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Beam generation and structural optimization of size-selected Au₉₂₃ clusters

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A size-selected beam of Au_{923±20} clusters is generated in a gas-phase condensation cluster source equipped with a lateral time-of-flight mass selector. The beam current reaches up to 9.13 nA for small clusters and 80 pA for Au_{923±20} clusters, which are then analyzed using a scanning transmission electron microscope. Four types of metastable structures are observed for the Au_{923±20} clusters, including icosahedron (Dh), cuboctahedron and icosahedron (Ih). The proportion of bulk-favorable cuboctahedron (*i.e.* face center cubic (Fcc)) structure takes up only 10–20%, while the penta-rotating symmetrical structures (Dh/Ih) are the dominant ones which take up over three quarters. Changing the beam condition may optimize the clusters from Dh-dominant to the Ih-dominant phase, which paves the way towards nanoparticle control beyond the diameters.

Clusters are a series of materials composed of several to hundreds of atoms. The limited atom numbers and large specific surface area indicate that clusters cannot easily form periodical structures and can show numerous other structures.^{1–3} Finding new structures is the fundamental task of the cluster science, based on which numerous novel properties have been discovered.^{4,5} One of the attractive element is gold. Due to the strong relativistic effect, the decrease in the energy of the 6s electron and the increase in that of the 5d electrons along with s-d hybridization occur in gold,^{6,7} playing a key role in the formation of novel structures, such as cages,⁸ planar structures,⁷ Dh, Ih⁵ and tetrahedron.⁹ The evolution of the cluster structure of gold with respect to size/atomic number,^{10,11} temperature,^{12,13} substrates^{14,15} and even gas environments¹⁶ has been studied in detail, but only few studies commit to the directional structure control of these clusters.¹⁷ The precise control of the cluster structural transformation with certain atomic numbers and high yield are of great importance for the

precise study of properties and further practical applications. Not only catalysis¹⁸ but optical nanoclusters¹⁹ also require the precise controlled clusters with both atomic number and nanostructure precision.

Au₉₂₃ is a magic number cluster²⁰ with sufficient yield in our experiment. The structures of the as-prepared Au₉₂₃ through a magnetron sputtering cluster source in STEM include Dh, Fcc and Ih.²¹ The direct structure control of such a big cluster from the original Au atoms will be of great importance beyond the size control, which has been achieved for a long time. Here, we study the structural optimization of the size-selected Au₉₂₃ clusters through our magnetron sputtering cluster source. The original highest proportion of the Dh structure with considerable yield can be conversed to another penta-rotating symmetrical Ih structure with proper conditions; however, the bulk-favorite Fcc structure remains only a little part between 10% and 20%.

A magnetron sputtering gas-phase condensation cluster source equipped with a lateral time-of-flight mass selector^{22,23} is used for cluster preparation and its schematic is shown in Fig. 1(a). Clusters with a mass resolution of about 4% can be produced. The process of cluster formation is shown in Fig. 1(b). Energetic gold atoms are sputtered out by bombarding the gold target with ionized argon atoms and then cooled down through the multi-body collision with He atoms.^{24,25} Then, the cooled Au atoms aggregate with each other and gradually grow up. By controlling the collision time we can enrich clusters within a certain atomic number range. The as-prepared clusters finally soft-land on ultrathin carbon films. The morphology and size of the as-synthesized Au_{923±20} clusters are studied using a spherical-aberration-corrected STEM, carried out on an FEI Themis Z at the National Graphene Center in Wuxi, China. In order to diminish the statistical error for size and structural statistics, over 50 images are collected for each group of clusters under different synthesis conditions. Four types of structures are observed, including Dh, Fcc, Ih and an unknown structure, which cannot be classified due to the blurred image. To classify the structures, the simulated STEM image atlas of each possible

