-La are calculated to be of similar or greater nuclear explosion detection sensitivity than <sup>131</sup>I. This comparison is misleading, because of the much higher probability of iodine leakage. Table 4 shows a summary of the frequency of various isotopes being detected in the leakage from US tests at the Nevada Test Site, as compiled from DOE/NV-317. Iodine, specifically <sup>131</sup>I, was reported three times as often as <sup>140</sup>Ba, and yet <sup>140</sup>Ba may have only been observed in gross venting scenarios. Of the 356 individual isotope detections, 219 were iodine isotopes. Of the 80 identified leaking tests, 70 were listed as having measurable iodine. The preference for iodine leakage may be much stronger even than these statistics indicate, and the more complete the containment, the stronger the preference may be.

 
 Table 3
 Laboratory results in minimum detectable fission atoms for an IMS filter with a 7 day decay followed by 7 days of measurement

Lab	mwe	<sup>99</sup> Mo 739.5 keV	<sup>140</sup> Ba 537.3 keV	<sup>140</sup> La 487 keV	<sup>131</sup> I 364.5 keV
IRSN/ LMRE (O)	7.5	5796	4699	4399	5159
MPI-K	15	4570	3297	11465	9199
IAEA-MEL	35	1971	1231	1350	1232
VKTA	110	-	-	-	-
UI	300	4583	3796	3795	6716
IRMM	500	1827	1548	1725	5014
IFIN-HH	560	23491	17379	14371	19419
РТВ	1100	1643	1599	2016	6500
LNGS	3800	2930	2969	2859	10916
IRSN/ LMRE (M)	4800	1506	1348	1465	5354

Some variations in natural atmospheric radioactivity were observed in the filters. Some extra decay occurred in some samples, but this was shown to be insignificant for this experiment, as described in the text

 $\label{eq:table_transform} \begin{array}{l} \textbf{Table 4} \ Frequency \ of \ release \ of \ top \ isotopes \ as \ compiled \ from \ DOE/NV-317 \end{array}$ 

	Isotope	Frequency	T 1/2	
1	<sup>133</sup> I	63	20.3	h
2	<sup>131</sup> I	62	8.04	d
3	<sup>135</sup> I	62	6.61	h
4	<sup>138</sup> Cs	27	32.2	m
5	<sup>140</sup> Ba	19	12.8	d
6	<sup>88</sup> Rb	15	17.8	m
7	<sup>103</sup> Ru	14	39.4	d
8	<sup>132</sup> Te	13	77.9	h
9	$^{132}I$	11	143	m
10	<sup>91</sup> Sr	8	9.7	h
11	$^{134}$ I	8	52.6	m
12	<sup>95</sup> Zr	6	64.4	d
13	<sup>106</sup> Ru	5	368	d
14	<sup>139</sup> Ba	5	84.9	m
15	<sup>141</sup> Ce	5	32.4	d
16	<sup>137</sup> Cs	4	30.1	У

The comparison of Tables 3 to 1 is interesting. It might be thought that the deepest lab would return the most sensitive results, but that may be impossible for a relatively short measurement (7 days) and where natural radioactivity (<sup>7</sup>Be, <sup>40</sup>K, etc.) remains after the decay period. In fact, the shallowest laboratory had excellent detection capability for every isotope, far better than the Table 1 estimates for the shallow p-type detector, even approaching the deep underground detector (TWIN) in sensitivity. In fact, the <sup>131</sup>I signal could endure a factor of ten reduction for fall-out losses near the vent location and a factor of three in speciation loss and still be easily detectable by the shallowest laboratory.

This presents a conclusion that in the scenario discussed above, the use of ultra-low background technology could make <sup>131</sup>I detectable by the IMS. Detector size, depth, peak-to-Compton ratio, cosmic veto systems and Compton suppression systems may affect the relative performance of the participating laboratories. These and other factors must be considered to create a tailored, cost effective approach that could be adapted to each of the 16 CTBT radionuclide laboratories. In addition, the laboratory results show that the original calculation approach was conservative. This may suggest that the computed station sensitivity enhancements could be similarly improved.

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