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The continuous and discrete molecular orbital x-ray bands from Xe^{q+} ($12 \leq q \leq 29$) + Zn collisions

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In this paper, the x-ray emissions are measured by the interaction of 1500–3500 keV Xe^{q+} ($q = 12, 15, 17, 19, 21, 23, 26$ and 29) ions with Zn target. When $q < 29$, we observe $L\text{I}$, $L\alpha$, $L\beta_1$, $L\beta_2$ and $L\gamma$ characteristic x-rays from Xe^{q+} ions and a broad M-shell molecular orbital (MO) x-ray band from the transient quasi-molecular levels. It is found that their yields quickly increase with different rates as the incident energy increases. Besides, the widths of the broad MO x-ray bands are about 0.9–1.32 keV over the energy range studied and are proportional to $v^{1/2}$ ($v =$ projectile velocity). Most remarkably, when the projectile charge state is 29, the broad x-ray band separates into several narrow discrete spectra, which was never observed before in this field.

Research on charge-exchange processes during ion-atom collisions has attracted considerable interest for recent decades^{1–5}. For slow highly charged ion interactions with the target, quasi-molecular states that are formed during collisions have proved to be successful in qualitative describing electron-vacancy exchange processes^{6–12}. In this frame, as the collision partners approach each other, the atomic orbitals evolve into the molecular orbitals and some atomic levels may turn into two or more molecular levels^{6,7}.

When a vacancy participates in such an interaction, an electron in a low-lying level may be promoted to the vacant orbital through so-called electron promotion^{1,6,7}. For example, most K-shell vacancy productions are attributed to the promotion of $2p\sigma \rightarrow 2p\pi$ at slow collisions. $2p\sigma$ and $2p\pi$ molecular orbitals are from the low-Z partner 1s and high-Z partner 2p atomic orbitals, respectively. This mechanism has been used to interpret a wealth of experimental results, such as the origin of characteristic lines, the energy level matching effect, the charge state effect and the atomic number effect, among others^{1,9,10,12–14}. Alternatively, the vacancy brought into the quasi-molecular system also is possible to be filled by an electron occupying a high-lying level, accompanied by molecular orbital (MO) x-ray emission, i.e., MO transition. Saris *et al.* first reported this phenomenon by measuring x-ray emissions from Ar ions of energy 70–600 keV impinging on C, Al, Si and Fe targets¹¹. In their experiment, they observed a broad MO x-ray band, which was attributed to the radiative filling of a vacancy in the $2p\pi$ level of the Ar-Ar system during the collision. To date, several experimental investigations of MO x-rays have been reported^{15–18}.

As we know, the formation of the quasi-molecule may allow us to understand the properties of the super-heavy atoms in advance and x-ray emission can provide some deeper insights on the dynamical process of its formation. This is why a lot of experiments were devoted to x-ray emissions in slow collisions. Unfortunately, the molecular orbitals are difficult to calculate, especially for higher shells and multi-electrons systems^{19–21}. Therefore, the precise theoretical calculations in electron promotion and MO transition processes are still lacking for slow highly charged ion-atom collisions. More experiments are necessary for its understanding.

Although, under certain conditions, both the processes of electron promotion and MO transition can occur in the same ion-atom collision and there were many experiments focusing on the collision between slow highly charged ion and atom, these two processes have never been observed and studied simultaneously in experiment up to now.

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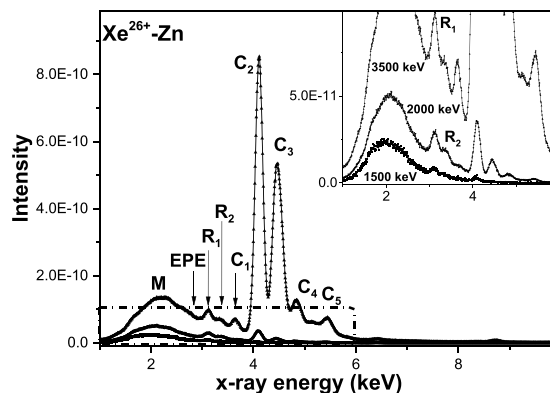


Figure 1. Typical x-ray spectra from collisions of $\text{Xe}^{q < 29+}$ ions with a Zn target at 1500 keV, 2000 keV and 3500 keV (from bottom to top). The insert graph shows a zoomed-in view of the x-ray spectral structure in the dotted box. M: MO x-ray band; EPE: end point energy; R_1 and R_2 : radiative Auger effect peaks; C_1 , C_2 , C_3 , C_4 , and C_5 : $L\alpha$, $L\beta_1$, $L\beta_2$ and $L\gamma$ x-ray peaks, respectively, of Xe^{q+} ions.

In order to check whether these two processes can happen in the same slow ion-atom collision and understand how they happen, in this paper, the collisions between 1500–3500 keV Xe^{q+} ions ($q = 12\text{--}29$) with Zn target are studied by measuring the x-rays produced from it.