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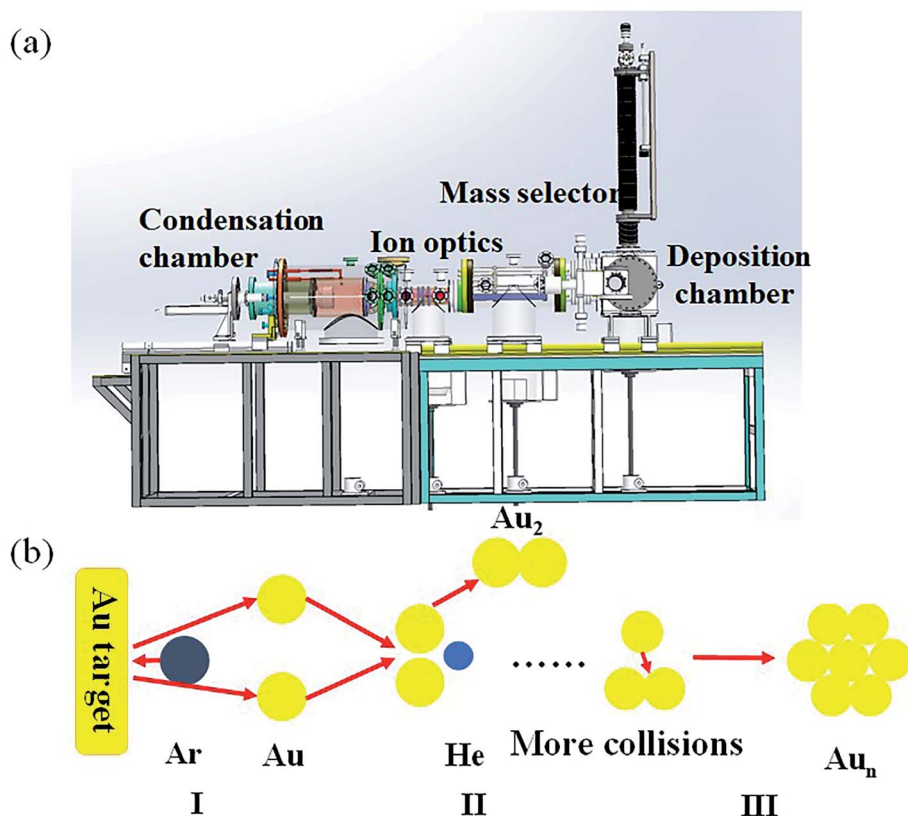


Fig. 1 Schematic of the magnetron sputtering gas-phase condensation cluster source with time-of-flight mass-selector and the formation process of clusters. (a) The schematic of the cluster source. The source is composed of four parts, namely condensation chamber, ion optics, TOF mass selector and the deposition chamber. (b) The formation process of clusters in the condensation chamber. (I) The hot gold atoms (golden) are sputtered out by argon atoms (gray). (II and III) The sputtered atoms are cooled down by multi-body collisions with helium gas (blue) and aggregate with each other.

structure is obtained. By comparing the simulation and experimental images, the structure of each cluster in the STEM image can be assigned. Except for a normally prepared Au_{923±20} clusters (group 1) with the highest beam current of about 80 pA, we prepared other three groups of Au_{923±20} clusters by changing the preparation conditions from group 1, which are shown in Table 1 below with five sets of condensation parameters. We primarily changed the sputtering power, He gas flow and the condensation length while slightly changing Ar gas to sustain a steady and detectable deposition current. The total pressure in the condensation chamber is changed when we change the helium and argon gas flows.

The yield of the Au_{923±20} clusters *versus* the condensation length is shown in Fig. 2(a). The collision times between gold

atoms, clusters and helium atoms increases with the condensation length, leading to a higher cluster yield. The mass spectra of small gold clusters and the deposition current of the Au_{923±20} clusters *versus* time are shown in Fig. 2(b). For small clusters, there exist discrete mass spectra, while the yields of Au₂ and Au₃ are 9.13 nA and 6.64 nA, respectively. The inset shows the optimized deposition beam current of the Au_{923±20} clusters *versus* time. When we switch off and switch on the mass selection pulse to stop and restart cluster deposition, the beam current shows an obvious decrease to zero, followed by rebound. In this way, we confirm that the deposition current of the Au_{923±20} clusters is real rather than the background noise, and the maximum steady deposition current for Au_{923±20} is about 80 pA. In Fig. 2(c), the STEM images of the Au_{923±20}

Table 1 The variation of synthesis conditions for the different structure proportions of Au_{923±20} clusters

Group number	Sputtering power (W)	Helium flow (sccm)	Argon flow (sccm)	Condensation pressure (Pa)	Condensation length (cm)
1	20	154	176	54	27
2	60	15	184	17	12
3	80	11	179	16	12
4	120	11	183	17	12