Experiment method

The experiment was performed using the 320 kV high voltage platform of an all permanent Electron Cyclotron Resonance Ion Source (ECRIS) at the Institute of Modern Physics (IMP), Chinese Academy of Sciences (CAS). It was designed to be operated at 14.5 GHz with the purpose of producing medium charge state and high charge state gaseous and metallic ion beams²². Xe^{q+} ions of charge states 12, 15, 17, 19, 21, 23, 26 and 29 in the energy range of 1500 to 3500 keV were provided by the ECRIS. The essentials of the experimental apparatus were described in detail elsewhere⁸. The Xe^{q+} ion beam, extracted and selected by a 90° analyzing magnet, was first highly collimated by two sets of four jaw slits and used to bombard the surface of a Zn target at 45° . The target used in our experiment, with a purity of 99.99% and a polished surface, had a thickness of $50\ \mu\text{m}$ and a size of $20 \times 21\ \text{mm}^2$. The emitted x-rays were counted by a Si(Li) detector placed at 90° relative to the incident direction of the beams. The Si(Li) detector had an energy resolution of 165 eV at 5.9 keV and a detection range of 1–60 keV. The energy calibration of the detector was made by using radioactive sources ^{241}Am and ^{55}Fe . The pressure of the target chamber was maintained at 2×10^{-8} mbar. In order to reduce the statistical error, each measurement was taken for 3000 seconds.

Results and Discussion

Typical x-ray spectra obtained with $\text{Xe}^{q < 29+}$ ions incident upon the target at different incoming energies are shown in Fig. 1. All spectra have been normalized by incident particle number. From top to bottom, the spectra were induced by 3500, 2000 and 1500 keV Xe^{q+} ions, respectively. To clearly show the low energy excited spectral structure, a zoomed-in view of the spectra in the dotted box is presented in the inset of Fig. 1. Obviously, in Fig. 1, every measured spectral structure can be divided into two components, namely, a non-characteristic x-ray band M and the characteristic x-ray lines C_1 , C_2 , C_3 , C_4 , and C_5 , by their full widths at half-maximum (FWHM). FWHMs of the x-ray bands are approximately 0.90–1.32 keV, and their end point energy of 2.7 keV agrees well with the energy of the M-shell x-rays in the united Po atom limit²³. According to the Fano-Lichten description of heavy-ion collisions, we know that radiative decay of the vacancies can occur during the MO formation process, and the emitted x-ray spectra have a broad FWHM and end point energy (EPE) that matches with the united atom (UA) characteristic x-ray energy. Hence, in our results, the pronounced non characteristic M bands are expected as M-shell MO x-rays of united Po atoms. Meanwhile, the peaks, marked as C_1 , C_2 , C_3 , C_4 , and C_5 , which respectively peak at 3.6, 4.0, 4.4, 4.8 and 5.4 keV, are identified as $L\alpha$, $L\beta_1$, $L\beta_2$ and $L\gamma$ characteristic x-ray of the Xe^{q+} ions²⁴. From the above figure, as the energy of the Xe^{q+} ions increasing, a large increase is simultaneously observed in the yields of the characteristic and non-characteristic x-rays. Moreover, the FWHM of the MO x-ray spectrum is also increasing with increasing energy. In addition, plots of the M band FWHM versus $v^{1/2}$ for different charge states are shown in Fig. 2. From this figure, one can see that the FWHM of the band M is proportional to the square root of the projectile velocity. This finding is consistent with that predicted by Betz *et al.*^{25,26}. But in our experiment, the slope is different from that given by Betz *et al.* because they focused on the K-shell measurement.

To understand the origin of the M band and the Xe^{q+} L-shell characteristic lines, an approximate correlation diagram representing the Xe-Zn molecule based on MO theory is presented in Fig. 3⁷.

To be specific, in this theory, as the inter-nuclear distance shrinks, the target 2p orbital evolves to $2p\pi$, $3d\pi$ and $4f\pi$ MOs, of which the unoccupied $3d\pi$ is filled by the $4f\pi$ to generate MO x-rays (denoted by arrow (A) in Fig. 3). When the vacancies decay at the closet distance, the high energy limit, i.e., the tail of the MO line, occurs; this energy is equal to that of the M x-rays of the united Xe-Zn atom (Po, $Z = 84$), in our results. After collision, a broad molecular x-ray spectrum is formed. Meanwhile, during the interaction, the electron promotion process