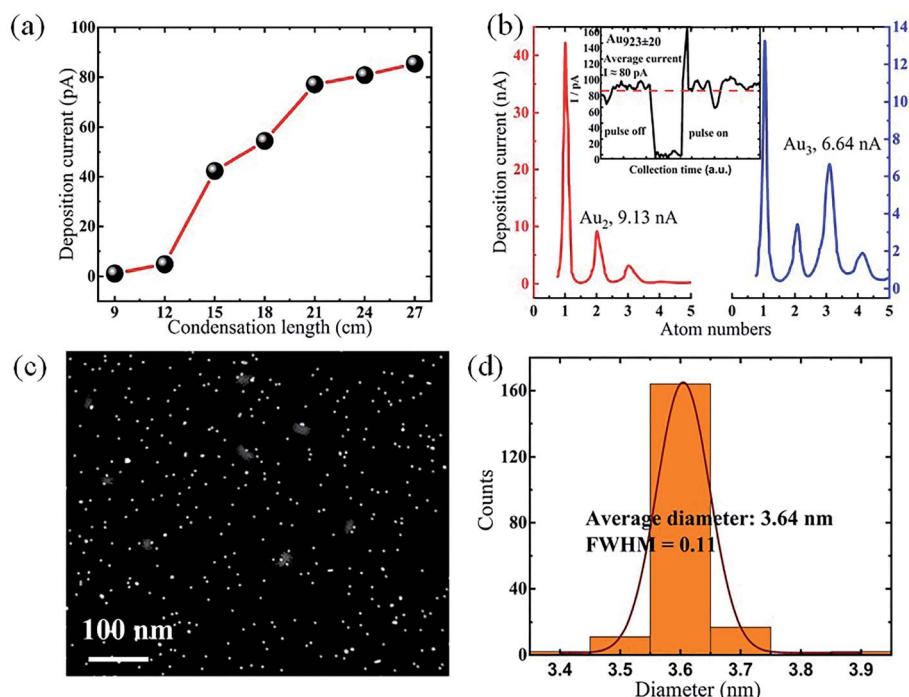


Fig. 2 Influences of the gold cluster yield and the size statistics. (a) The influence of the condensation length on the Au_{923±20} cluster yield. The deposition current increases monotonously as the condensation length increases. (b) Yields of gold clusters. The controllable maximum beam current of Au₂ and Au₃ clusters are 9.13 and 6.64 nA, respectively. The inset shows the maximum steady deposition current of Au_{923±20} clusters versus time. When we switch off the pulse to stop collection, an obvious decrease in the deposition current from 80 pA to about 0 appears. (c) The uniform distribution of the as-prepared Au_{923±20} clusters in the STEM image of in a low magnification. (d) Size statistics of Au_{923±20} clusters. The average size is 3.67 nm and the FWHM is 0.11 nm.

clusters with low magnification clearly indicate their uniform spherical shape. The average size is 3.64 ± 0.11 nm, as shown in Fig. 2(d), after measuring over 100 clusters. The narrow FWHM with the dense distribution shows that we successfully synthesized the Au_{923±20} clusters with a high yield. Three sets of

matches between simulated and experimental STEM images for typical structures of the Au_{923±20} clusters are shown in Fig. 3(a)–(c). Each simulated structure has some typical characteristics to distinguish from others. Taking Fig. 3(c) for instance, the circular rings with a bright dot in it is a typical feature of the Ih structure in STEM. In this manner, we classify the cluster in Fig. 3(c) as an Ih structure, while the three turns of stripes and the straight stripe distinguish them from each other (Fig. 3(a) and (b)). In Fig. 3(d), the fuzzy STEM image is attributed to either the naturally amorphous structure or the specific angle facing the electron beam cannot match each of the simulated images; therefore, we classify this as an unknown structure. By comparing simulated and experimental images, clusters with different structures observed in STEM are classified and counted. The proportions of the above four structures considerably change in the four groups, as shown in Fig. 4(a)–(d). The dark gray, red, blue and green parts in the pie charts represent Dh, Fcc, Ih and an unknown structure, respectively. In Fig. 4(a), namely, group 1, a statistical result of the normal synthesis condition is observed where the Dh structure occupies a majority of all clusters with a proportion of 65.45%. The Ih structure only accounts for a small part of 10.92%. Moreover, the bulk-favorable cuboctahedron (*i.e.* Fcc) structure only takes a small part of 18.18%. However, in group 2, with a condensation length of 12 cm, the lowest helium gas flow, a sputtering power of 60 W, and an obvious increase of Ih structure from 10.92% to 31.65% were observed. Simultaneously, the Dh

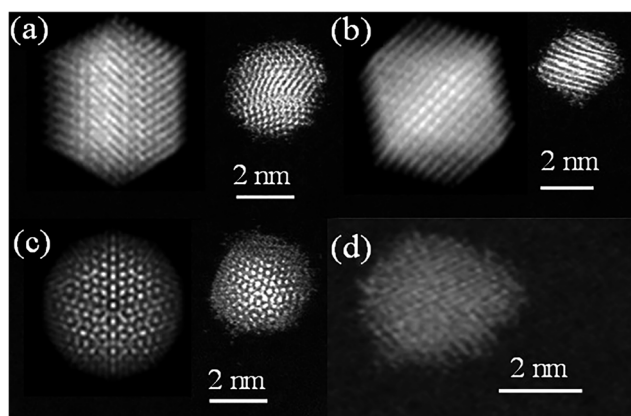


Fig. 3 Comparison between the simulated STEM images and the experimental ones for different structures. (a) On the top left corner is the simulated STEM image of a Dh structure, while the top right one is an experimental result. (b) The simulated STEM image and an experimental result of an Fcc structure. (c) The simulated STEM image and an experimental result of an Ih structure. (d) The experimental STEM images that cannot match with each simulated STEM ones.

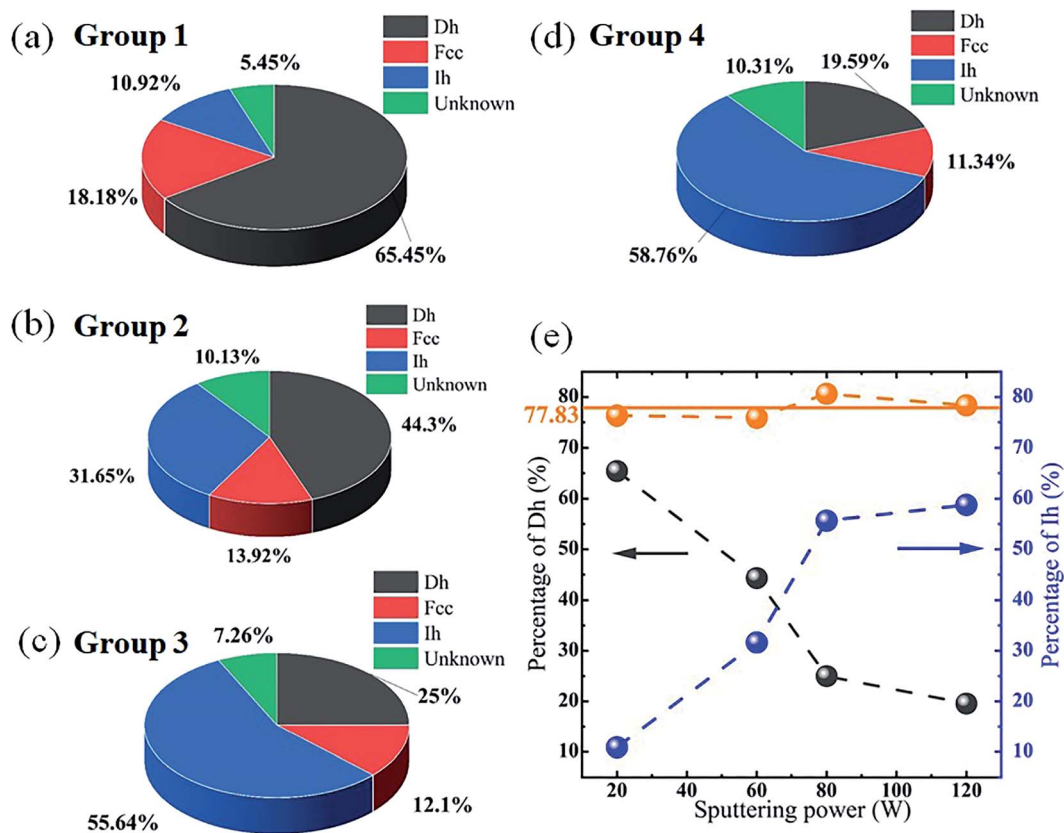


Fig. 4 The structural statistic of the $\text{Au}_{923\pm 20}$ clusters in different synthesis conditions. (a) The normal synthesis condition for the $\text{Au}_{923\pm 20}$ clusters with 20 W sputtering power, 27 cm condensation length, and high helium gas flow. (b)–(d) The changed synthesis condition with the lowest helium gas flow and the shortest condensation length of 12 cm. The sputtering power increased from 20 to 60, 80 and 120 W. (e) The dark gray and blue balls refer to the proportion of Dh and Ih structures, while the orange one refers to the sum of proportions of the two structure. The orange-dashed line represents the average proportion of the sum.