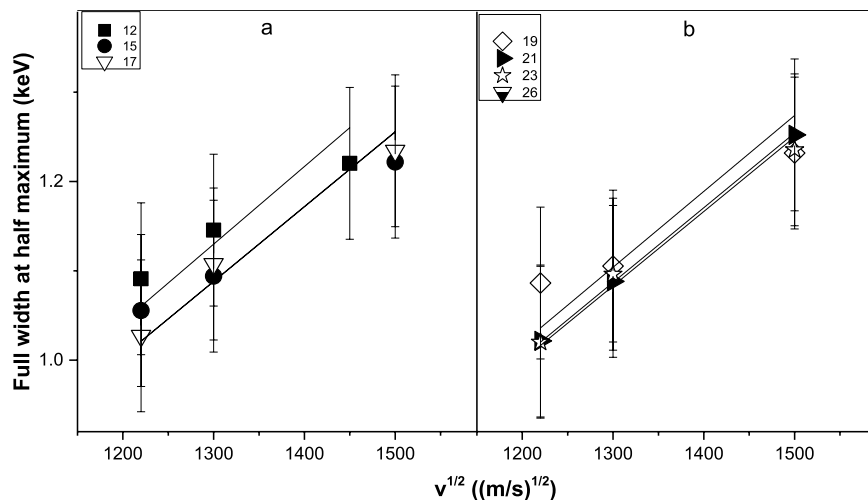


Figure 2. Full width at half maximum (FWHM) of the M band for $\text{Xe}^{12, 15, 17+}$ (a) and $\text{Xe}^{19, 21, 23, 26+}$ (b) vs $v^{1/2}$. The solid lines represent the fitting results obtained using $\text{FWHM} = kv^{1/2}$.

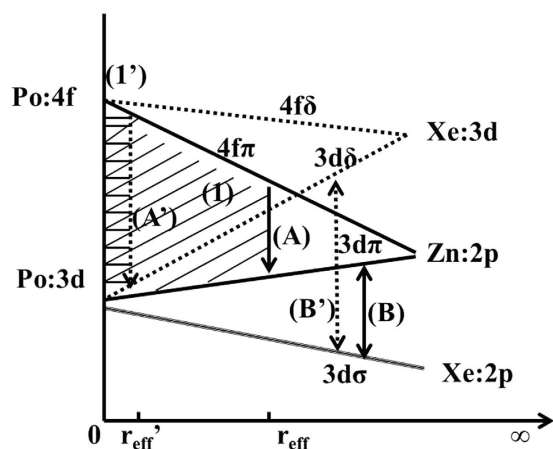


Figure 3. Molecular correlation diagram (a qualitative estimate of the $\text{Xe} + \text{Zn}$ molecular levels). (1) and (1'): MO transition region; (A) and (B): MO transition and electron promotion for a $\text{Xe}^{q < 29+} + \text{Zn}$ atom collision; (A') and (B'): MO transition and electron promotion for a $\text{Xe}^{29+} + \text{Zn}$ atom collision; r_{eff} (r'_{eff}): effective range of MO transition induced by $\text{Xe}^{q < 29+}$ (Xe^{29+}).

(denoted by arrow (B)) is also expected. The 2p electron of Xe^{q+} is promoted by $3d\sigma \rightarrow 3d\pi$ rotational coupling^{1,6,7}. After collision, the Xe^{q+} ions carry 2p vacancies, whose radiative decay leads to the emission of Xe^{q+} L-shell x-rays.

Obviously, for the $4f\pi \rightarrow 3d\pi$ and $3d\sigma \rightarrow 3d\pi$ transitions to occur, the target atom must carry 2p vacancies. As described by W. Meyerhof *et al.*^{27,28}, these 2p vacancies could be created through a one-step or a two-step process in the first collision, and they may then be carried into the second collision and participate in subsequent interactions. Here, in our paper, we also suggest that the 2p vacancies have been created before the process described in Fig. 3, as the production of both the MO and characteristic x-rays need the 2p vacancies, and we do not look further into how they are produced.

According to the classical over-barrier (COB) model^{29,30}, when the distance between the incident ion and the solid surface reaches a critical distance $r_c \sim \sqrt{2q}/W$, electrons can be resonantly captured from the surface into projectile highly excited states $n_c \sim q/\sqrt{2W}$, where W is the work function of the metal surface (in atomic units). The charge transfer process continues until the ion is almost fully neutralized and a hollow atom/ion (HA) is formed. The potential energy of the ion is mainly released into the metal surface. The time τ_c required for the HA to arrive at the surface is approximately r_c/v . In the experiments reported here, $W(\text{Zn})$ is 0.123 a.u. and v is $1.48 \sim 2.26 \times 10^6$ m/s; therefore, for $q = 12$, $r_c \sim 39.83$ a.u., $n_c \sim 24$, and $\tau_c = 0.93 \sim 1.42 \times 10^{-15}$ s and for $q = 26$, $r_c \sim 58.63$ a.u., $n_c \sim 54$, and $\tau_c = 1.37 \sim 2.10 \times 10^{-15}$ s. That is, in our experiment, the time required for a HA to arrive at the target surface is about 10^{-15} s. As we know, the time for a HA to decay to the ground state is about $10^{-13} \sim 10^{-14}$ s, which is larger than τ_c . This indicates that the x-rays from the HA decays are ignorable. When the HA impacts onto the target surface, its electrons in high Rydberg states are mostly stripped off. Subsequently, only

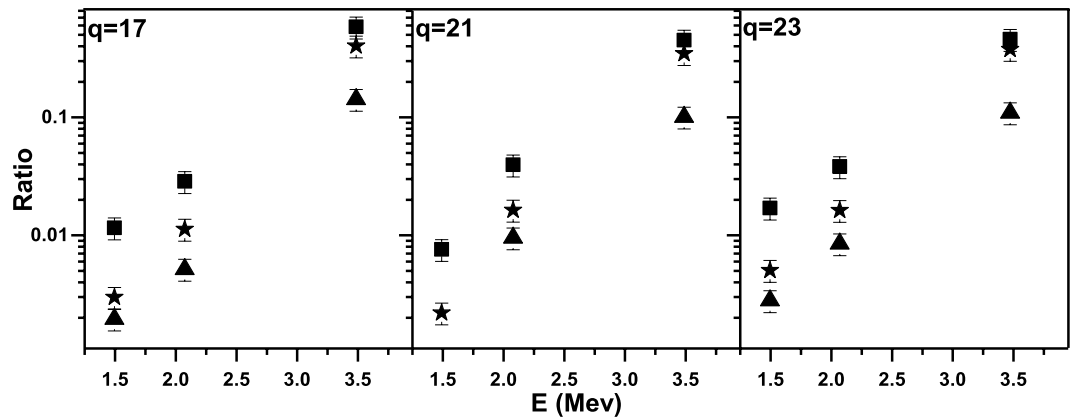


Figure 4. Variations in the ratios $\sigma_{L\alpha}/\sigma_{MO}$, $\sigma_{L\beta_1}/\sigma_{MO}$, $\sigma_{L\beta_2}/\sigma_{MO}$ as functions of the incident energy for charge states of 17, 21 and 23. Square: $\sigma_{L\alpha}/\sigma_{MO}$; Star: $\sigma_{L\beta_1}/\sigma_{MO}$; Triangle: $\sigma_{L\beta_2}/\sigma_{MO}$.

when the ions enter into the target (forming a new impact hollow atom/ion), electron-vacancy exchange process described by Fig. 3 occurs. This analysis proves from another perspective that the MO x-rays and characteristic x-rays observed in the present experiment originate mainly from below-surface.

From Fig. 3, one can see that the target 2p vacancy is shared by the $4f\pi \rightarrow 3d\pi$ and $3d\sigma \rightarrow 3d\pi$ transition processes. This leads to competition between the production of MO x-rays and characteristic x-rays, which is also reflected in Fig. 1. This competition can be observed for all projectile charge states. From the correlation diagram, it is clearly evident that the coupling process requires a Zn 2p vacancy ($3d\pi$ orbital) and a Xe 2p electron ($3d\sigma$ orbital), whereas the 2p electron ($4f\pi$ orbital) and the 2p vacancy ($3d\pi$ orbital) of the MO transition both originate from the Zn atom. Thus, the x-ray yields and the competition between them are primarily driven by the number of Zn 2p electrons. As the incident energy increase, the number of Zn 2p vacancies also increase, which leads to an increase in the production of characteristic x-rays and MO x-rays. However, the consequent decrease in the number of Zn 2p electrons causes the growth of the MO x-ray production to be slower than that of the characteristic photon production. As a result, the ratio of MO x-ray cross sections (σ_{MO}) to characteristic x-ray cross sections (σ_{cha}) is decreasing with increasing incident energy. This nicely explains the measured dependence of σ_{MO}/σ_{cha} on the projectile energy in our experiment, as shown in Fig. 1. For greater clarity, we present the following quantitative analysis.

Assuming that the Xe^{q+} ions slow down along a straight trajectory and emit x-rays isotopically and neglecting the energy loss straggling, the x-ray production cross sections σ could be extracted from the x-ray yield $Y(E)$ using the standard formula¹:

$$\sigma = (n)^{-1}[dY(E)/dE](dE/dR) + (\bar{\mu}/n)Y(E) \quad (1)$$

$$Y(E) = \frac{N_x 4\pi}{N_p \Omega \varepsilon \mu} \quad (2)$$

The quantity $Y(E)$ represents the x-ray yield; N_x is the total detected x-ray counts for each spectrum; N_p is the total number of incident particles; Ω expresses the solid angle seen by the detector from the target, which is 23.8 msr in the present experiment; ε expresses the detector efficiency of Si(Li) detector calculated using manufacture's specifications; $\bar{\mu}$ is the self-absorption coefficient of the target for its own x-rays and is acquired from NIST; μ is the photon filter transmission coefficient in 2 cm air and a 50 μm beryllium window³¹; n refers the target atom density; dE/dR represents the stopping power for incoming ions in the target, which is calculated using the SRIM-2010 program³². The uncertainty in the x-ray production cross sections including statistical and systematic errors is about 15%, which results in a total error of 21% in the ratio of σ_{MO}/σ_{cha} . Figure 4 shows the ratio of $L\alpha$ x-ray cross sections ($\sigma_{L\alpha}$) to MO x-ray cross sections ($\sigma_{L\alpha}/\sigma_{MO}$), $L\beta_1$ x-ray cross sections ($\sigma_{L\beta_1}$) to MO x-ray cross sections ($\sigma_{L\beta_1}/\sigma_{MO}$) and $L\beta_2$ x-ray cross sections ($\sigma_{L\beta_2}$) to MO x-ray cross sections ($\sigma_{L\beta_2}/\sigma_{MO}$) as a function of the incident energy, for charge states Xe^{17+} , Xe^{19+} and Xe^{23+} . Obviously, these ratios are increasing with increasing incident energy.