structure shows a rapid decrease to 44.3%, which may indicate a structure transformation from Dh to Ih. In this condition, the Fcc structure had a slight decrease to 13.92%. When we continue to increase the sputtering power to 80 W while keeping other parameters almost unchanged, the structure proportion continues to change. As shown in Fig. 4(c), the Ih structure reaches 55.64%, while the Dh structure declines to only 25%. Furthermore, the Fcc structure shows a very slight decrease to 12.1%. When further increasing the sputtering power to 120 W, the transformation to the Ih structure slows down and only reaches 58.76%. In Fig. 4(e) the visualizing trend about the two structures Dh and Ih *versus* different sputtering power is shown. The dark gray and blue balls refer to the proportion of the Dh and Ih structures. The orange one refers to the sum of the proportions of the two structures, while the orange dashed line represents the average proportion of it. It is obvious that the proportions of Dh and Ih structures show considerable changes in a variety of synthesis conditions when the Ih structure rises from 10.92% to 58.76% and the Dh structure falls from 65.45% to 19.59%. Furthermore, we note that the sum of these two structures occupy a quite stable percentage of about 77.83% for all groups. Therefore, it might be possible that the transformation from Dh to Ih is the major process in our structural control.

Discussion

The yield of clusters with certain atom numbers reaches maximum when we have a suitable condensation condition to control the collision time during the cluster growth. However, the cluster yield is partially limited because only part of atoms aggregates into the (923 ± 20) range. The beam consumption in the chamber occupies a large part in the total loss; therefore, the aerodynamic optimization of chamber is required to effectively guide the beam out of the chamber. As for the structure control, the $\text{Au}_{923\pm 20}$ clusters prepared in the gas phase are mostly in penta-rotating symmetrical structures (Dh or Ih structures) and the proportion of the sum of these two structures takes up over three quarters. Nevertheless, the proportion of bulk-favorable cuboctahedron structure takes up only 10–20%. It needs further study that at which atom numbers the bulk-favorable structure begin to be the thermodynamically most stable structure. This atom numbers could be a limit for bulk-like Au nanoparticles.¹⁰ The effective structure control indicates that the proportions of different cluster structures are controllable by changing the synthesis condition in the condensation chamber. As shown by previously reported studies,^{5,26} the vibrational frequencies in Ih structures are smaller than Dh ones, while Dh structures have softer vibrations than Fcc ones. Considering the entropic

contributions in cluster structure transitions, we can simply make a point that for a given atomic numbers of clusters, Ih structures prefer higher temperature than Dh and Fcc ones. In our experiments, by decreasing the condensation length as well as the flow of helium gas, the collision times of clusters with helium gas decreased, thus the heat carried away from clusters reduced, creating high temperature during cluster growth.^{27,28} The increased sputtering power leads to the excessive sputtered energetic gold atoms, which further results in insufficient cooling within the gold atom gas. In this manner, we increase the average temperature of clusters, and the preferred structural transition from Dh to Ih happens. The bulk-favorite Fcc structures are influenced in the same way by the gradually significant entropic contributions. Note that the increased temperature suppresses the previously small proportion of it. These changes lead to an evolution in Dh fractions.^{29,30} However, there remain some problems. First, the cluster yield shows a great decrease to several pA when we change the synthesis conditions because the increased temperature, the insufficient helium gas and the shortest condensation length influence the average collision times in the aggregation of atoms,^{27,31} creating a harsh environment for the cluster growth, particularly for such a large cluster. Second, the proportion of penta rotation structures is almost invariable, which may indicate that there exists an upper limit in the Ih structure if the most probable structure transforming happens from the Dh to the Ih structure. Finally, the separation³² and preservation of the metastable clusters^{33,34} are needed for further applications. In our study, the structure change during the STEM can easily happen during the irradiation of electron beam.

Conclusion

In this study, atomically precise Au_{923±20} clusters are generated in a gas-phase condensation cluster source equipped with a lateral time-of-flight mass selector. The beam current reaches up to 9.13 nA for small clusters and 80 pA for the Au_{923±20} clusters. For Au_{923±20}, the average size is confirmed by STEM to be 3.64 ± 0.11 nm with a narrow distribution FWHM = 0.11 nm. By adjusting the synthesizing conditions in the condensation chamber, we successfully enrich the Ih clusters of size 923 ± 20 gold atoms to be the absolute majority from 10.92% to 58.76%, while the original majority Dh structure falls from 65.45% to 19.59%.

Conflicts of interest

The authors declare no conflict of interest.

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