In addition, two 'bumps' located on the low energy side of the Xe^{q+} L x-ray lines are observed in our experiment, denoted by R_1 and R_2 in Fig. 1. Usually, an inner shell vacancy decays either through a radiative or a non-radiative transition. However, as suggested by T. Åberg *et al.*, it may also decays by the simultaneous emission of a photon and an electron. The emitted photon has energy slightly lower than that of a characteristic x-ray photon and forms low energy peaks. This process is called the Radiative Auger Effect (RAE)^{33,34}. In our results, peaks R_1 and R_2 originate from the RAE, as they are located on the low energy side of the projectile L-shell x-ray peaks.

Significantly, several new features are observed in the excited spectra for projectiles of charge state 29, as shown in Fig. 5.

Clearly, in the same energy range as the MO x-ray bands for the $q < 29$ cases, several narrow discrete lines are observed for an incident charge state of 29. It seems that the broad MO x-ray line 'splits' into several narrow lines

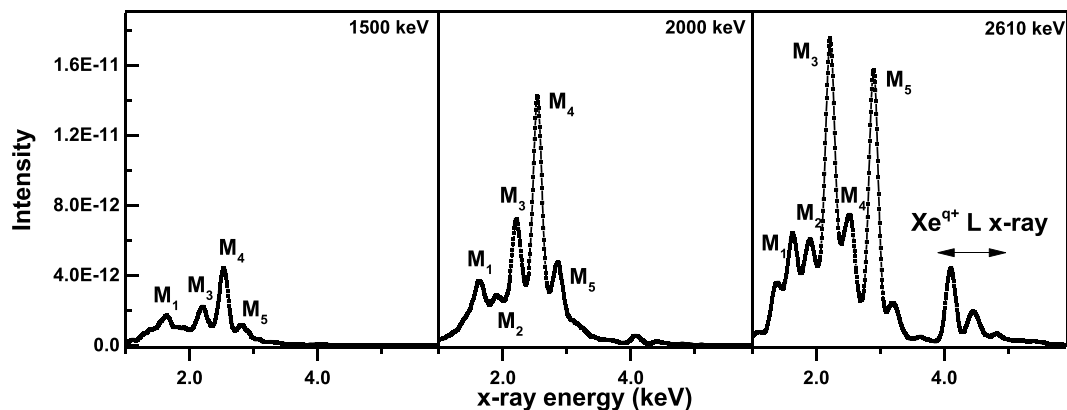


Figure 5. X-ray spectra from collisions of Xe^{29+} ions with Zn target at 1500 keV, 2000 keV and 2610 keV. M_1 : $4p_{1/2} \rightarrow 3d_{5/2}$; M_2 : $4p_{3/2} \rightarrow 3d_{3/2}$; M_3 : $4f_{5/2} \rightarrow 3d_{5/2}$; M_4 : $4f_{7/2} \rightarrow 3d_{3/2}$; M_5 : $4d_{5/2} \rightarrow 3p_{3/2}$. The spectra have been normalized by the incident particle number.

q	peak	IE (keV) EPE (keV)		1500		2000		2610 (3500)	
		Exp.	The.	σ_{MO} (b)	r'_{eff} (Å)	σ_{MO} (b)	r'_{eff} (Å)	σ_{MO} (b)	r'_{eff} (Å)
29	M_1 ($4p_{1/2} \rightarrow 3d_{5/2}$)	1.75	1.83	7.05	0.80 (−2)	13.3	1.04 (−2)	18.5	1.21 (−2)
29	M_2 ($4p_{3/2} \rightarrow 3d_{3/2}$)	2.03	2.10	1.05	0.42 (−2)	3.88	0.69 (−2)	9.35	0.96 (−2)
29	M_3 ($4f_{5/2} \rightarrow 3d_{5/2}$)	2.40	2.50	0.28	0.44 (−2)	0.98	0.70 (−2)	1.93	0.92 (−2)
29	M_4 ($4f_{7/2} \rightarrow 3d_{3/2}$)	2.70	2.79	0.60	0.56 (−2)	1.39	0.78 (−2)	0.95	0.72 (−2)
29	M_5 ($4d_{5/2} \rightarrow 3p_{3/2}$)	2.90	2.82	0.27	0.42 (−2)	0.92	0.66 (−2)	1.78	0.86 (−2)
Total σ_{MO} or r'_{eff}		—	—	9.25	0.264 (−1)	20.47	0.387 (−1)	32.54	0.467 (−1)
21	M ($4f_{7/2} \rightarrow 3d_{3/2}$)	2.70	2.79	67.5	0.27 (−1)	102	0.33 (−1)	242	0.48 (−1)

Table 1. Summary of peak M_n transition configurations²⁴, EPEs, σ_{MO} values and r'_{eff} values for $\text{Xe}^{29+} + \text{Zn}$. For comparison, the values for the spectrum M induced by Xe^{21+} and the total σ_{MO} and total r'_{eff} for Xe^{29+} are also listed. Numbers in parentheses indicate powers of ten. The errors of x-ray production cross sections are about 15%, including statistical and systematic errors.

when charge state $q = 29$, which is a phenomenon that has never before been reported in a quasi-molecular radiation study. Although the FWHMs of these narrow lines are similar to those of the characteristic lines $L\alpha$, $L\beta_1$, $L\beta_2$ and $L\gamma$, they cannot be identified as characteristic lines of either target atoms or of the projectiles, as their centroid energies do not correspond to the characteristic x-rays of target atoms or the projectile ions. Taken together, these observations permit us to attribute these new narrow bands to the MO x-rays, as well. Hence, the molecular correlation diagram will once again be employed in the following analysis. In Fig. 3, as the inter-nuclear distance shrinks, the projectile orbital also evolves into molecular orbitals. Unlike $\text{Xe}^{q < 29+}$, Xe^{29+} ion itself has three 3d vacancies. That is, for the $\text{Xe}^{29+} + \text{Zn}$ system, the predominant emission process for the MO x-rays is $4f\delta \rightarrow 3d\delta$. By contrast, the $3d\sigma \rightarrow 3d\delta$ electron promotion process leads to projectiles carrying 2p vacancies. After separate, L-shell characteristic x-rays of the projectile ions emit.

Moreover, further analysis of the experimental data in Figs 1 and 5 using equations (1) and (2) reveals that, at the same incident energy, (i) The MO x-ray production cross sections induced by $\text{Xe}^{q < 29+}$ are of the same order of magnitude; (ii) Total $\sigma_{\text{MO}}(\text{Xe}^{29+})$ of all spectra M_n is approximately one order of magnitude smaller than $\sigma_{\text{MO}}(\text{Xe}^{q < 29+})$ (see Table 1). These findings indicate that q influences not only the MO x-ray spectral profile and its origin, but also its production cross sections. Importantly, the production cross sections are closely related to the effective geometric range r_{eff} where MO transition may occur; see the shaded area in Fig. 3. Specifically, the relationship between them can be expressed in the following form³⁵:

$$\sigma_{\text{MO}} = \sigma_{\text{eff}} \varpi_{\text{MO}} \tau_{\text{eff}} / \tau_{\text{MO}} \quad (3)$$

where, $\sigma_{\text{eff}} = \pi r_{\text{eff}}^2$ and $\tau_{\text{eff}} = r_{\text{eff}} / v$.

Then, the effective geometric range induced by Xe^{q+} ions is

$$r_{\text{eff}} = (\sigma_{\text{MO}} \nu \tau_{\text{MO}} / \pi \varpi_{\text{MO}})^{1/3}. \quad (4)$$

Where, τ_{MO} describes the average lifetime of a vacancy in the quasi-molecule; τ_{MO} represents an average fluorescence yield and the production cross sections σ_{MO} (see Table 1) is extracted from the experimental data in Figs 1 and 5 using equations (1) and (2).

To look for the further reason causing MO x-ray division, the r_{eff} for $q < 29$ and $q = 29$ is calculated using equation (4), respectively.

Taking 2000 keV $\text{Xe}^{21+} + \text{Zn}$ as an example, we have $v = 1.71 \times 10^8$ cm/s, $\sigma_{\text{MO}} = 1.02 \times 10^{-22}$ cm² (from Table 1), $\tau_{\text{MO}} = 3.6 \times 10^{-2}$ and $\tau_{\text{MO}} = 3 \times 10^{-16}$ s³⁶⁻³⁸, consequently, $r_{\text{eff}} \approx 0.33 \times 10^{-9}$ cm. As we know, for the filling of a MO vacancy to occur during an interaction, the MO vacancy lifetime τ_{MO} should be comparable to the effective collision time τ_{eff} evaluated as r_{eff}/v . However, in our experiment, τ_{MO} is about one order of magnitude larger than τ_{eff} . Two factors may lead to this discrepancy. (i) A straight-line trajectory, which results in the shortest possible route and time, is assumed in our calculation of r_{eff} and τ_{eff} . In a real event, however, Coulomb (97) deflection plays an important role, especially in a slow bombardment¹, and this effect will significantly increase the interaction time and the effective range. (ii) τ_{MO} is appreciably larger in a UA than in a quasi-molecule³⁵. In Eq. (4), the average lifetime of a Po 3d vacancy is employed as there is no way to know the true τ_{MO} value in a quasi-molecule. Hence, in a real impact event, the difference between the two values should be much smaller than the calculated ones. It is clear that the present experiment conditions satisfy the space and time requirements for MO transition to occur. In addition, the EPEs, the transition configurations, the production cross sections σ_{MO} and the effective ranges r_{eff} for $q = 29$ are listed in Table 1.

Table 1 tells us that, at the same incident energy, (i) The total $\sigma_{\text{MO}}(\text{Xe}^{29+})$ for all peaks M_n is about one order of magnitude smaller than $\sigma_{\text{MO}}(\text{Xe}^{q < 29+})$; (ii) The r_{eff} corresponding to each narrow spectrum M_n is lower than that of peak M by approximately a factor of 5 and the total r_{eff} for Xe^{29+} is very similar to r_{eff} for Xe^{21+} . For comparison, the effective ranges (r'_{eff} and r_{eff}) for both peaks M_4 and M at 2000 keV are roughly indicated in Fig. 3. From the molecular correlation diagram, it is evident that a smaller r_{eff} can induce a narrower x-ray energy variation, i.e., a smaller FWHM. According to the above analysis, it is very easy to understand why, in our experiment, a broad MO x-ray band is observed for $\text{Xe}^{q < 29+}$, whereas several discrete peaks are observed for Xe^{29+} . The small MO transition effective range r_{eff} for Xe^{29+} ions is the key to explaining these experimental results. Moreover, our measured MO x-ray cross sections show a remarkably close match with this requirement (see Table 1). However, in this field, although some relevant experimental results have been reported, quantitative calculations of spectral structure and intensity are still sorely lacking, especially for multi-electron systems. To date, theoretical efforts have been primarily focused on few-electron quasi-molecules. Hence, with regard to our result, the question of why the effective range r_{eff} is small for Xe^{29+} remains open to theoretical investigation.

Conclusion

In summary, the non-characteristic and characteristic x-ray spectra measured in collisions of 1500–3500 keV Xe^{q+} ($q = 12-29$) ions with solid target Zn are investigated in detail. In the study, two different electron-vacancy exchange processes, namely, electron promotion and MO transition, corresponding to the characteristic x-rays and MO x-rays emissions respectively, are observed simultaneously in the same formation of the quasi-molecule Xe-Zn. Both the yields of characteristic x-rays and MO x-rays show a marked rise with an increase in the incident energy and the former grows faster than the latter. Meanwhile, the FWHMs of the MO band are measured and found to be proportional to the square root of the projectile velocity when the projectile charge states $q < 29$. Moreover, the present work finds very striking difference appeared in MO x-ray spectra produced by $\text{Xe}^{q < 29+}$ and Xe^{29+} ions. Several narrow discrete lines are observed for an incident charge state 29, whereas a broad x-ray band is observed for $\text{Xe}^{q < 29+}$. It seems that the broad MO X-ray line ‘splits’ into several narrow lines as the charge state $q = 29$. This finding puts forward a new question of how the MO transition depends on the projectile charge state. Although the present work tries to answer the question and gives some explanations, more experiments and, in particular, deeper theoretical investigations are definitely necessary for its full understanding.

References

- Garcia, J. D., Fortner, R. J. & Kavanagh, T. M. Inner-shell vacancy production in ion-atom collisions. *Rev. Mod. Phys.* **45**, 111–177 (1973).
- Eichler, J. & Stöhlker, T. Radiative electron capture in relativistic ion-atom collisions and the photoelectric effect in hydrogen-like high-Z systems. *Physics Reports* **439**, 1–99 (2007).
- Horvat, V., Watson, R. L. & Peng, Y. $K\alpha$ satellite and hypersatellite distributions of Ar excited in heavy-ion collisions. *Phys. Rev. A* **79**, 012708-1-7 (2009).
- Wang, J. J., Zhang, J., Gu, J. G., Luo, X. W. & Hu, B. T. Highly charged Ar^{q+} ions interacting with metals. *Phys. Rev. A* **80**, 062902-1-9 (2009).
- Watanabe, H. *et al.* X-ray emission in collisions of highly charged I, Pr, Ho, and Bi ions with a W surface. *Phys. Rev. A* **75**, 062901-1-5 (2007).
- Lichten, W. Molecular wave functions and inelastic atomic collisions. *Phys. Rev.* **164**, 131–142 (1967).
- Barat, M. & Lichten, W. Extension of the electron-promotion model to asymmetric atomic collisions. *Phys. Rev. A* **6**, 211–229 (1972).
- Song, Z. Y. *et al.* Charge state effect on K-shell ionization of aluminum by 600–3400 keV xenon^{q+} ($12 < q < 29$) ion collisions. *Eur. Phys. J. D* **64**, 197–201 (2011).
- Mace, J., Gordon, M. J. & Giapis, K. P. Evidence of simultaneous double-electron promotion in F^+ collisions with surfaces. *Phys. Rev. Lett.* **97**, 257603-1-4 (2006).
- Commisso, M. *et al.* Plasmon excitation and electron promotion in the interaction of slow Na^+ ions with Al surfaces. *Nucl. Instrum. Methods Phys. Res. Sect. B* **230**, 438–442 (2005).
- Saris, F. W., van der Weg, W. F., Tawara, H. & Laubert, R. Radiative transitions between quasimolecular levels during energetic atom-atom collisions. *Phys. Rev. Lett.* **28**, 717–720 (1972).
- German, K. A. H., Weare, C. B. & Yarmoff, J. A. Inner-shell electron promotion in low energy Li^+ -Al(100) collisions. *Phys. Rev. Lett.* **72**, 3899–3902 (1994).

13. Sun, J. *et al.* K and L x-ray emission from hollow atoms produced in the interaction of slow H-like (F^{22+}) and bare (F^{53+}) ions with different target materials. *Phys. Rev. A* **77**, 032901-1-5 (2008).
14. Sun, H. L. *et al.* Charge-state dependence of K-shell x-ray production in aluminum by 2–12-MeV carbon ions. *Phys. Rev. A* **53**, 4190–4197 (1996).
15. Meyerhof, W. E. X-rays from coalescing atoms. *Science* **193**, 839–848 (1976).
16. Prior, M. H. *et al.* Quasimolecular x-ray spectrum from 117-keV Ne^{9++} Ne collisions. *Phys. Rev. A* **47**, 2964–2967 (1993).
17. Stöckli, M. P. & Anholt, R. Observation of peaked $2p\sigma$ molecular-orbital spectra. *Phys. Lett.* **106A**, 130–132 (1984).
18. Schulze, K., Anton, J., Sepp, W.-D. & Fricke, B. Energy dependence of the molecular-orbital x-ray interference structure in U^{92+} -Pb collisions. *Phys. Rev. A* **58**, 1578–1580 (1998).
19. Devdariani, A. Radiative transitions in few-electron quasi-molecules. *Adv. Space Res.* **54**, 1173–1179 (2014).
20. Devdariani, A., Dalimier, E. & Sauvan, P. Optical Transitions and Charge-Exchange in Highly Charged Quasi-Molecules. *International Journal of Spectroscopy* **2010**, 1–12 (2010).
21. Devdariani, A. *et al.* Dipole transition-matrix elements of the one-electron heterodiatomic quasimolecules. *Phys. Rev. A* **71**, 022512-1-12 (2005).
22. Sun, L. T. *et al.* Commissioning test of LAPECR2 source on the 320 kV HV platform. *High Energy Phys. Nucl. Phys.* **31**(Supp I), 55–59 (2007).
23. Macdonald, J. R., Brown, M. D. & Chiao, T. Observation of a K X-Ray band emitted by the transient C-C system formed at keV energies. *Phys. Rev. Lett.* **30**, 471–474 (1973).
24. Zschornack, G. *Handbook of X-Ray Data*. (Springer-Verlag Berlin Heidelberg, 2007).
25. Betz, H.-D. *et al.* Spectral shape and cross section of molecular-orbital X-ray continua from heavy-ion collisions. *Phys. Rev. Lett.* **34**, 1256–1259 (1975).
26. Anholt, R. Theory of the angular distribution of molecular orbital K x rays seen in heavy-ion-atom collisions. *Z. Physik A* **288**, 257–276 (1978).
27. Meyerhof, W. E. *et al.* Observation of $Z = 70$ quasiatomatic K X rays from 30- and 60-MeV ${}_{35}Br + {}_{35}Br$ Collisions. *Phys. Rev. Lett.* **30**, 1279–1282 (1973).
28. Meyerhof, W. E. *et al.* Molecular-orbital K X-ray formation in heavy-ion collisions. *Phys. Rev. Lett.* **32**, 1279–1282 (1974).
29. Burgdörfer, J., Lerner, P. & Meyer, F. Above-surface neutralization of highly charged ions: The classical over-the-barrier model. *Phys. Rev. A* **44**, 5674–5685 (1991).
30. Lemell, C. *et al.* Image acceleration of highly charged ions by metal surfaces. *Phys. Rev. A* **53**, 880–885 (1996).
31. Gullikson, E. X-ray interactions with matter. Available at: http://henke.lbl.gov/optical_constants/ (1995).
32. Ziegler, J. F. SRIM, Program Version SRIM-2010 (2010).
33. Åberg, T. & Utriainen, J. Evidence for a “Radiative Auger Effect” in X-ray photon emission. *Phys. Rev. Lett.* **22**, 1346–1348 (1969).
34. Åberg, T. Theory of the radiative auger effect. *Phys. Rev. A* **4**, 1735–1740 (1971).
35. Mokler, P. H., Stein, H. J. & Armbruster, P. X rays from superheavy quasiatoms transiently formed during heavy-ion-atom collisions. *Phys. Rev. Lett.* **29**, 827–830 (1972).
36. Sögüt, Ö. *et al.* Fit values of M subshell fluorescence yields and Coster-Kronig transitions for elements with $20 \leq Z \leq 90$. *X-Ray Spectrometry* **31**, 62–70 (2002).
37. Perkins, S. T. *et al.* Tables and graphs of atomic subshell relaxation data derived from the LLNL evaluated atomic data library $Z = 1-100$. *Lawrence Livermore National Laboratory Report, UCRL 50400 30*, Livermore (1991).
38. Bambynek, W. *et al.* X-ray fluorescence yields, auger, and Coster-Kronig transition probabilities. *Rev. Mod. Phys.* **44**, 716–813 (1972).

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Author Contributions

Y.G. analyzed the data and wrote the manuscript. Z.Y. designed this experiment. Z.Y., Z.S., Q.X., B.Z., J.C. and B.Y. performed the experiment. B.H. and X.W. discussed the data. B.H. and J.Y. modified the manuscript.

Additional Information

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